

# Natural Supersymmetry

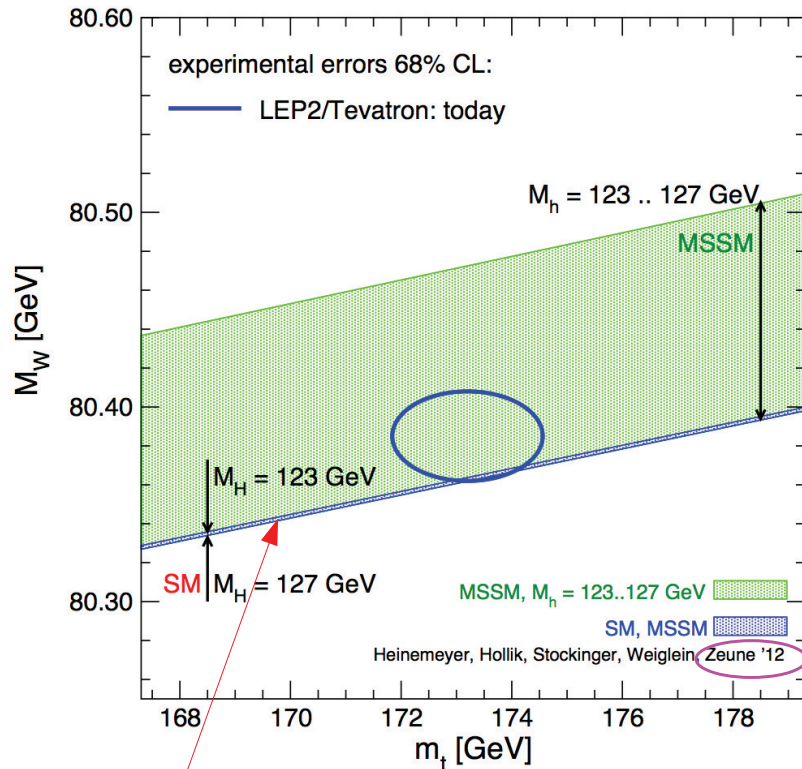
## Dark Matter, Neutrinos & LHC

Werner Porod

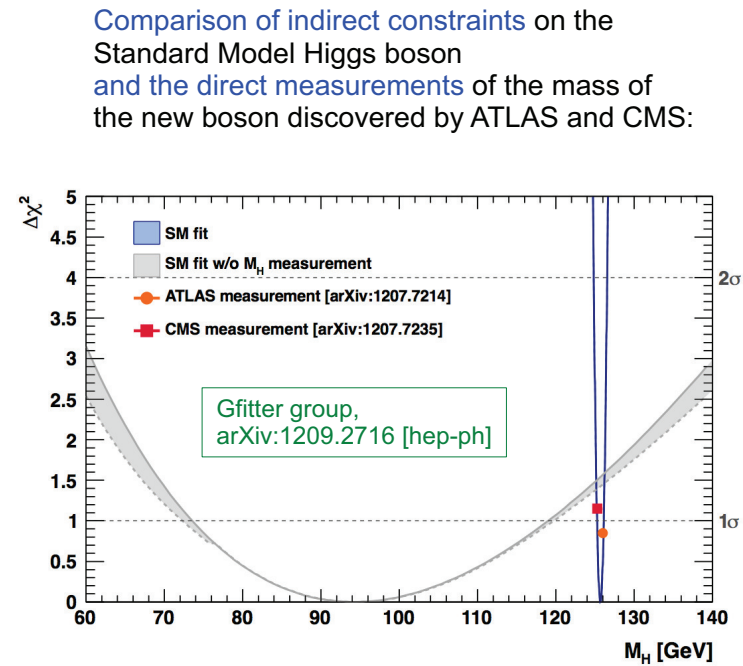
Universität Würzburg

- Why extending the SM at all, why supersymmetry
- MSSM
  - Higgs mass: consequences for GMSB & CMSSM
  - general MSSM, 'natural SUSY'
  - dark matter
- SUSY and extended gauge groups
  - implications for SUSY cascade decays
  - $Z'$  physics
  - 'Natural SUSY' and  $\tilde{\nu}_R$ -LSP

# W boson mass



In the context of the standard model, the mass of the new boson discovered by ATLAS+CMS is inside this blue band.

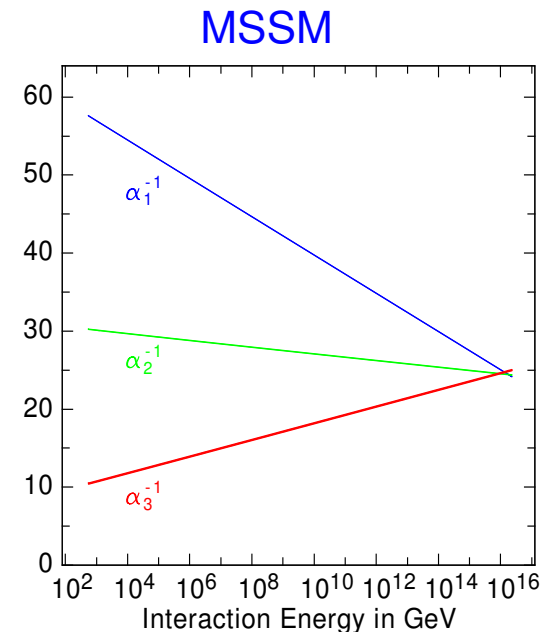
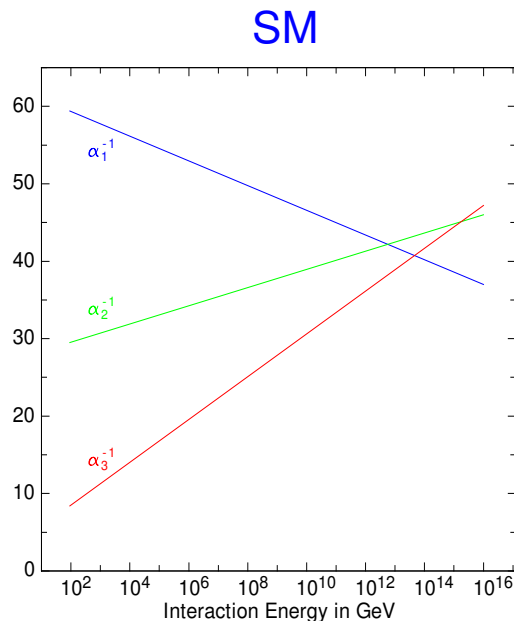


Consistent at the 1.3  $\sigma$  level.

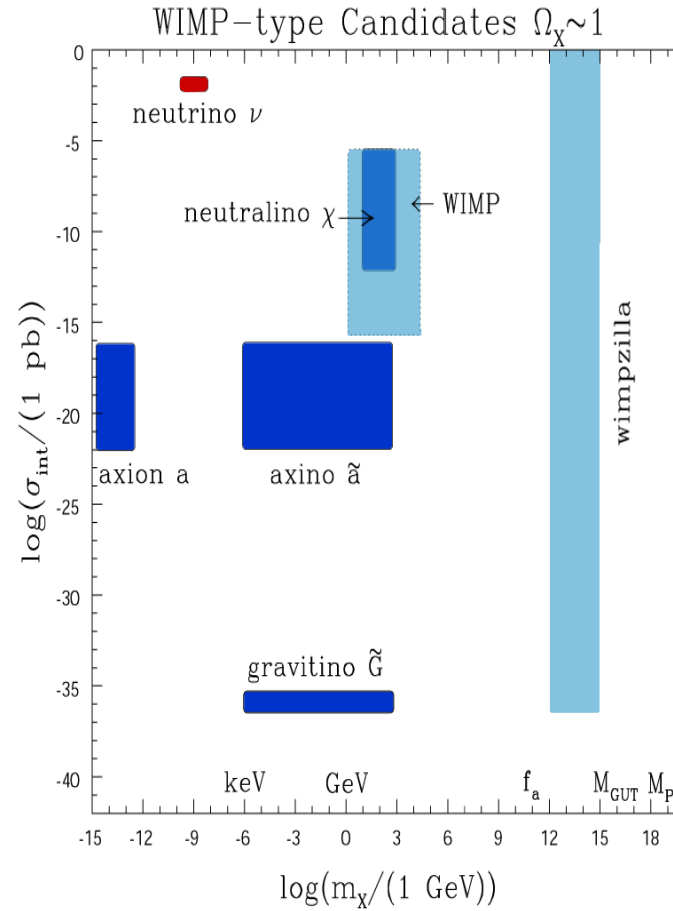
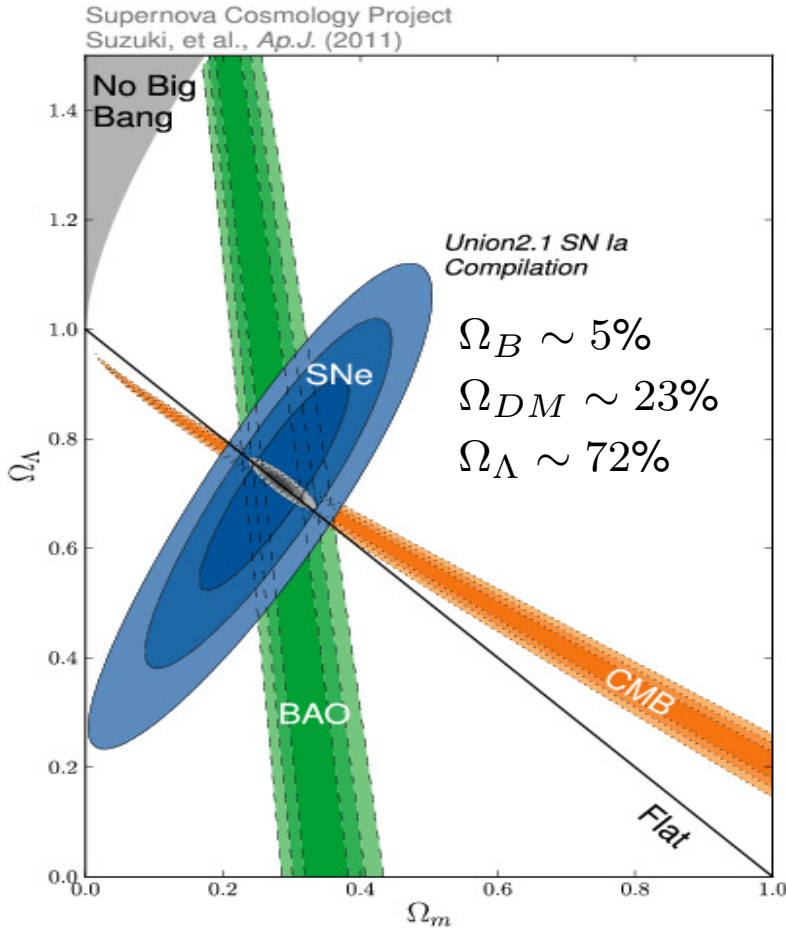
- How to combine gravity with the SM?  
⇒ local Supersymmetry (SUSY) implies gravity
- SM particles can be put in multiplets of larger gauge groups
  - in  $SU(5)$ :  $1 = \nu_R^c$ ,  $5 = (d_{\alpha,R}^c, \nu_{l,L}, l_L)$ ,  $10 = (u_{\alpha,L}, u_{\alpha,R}^c, d_{\alpha,L}, l_R)$
  - in  $SO(10)$ :  $16 = (u_{\alpha,L}, u_{\alpha,R}^c, d_{\alpha,L}, d_{\alpha,R}^c, l_L, l_R, \nu_{l,L}, \nu_R^c)$

However there are two problems in the SM but not in SUSY:

- proton decay (also in SUSY  $SU(5)$  a problem)
- gauge coupling unification



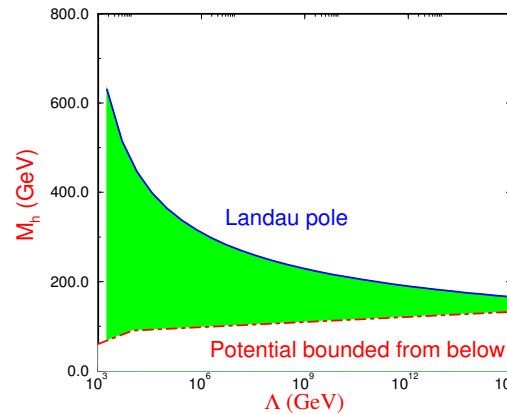
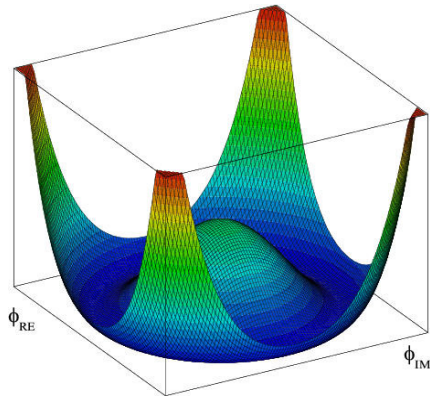
**What is the nature of dark matter ?**



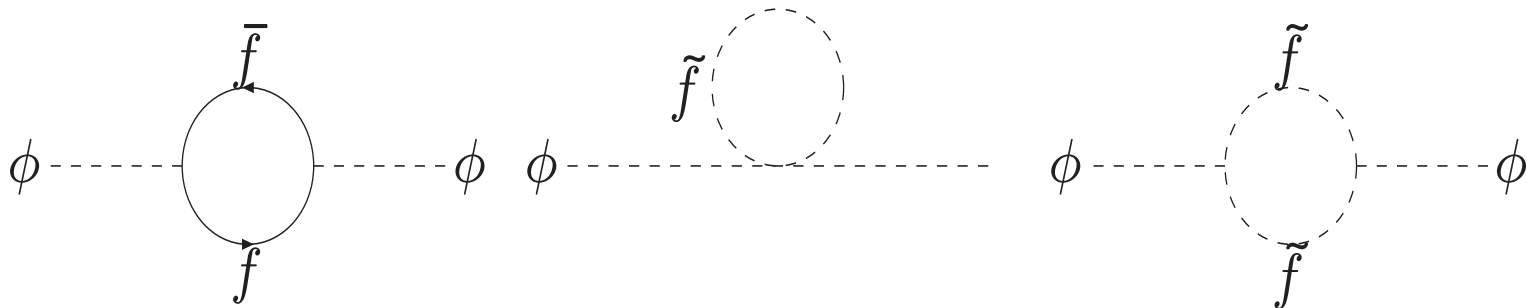
L. Roszkowski, astro-ph/0404052

**What is the origin of the observed baryon asymmetry?**

- SM &  $m_h = 125.1$  GeV: potentially meta-stable (G. Degrassi *et al.*, arXiv:1205.6497)



- ”Why does electroweak symmetry break?” or ”Why is  $\mu^2 < 0$  in the SM?”
- Hierarchy problem

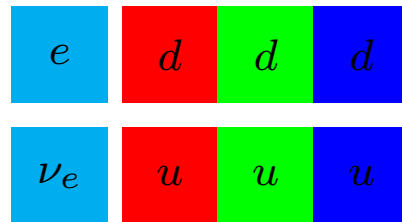


$\delta m_h^2 \propto \Lambda^2$ : Sensitivity to highest mass scale of unknown physics

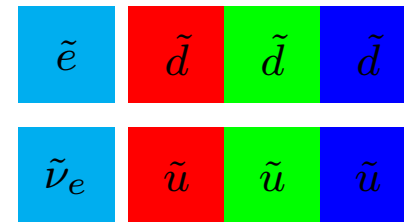
Standard Model

MSSM

matter:



$\Leftrightarrow$



gauge sector:



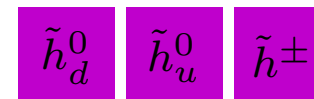
$\Leftrightarrow$



Higgs sector:



$\Leftrightarrow$



$R$ -Parity:  $(-1)^{(3(B-L)+2s)}$

$(\tilde{\gamma}, \tilde{z}^0, \tilde{h}_d^0, \tilde{h}_u^0) \rightarrow \tilde{\chi}_i^0, (\tilde{w}^\pm, \tilde{h}^\pm) \rightarrow \tilde{\chi}_j^\pm$



$$\begin{aligned}
 W_{MSSM} &= -\mu \hat{H}_d \hat{H}_u + \hat{H}_d \hat{L} Y_e \hat{E}^c + \hat{H}_d \hat{Q} Y_d \hat{D}^c - \hat{H}_u \hat{Q} Y_u \hat{U}^c \\
 W_{\mathcal{L}} &= \epsilon_i \hat{L}_i \hat{H}_u^b + \lambda_{ijk} \hat{L}_i \hat{L}_j \hat{E}_k^c + \lambda'_{ijk} \hat{L}_i \hat{Q}_j \hat{D}_k \\
 W_{\mathcal{B}} &= \lambda''_{ijk} \hat{U}_i^c \hat{D}_j^c \hat{D}_k^c
 \end{aligned}$$

$W_{\mathcal{L}} + W_{\mathcal{B}} \Rightarrow$  proton decay  $\Rightarrow R$ -parity

$$R \equiv (-1)^{3(B-L)+2s} \quad \text{or} \quad (-1)^{3B+L+2s}$$

soft SUSY breaking terms

$$\begin{aligned}
 -\mathcal{L}_{soft} &= \frac{1}{2} \left( M_3 \tilde{g} \tilde{g} + M_2 \tilde{W} \tilde{W} + M_1 \tilde{B} \tilde{B} \right) \\
 &+ m_{\tilde{Q}}^2 \tilde{Q}^* \tilde{Q} + m_{\tilde{u}}^2 \tilde{u}_R^* \tilde{u}_R + m_{\tilde{d}}^2 \tilde{d}_R^* \tilde{d}_R \\
 &+ m_{\tilde{L}}^2 \tilde{L}^* \tilde{L} + m_{\tilde{e}}^2 \tilde{e}_R^* \tilde{e}_R + m_{H_d}^2 |H_d|^2 + m_{H_u}^2 |H_u|^2 \\
 &- B\mu \epsilon_{ij} (H_d^i H_u^j + \text{h.c.}) \\
 &+ \epsilon_{ij} \left( H_d^i \tilde{Q}^j T_d \tilde{d}_R^* + H_u^j \tilde{Q}^i T_u \tilde{u}_R^* + H_d^i \tilde{L}^j T_e \tilde{e}_R^* + \text{h.c.} \right)
 \end{aligned}$$



general MSSM: more than 100 parameters

reduction assuming correlations between various parameters

● mSUGRA/CMSSM:  $M_{GUT}$

$$M_{1/2} = M_1 = M_2 = M_3$$

$$m_0^2 = m_{H_d}^2 = m_{H_u}^2, m_0^2 \cdot \mathbb{1}_3 = m_{\tilde{Q}}^2 = m_{\tilde{U}}^2 = m_{\tilde{D}}^2 = m_{\tilde{L}}^2 = m_{\tilde{E}}^2$$

$$T_f = A_0 Y_f \quad (f = u, d, e)$$

NUHM1/NHUM2:  $m_{H_d}^2, m_{H_u}^2 \neq m_0^2$

● GMSB,  $M \gtrsim 100 \text{ TeV}$

$$M_i = g(x, n) \alpha_i \Lambda$$

$$m_{\tilde{F}}^2 = f(x, n) \sum_i C_2(R) \alpha_i^2 \Lambda^2 \mathbb{1}_3$$

$$T_f \simeq 0$$

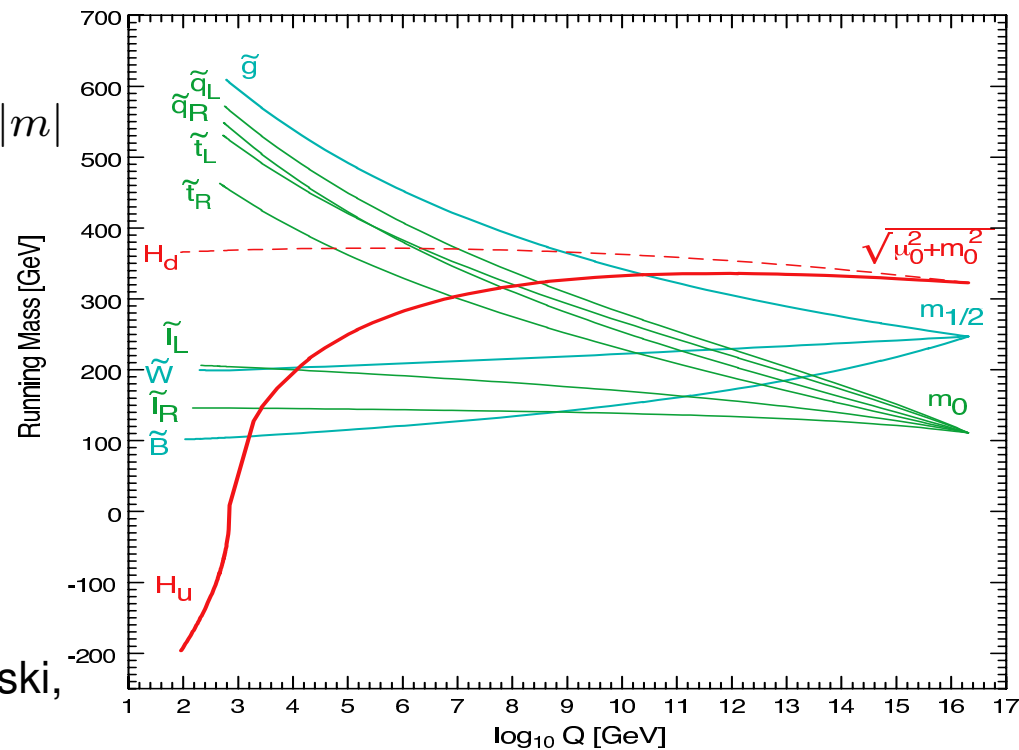
$n$  # of messenger fields,  $x = \Lambda/M$ ,  $\Lambda = O(100 \text{ TeV}) < M$

radiative electroweak symmetry breaking

$$\frac{d}{dt} \begin{pmatrix} m_{H_u}^2 \\ m_{\tilde{t}_R}^2 \\ m_{\tilde{Q}_L^3}^2 \end{pmatrix} = -\frac{8\alpha_s}{3\pi} M_3^2 \begin{pmatrix} 0 \\ 1 \\ 1 \end{pmatrix} + \frac{Y_t^2}{8\pi^2} \left( m_{\tilde{Q}_L^3}^2 + m_{\tilde{t}_R}^2 + m_{H_u}^2 + A_t^2 \right) \begin{pmatrix} 3 \\ 2 \\ 1 \end{pmatrix}$$

with  $t = \ln Q/m_Z$

$\text{sign}(m^2)|m|$



G. Kane, C. Kolda, L. Roszkowski,  
J. Wells, PRD 1994

- after EWSB:  
neutral CP-even:  $h, H$                       neutral CP-odd:  $A$                       charged:  $H^+, H^-$

- Higgs masses:  
at tree level  
 $m_A, \tan \beta = v_u/v_d$   
 $m_h \leq m_Z$   
at higher order:  
Ellis et al; Okada et al; Haber,Hempfling;  
Hoang et al; Carena et al; Heinemeyer et al;  
Zhang et al; Brignole et al; Harlander et al;  
Kant,Harlander,Mihaila,Steinhauser;...

$$m_h^2 \simeq m_Z^2 \cos^2(2\beta) + \frac{3m_t^4}{4\pi^2 v^2} \left[ \ln \left( \frac{M_S^2}{m_t^2} \right) + \frac{X_t^2}{M_S^2} \left( 1 - \frac{X_t^2}{12M_S^2} \right) \right]$$

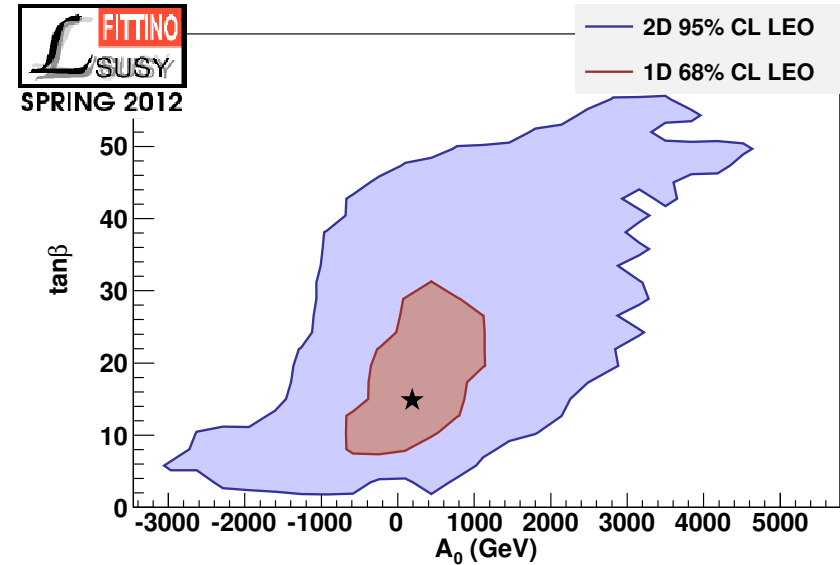
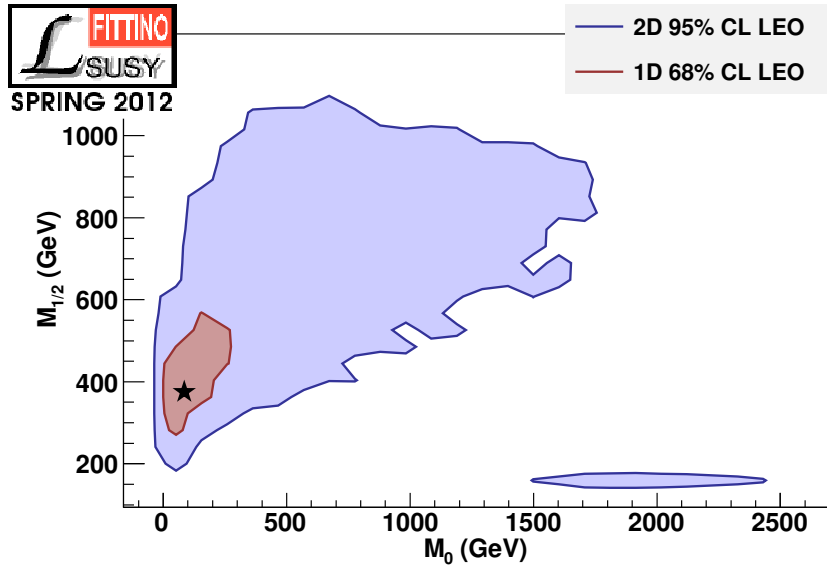
$$M_S^2 = m_{\tilde{t}_1} m_{\tilde{t}_2}, \quad X_t = A_t - \mu \cot \beta$$

$$m_H, m_A, m_{H^\pm} : O(v) \dots O(TeV)$$

$$m_{H^\pm}^2 = m_A^2 + m_W^2$$

$$v^2 = v_u^2 + v_d^2 = 4m_W^2/g^2$$

decoupling limit:  $m_A \gg v, \tan \beta \gg 1$



$\mathcal{B}(b \rightarrow s\gamma)$	$(3.55 \pm 0.34) \times 10^{-4}$
$\mathcal{B}(B_s \rightarrow \mu\mu)$	$< 4.5 \times 10^{-9}$
$\mathcal{B}(B \rightarrow \tau\nu)$	$(1.67 \pm 0.39) \times 10^{-4}$
$\Delta m_{B_s}$	$17.78 \pm 5.2 \text{ ps}^{-1}$
$a_\mu^{\text{exp}} - a_\mu^{\text{SM}}$	$(28.7 \pm 8.2) \times 10^{-10}$
$m_W$	$(80.385 \pm 0.015) \text{ GeV}$
$\sin^2 \theta_{\text{eff}}$	$0.23113 \pm 0.00021$
$\Omega_{\text{CDM}} h^2$	$0.1123 \pm 0.0118$

$\Rightarrow M_0 = 84_{-28}^{+145} \text{ GeV}, M_{1/2} = 375_{-88}^{+175} \text{ GeV},$

$\tan \beta = 15_{-7}^{+17} A_0 = 186_{-844}^{+831} \text{ GeV},$

$\chi^2/ndf = 10.3/8$

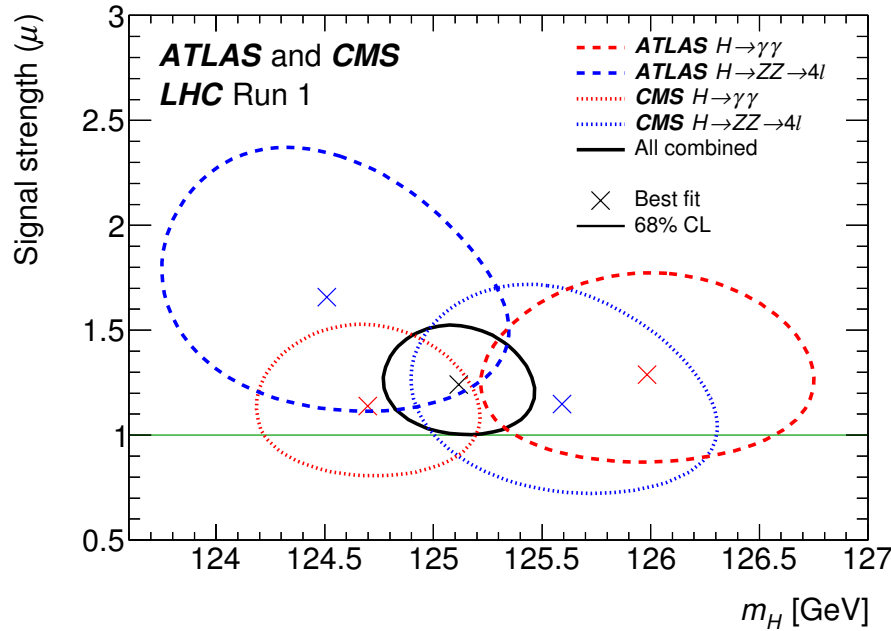
$\Rightarrow m_h = 116 \text{ GeV}$

P. Bechtle et al., arXiv:1204.4199

similar results by other groups

e.g. MasterCode, O. Buchmueller et al.

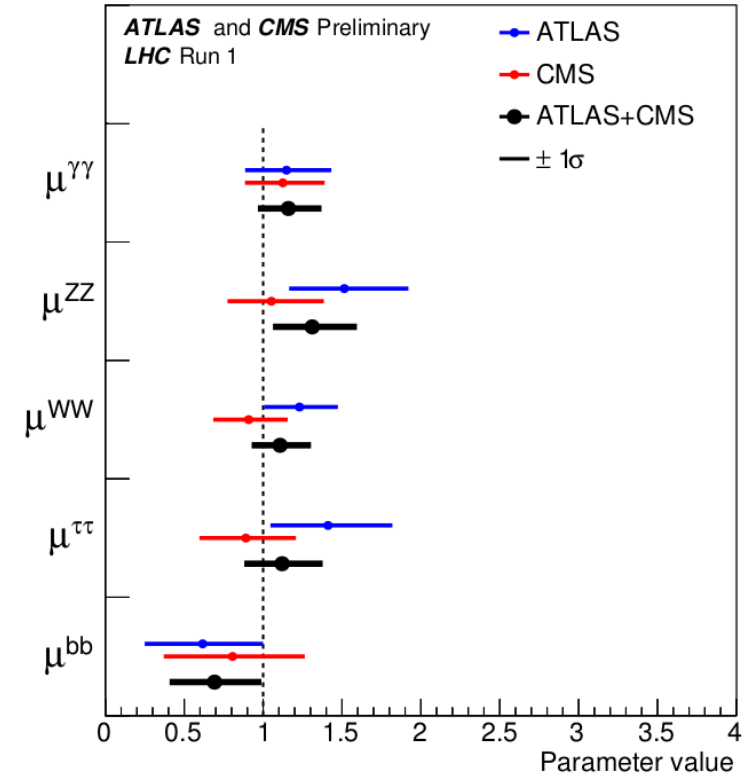
BayesFITS, L. Roszkowski et al.



$$m_H = 125.09 \pm 0.21 \text{ (stat)} \pm 0.11 \text{ (sys)} \text{ GeV}$$

PRL 114 (2015) 191803

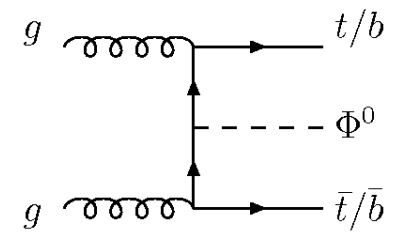
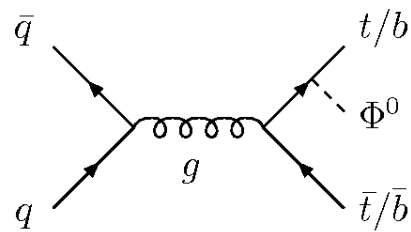
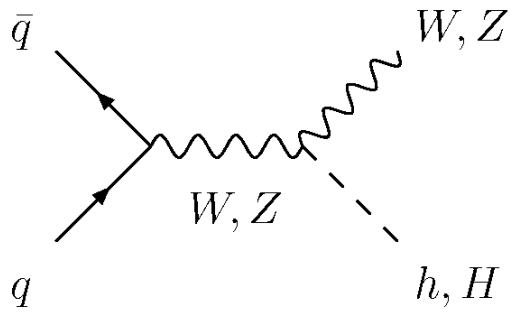
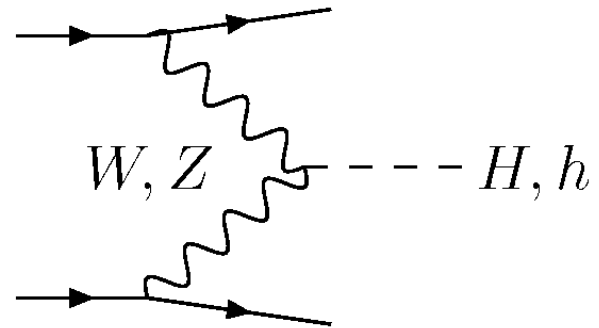
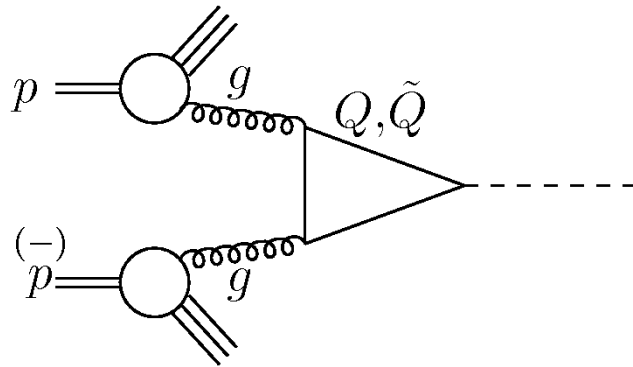
$$(125 \text{ GeV})^2 \simeq m_Z^2 + (86 \text{ GeV})^2 \Rightarrow \text{large corrections within MSSM}$$



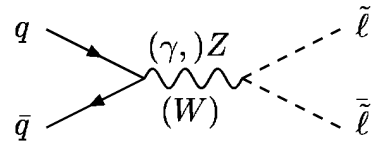
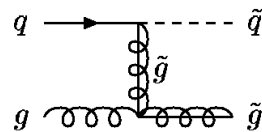
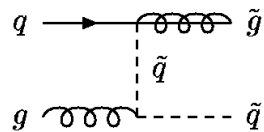
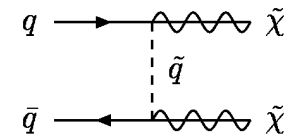
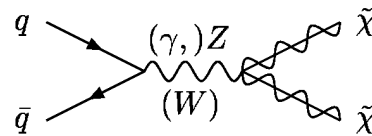
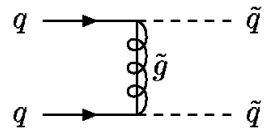
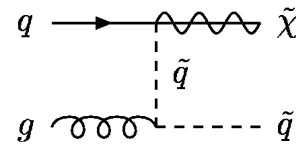
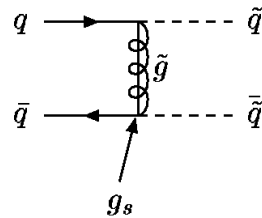
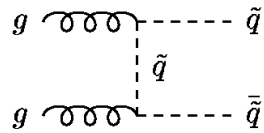
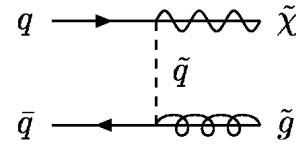
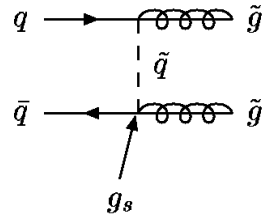
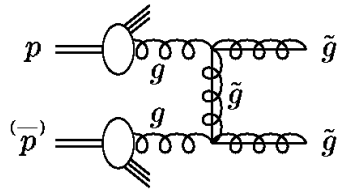
ATLAS-CONF-2015-044

CMS-PAS-HIG-15-002

production at hadron colliders

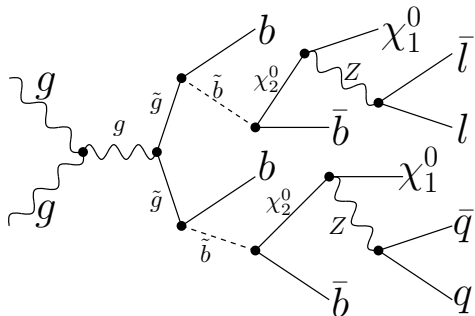
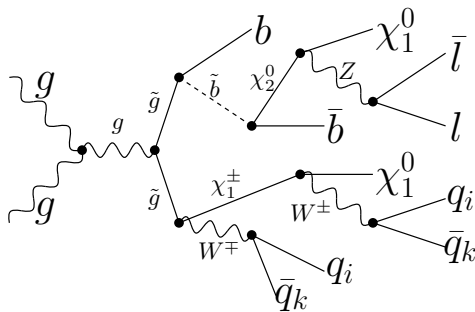
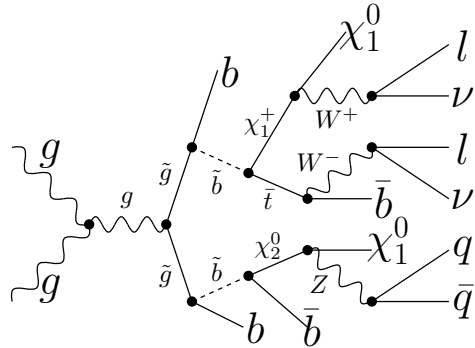


$$\Phi^0 = h, H, A$$

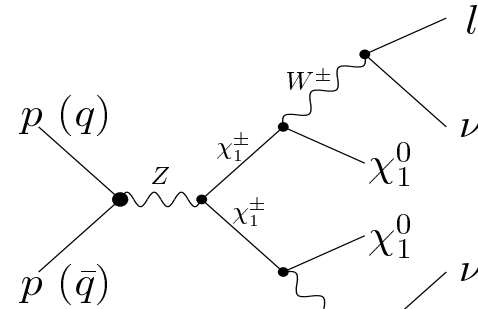
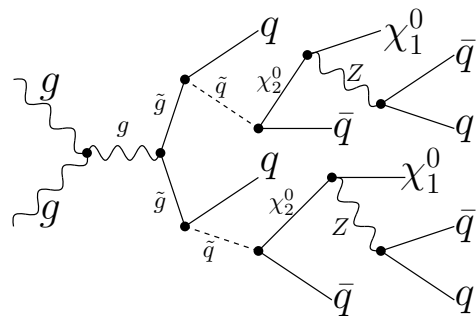
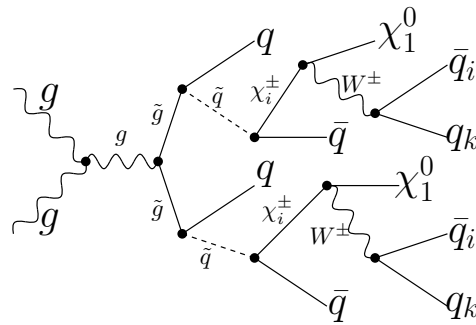
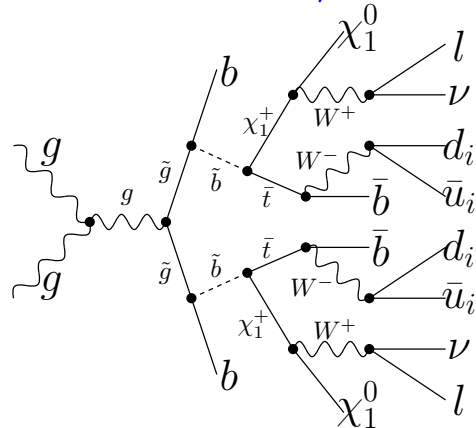




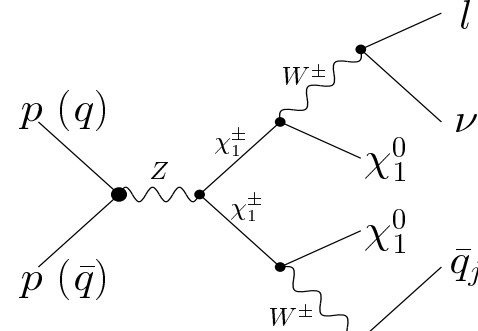
$$6j + 2l + \cancel{E/T}$$



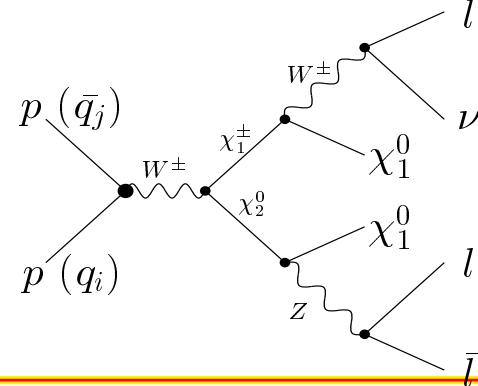
$$8j + 2l + \cancel{E/T}$$



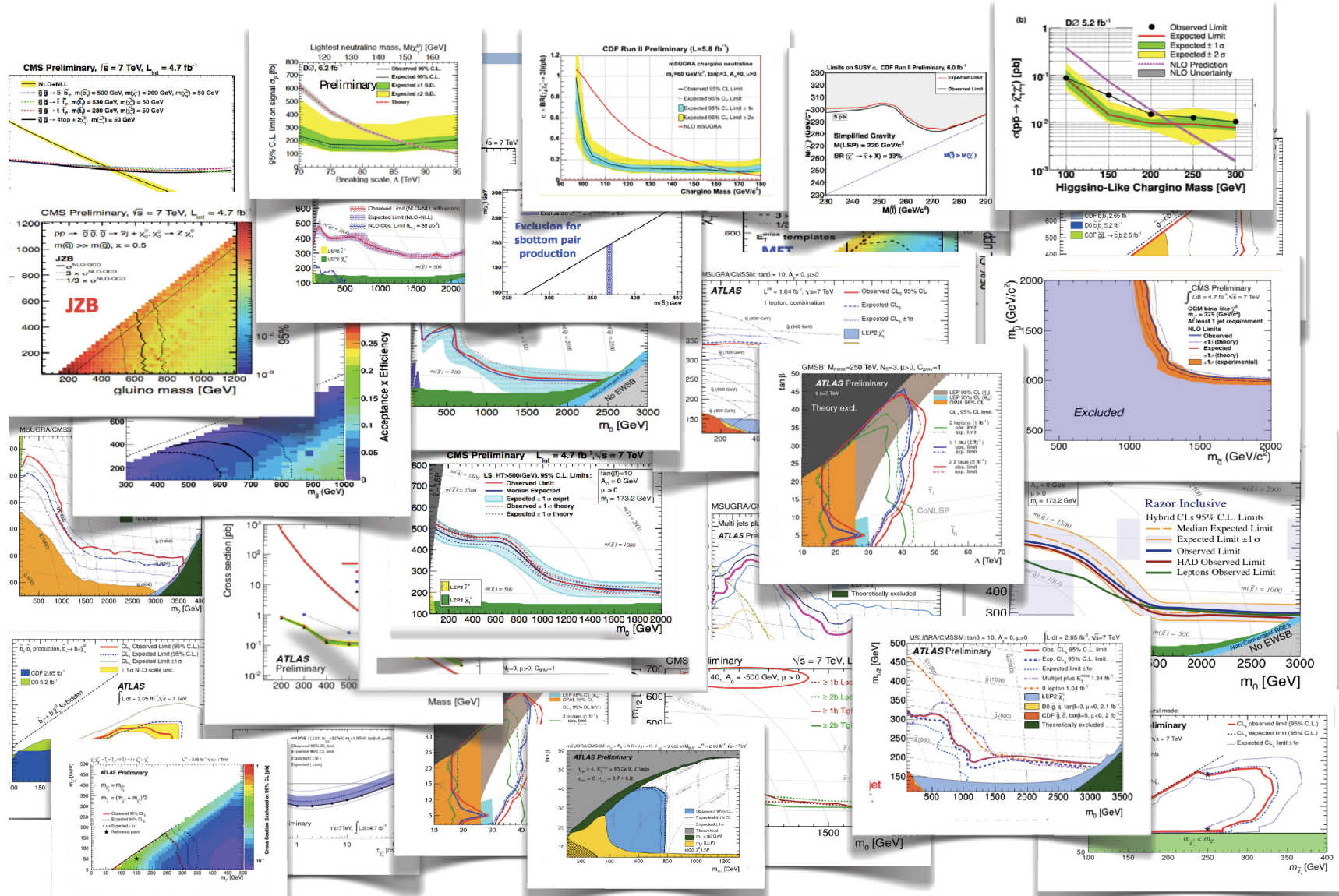
$$2l + \cancel{E/T}$$

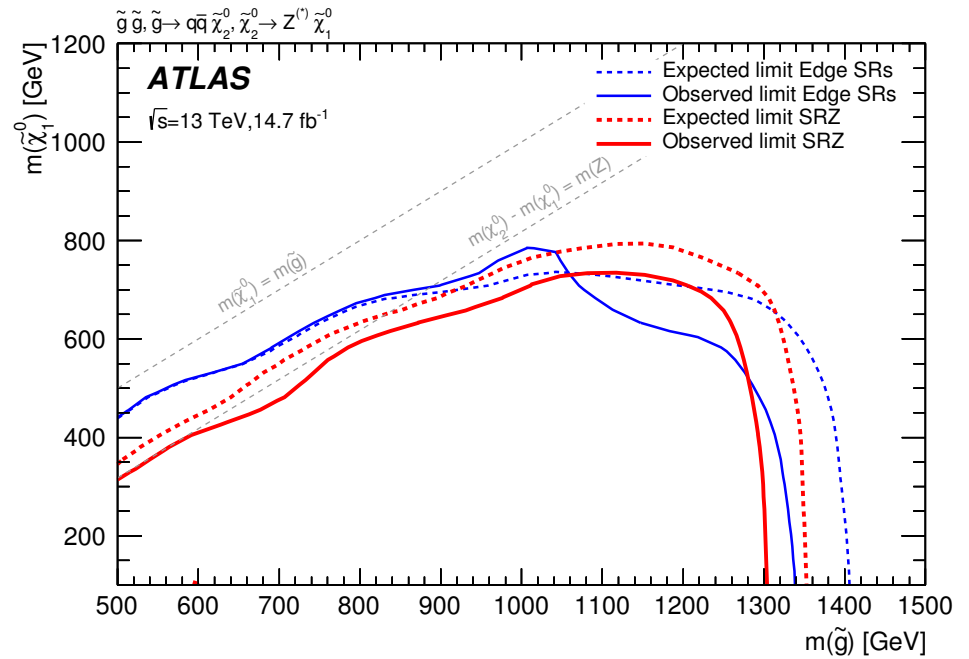
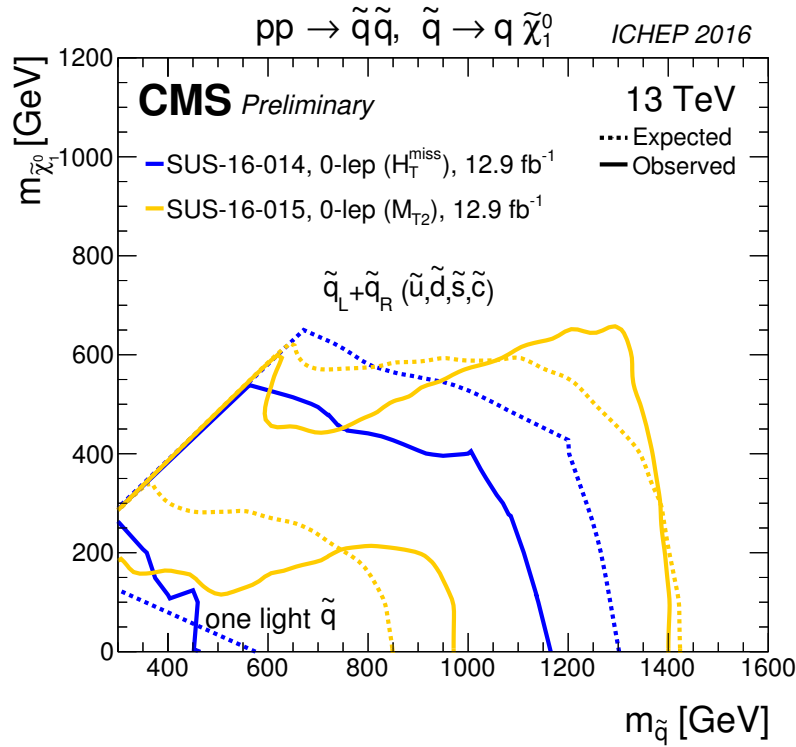


$$l + 2j + \cancel{E/T}$$



$$3l + \cancel{E/T}$$





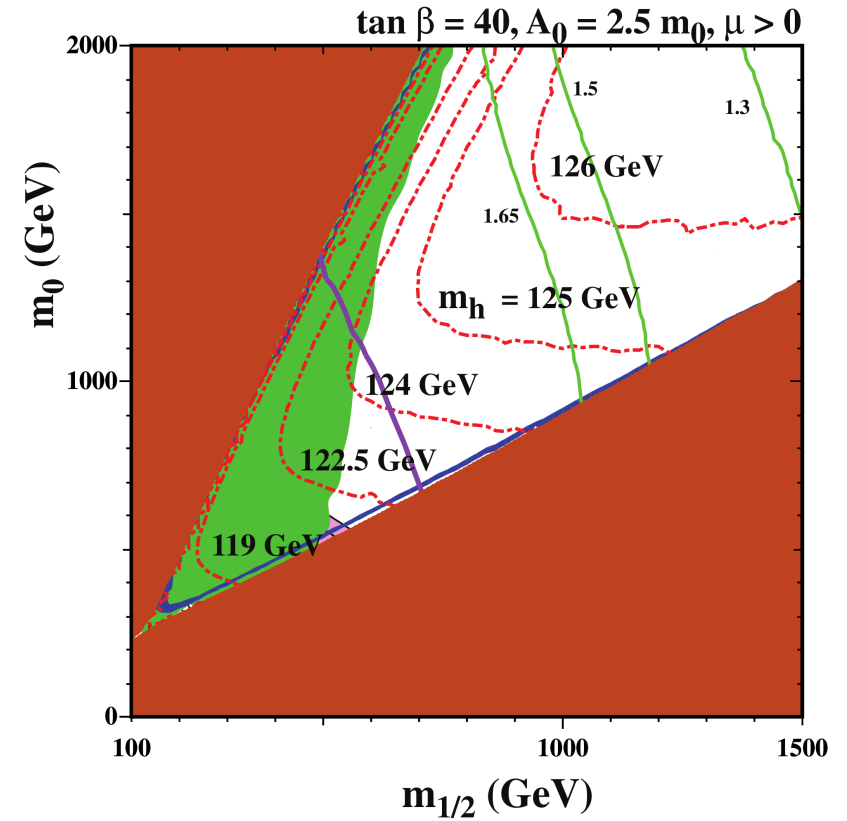
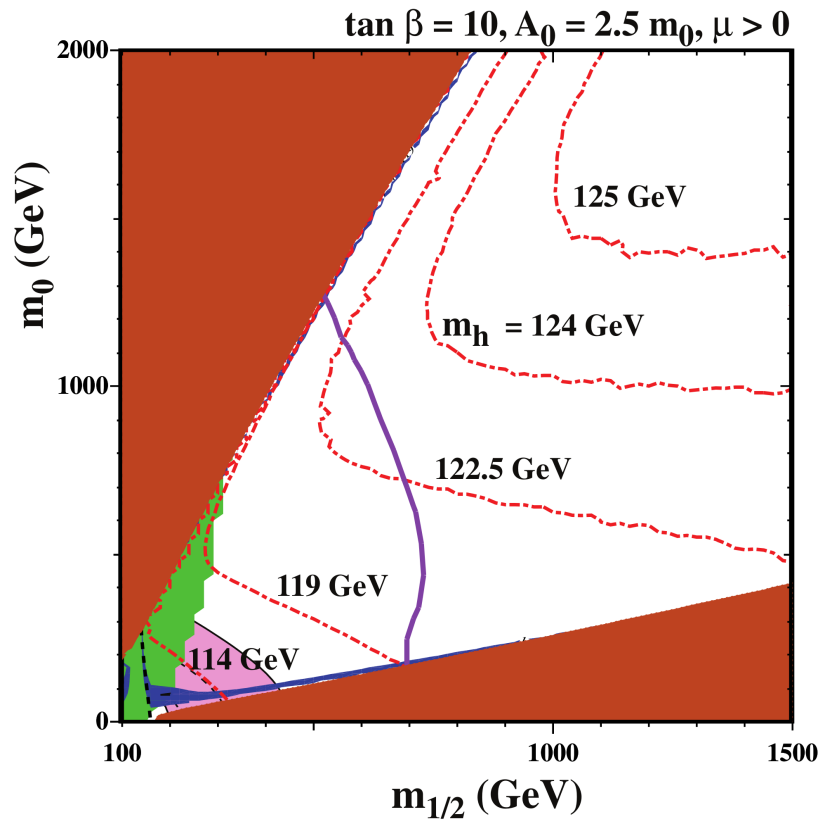
$m_h = 125.1 \text{ GeV} \Rightarrow$  large loop contributions  
 $\Rightarrow$  heavy stops and/or large left-right mixing for stops

- GMSB:  $m_{\tilde{t}_1} \gtrsim 6 \text{ TeV}$ ,  
 M. A. Ajaib, I. Gogoladze, F. Nasir, Q. Shafi, arXiv:1204.2856

more complicated models based on P. Meade, N. Seiberg and D. Shih,  
 arXiv:0801.3278  $\Rightarrow$  allow additional terms, choice not well motivated  $\Rightarrow$  generic MSSM

- CMSSM, NUHM models:  $|A_0| \simeq 2m_0$ ,  
 H. Baer, V. Barger and A. Mustafayev, arXiv:1112.3017; M. Kadastik *et al.*,  
 arXiv:1112.3647; O. Buchmueller *et al.*, arXiv:1112.3564; J. Cao, Z. Heng, D. Li,  
 J. M. Yang, arXiv:1112.4391; L. Aparicio, D. G. Cerdeno, L. E. Ibanez,  
 arXiv:1202.0822; J. Ellis, K. A. Olive, arXiv:1202.3262; ...

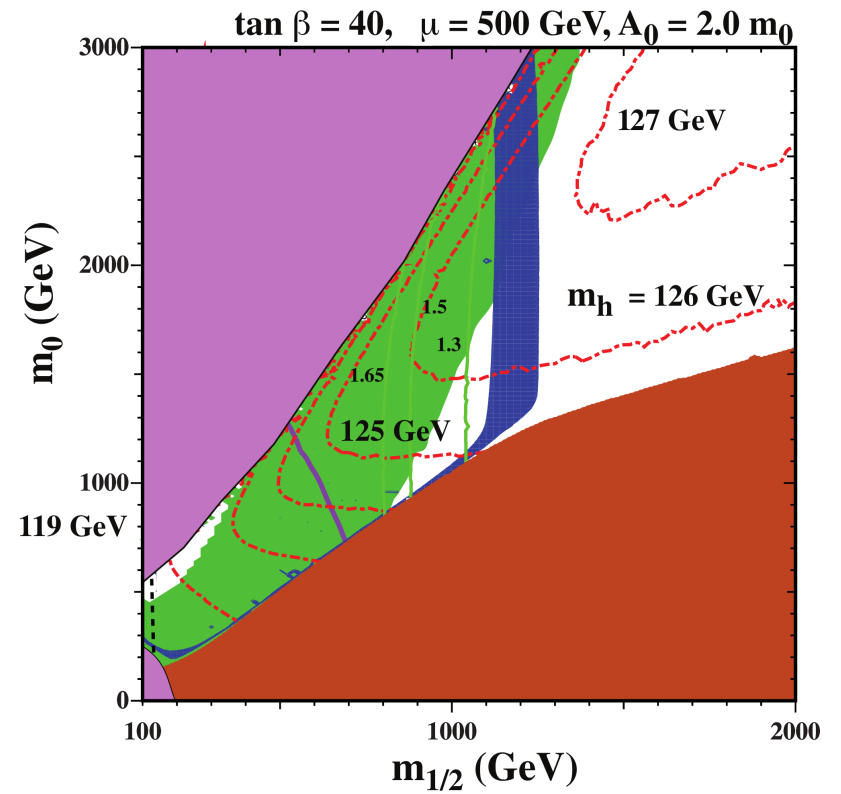
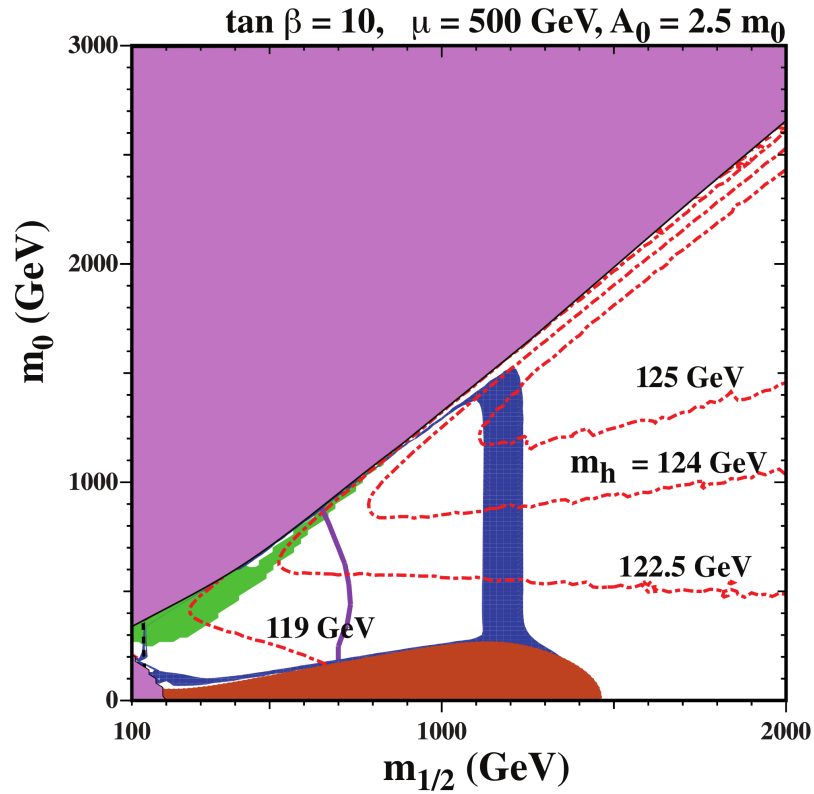
- general high scale models:  $A_0 \simeq -(1-3) \max(M_{1/2}, m_{Q_3, GUT}, m_{U_3, GUT})$   
 among other cases, details in F. Brümmer, S. Kraml and S. Kulkarni, arXiv:1204.5977



J. Ellis, F. Luo, K. Olive, P. Sandick, arXiv:1212.4476;  $m_t = 173.2$  GeV

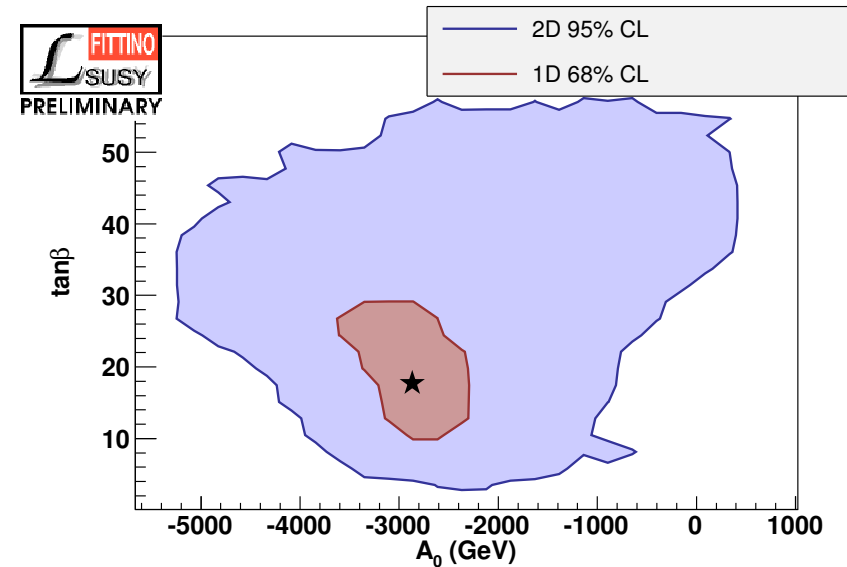
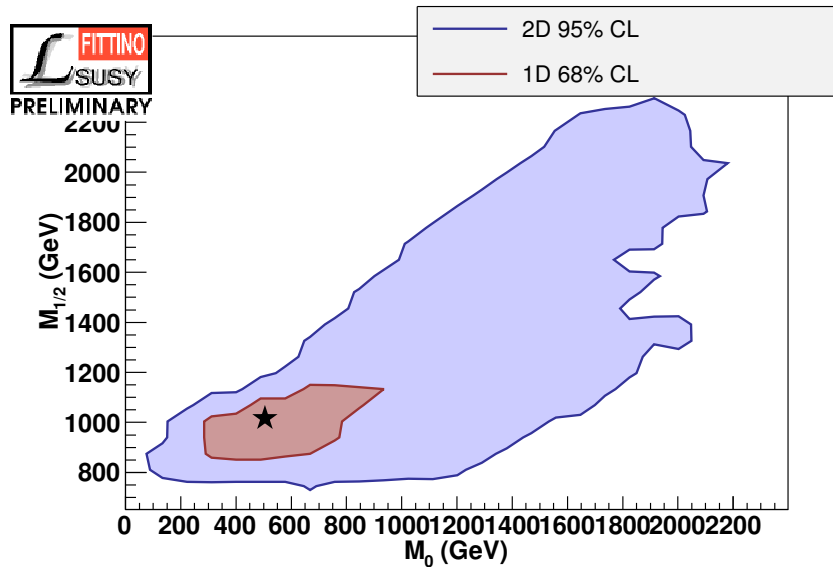


$m_{H_u}^2 \neq m_0^2 \Rightarrow \mu$  free parameter



J. Ellis, F. Luo, K. Olive, P. Sandick, arXiv:1212.4476;  $m_t = 173.2 \text{ GeV}$

Fitting low energy observables,  $m_h$ ,  $BR(h \rightarrow X)$ , LHC bounds



P. Bechtle et al., arXiv:1508.05951

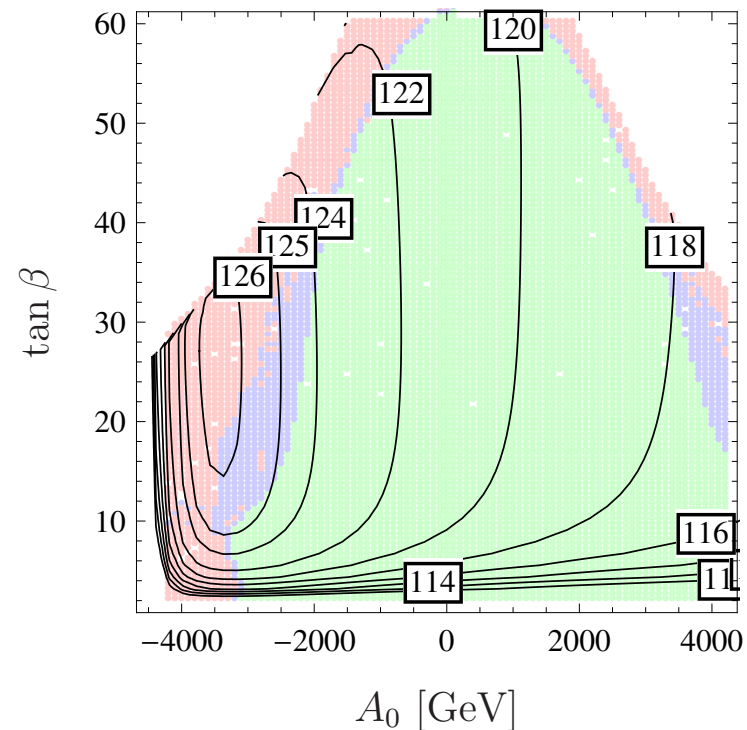
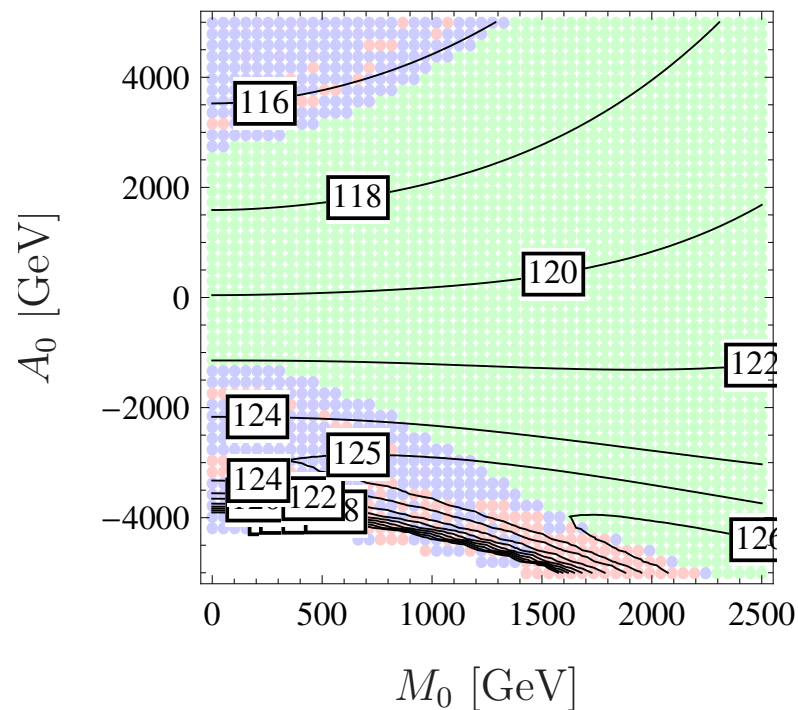
implications for LHC:  $m_{\tilde{g}}, m_{\tilde{q}} \gtrsim 2$  TeV,  $m_{\tilde{l}_R} \simeq 600$  GeV,  $m_{\tilde{\chi}_1^0} \simeq 450$  GeV

can be tested at LHC 13 TeV [14 TeV]

so far so good, but ...



- SUSY models contain many scalars  $\Rightarrow$  complicated potential
- usually some parameters ( $\mu, B$ ) are chosen to obtain correct EWSB
- does not exclude the existence of other minima breaking charge and/or color!



$$M_{1/2} = 1 \text{ TeV}, \tan \beta = 10, \mu > 0$$

$$M_{1/2} = M_0 = 1 \text{ TeV}$$

J.E. Camargo-Molina, B. O'Leary, W.P., F. Staub, arXiv:1309.7212



# VBF + MET: Compressed SUSY / DM

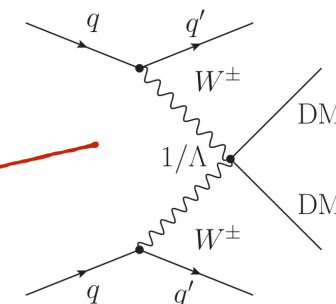
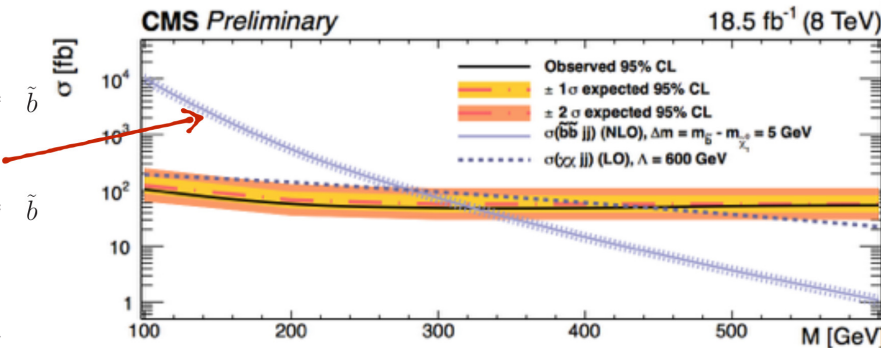
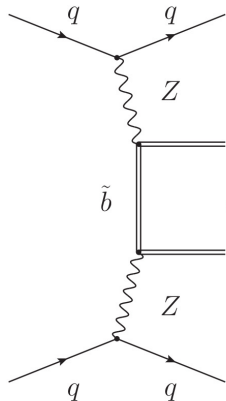
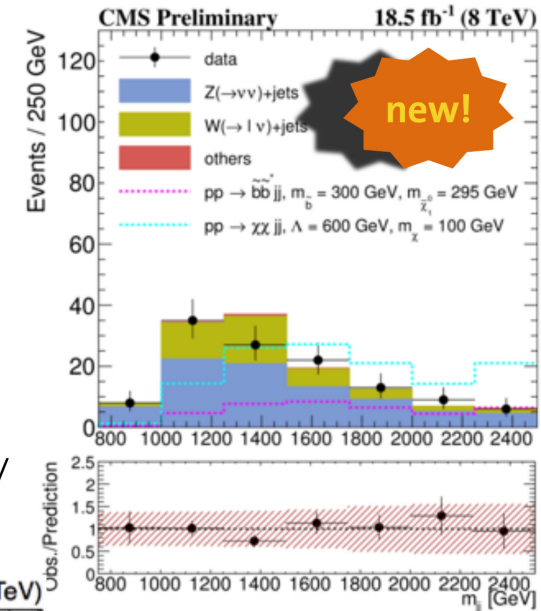


**Trigger:** MET65+VBFDiJet35

**Selection:** Two jets ( $p_T > 50$  GeV with  $\eta_1 \eta_2 < 0$ ; large rapidity gap  $|\eta_1 - \eta_2| > 4.2$  and invariant mass  $m_{12} > 750$  GeV; no b-tag); MET > 250 GeV; veto further jets ( $p_T > 30$  GeV)

**Dominant bgs:** ( $Z \rightarrow \nu\nu$ ) + jets & ( $W^\pm \rightarrow l^\pm \nu$ ) + jets estimated from data

Interpretation in models with DM production via contact interaction and  $\tilde{b}\tilde{b}\tilde{\chi}_1^0\tilde{\chi}_1^0$  production with  $m_{\tilde{b}} - m_{\tilde{\chi}_1^0} = 5$  GeV



SUS-14-019

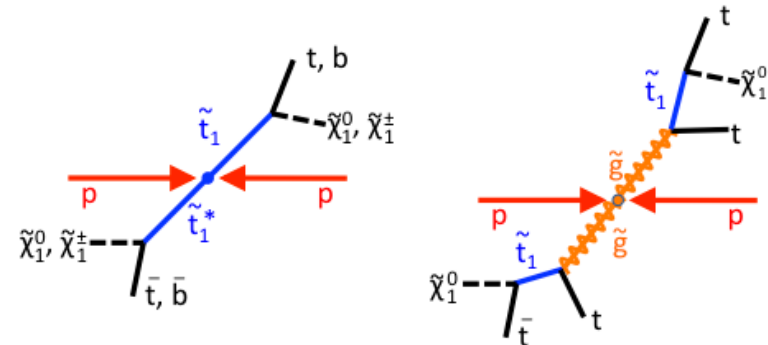
several studies, see e.g. S. Sekmen et al., arXiv:1109.5119; A. Arbey, M. Battaglia, A. Djouadi and F. Mahmoudi, arXiv:1211.4004; M. Cahill-Rowley, J. Hewett, A. Ismail and T. Rizzo, arXiv:1308.0297

- generic signatures are well known: multi-lepton, multi-jets + missing  $E_T$
- interesting feature of the 'Heavy Higgs case'  
production of  $h^0$  via SUSY cascade decays, e.g.  $\tilde{\chi}_2^0 \rightarrow h\tilde{\chi}_1^0$
- sub-class of general MSSM: 'natural SUSY' (see e.g. H. Baer, V. Barger, P. Huang, A. Mustafayev, X. Tata, arXiv:1207.3343; M. Papucci, J. T. Ruderman and A. Weiler, arXiv:1110.6926)

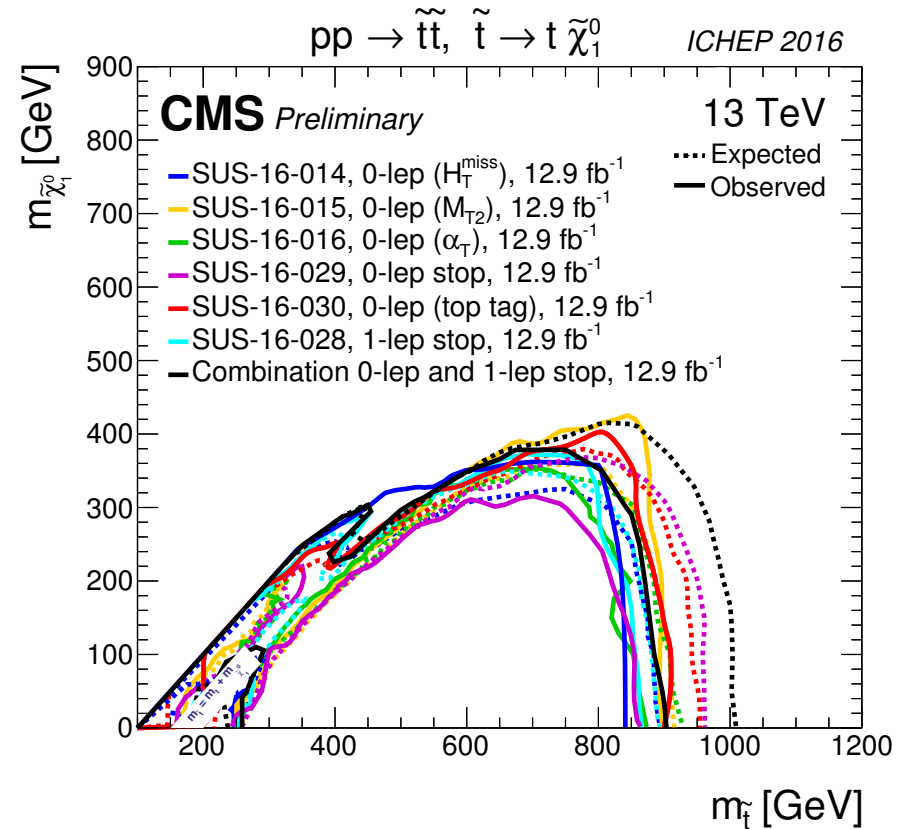
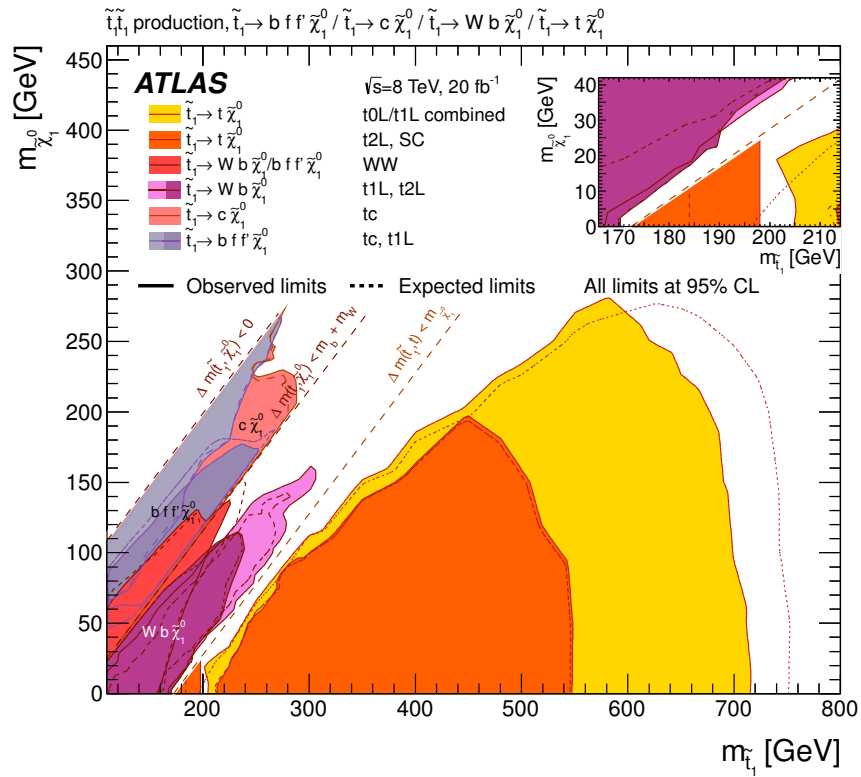
keep only SUSY particles light needed for 'natural Higgs':  $\tilde{t}_1, \tilde{b}_1, \tilde{g}, \tilde{h}^{+,0,-}$

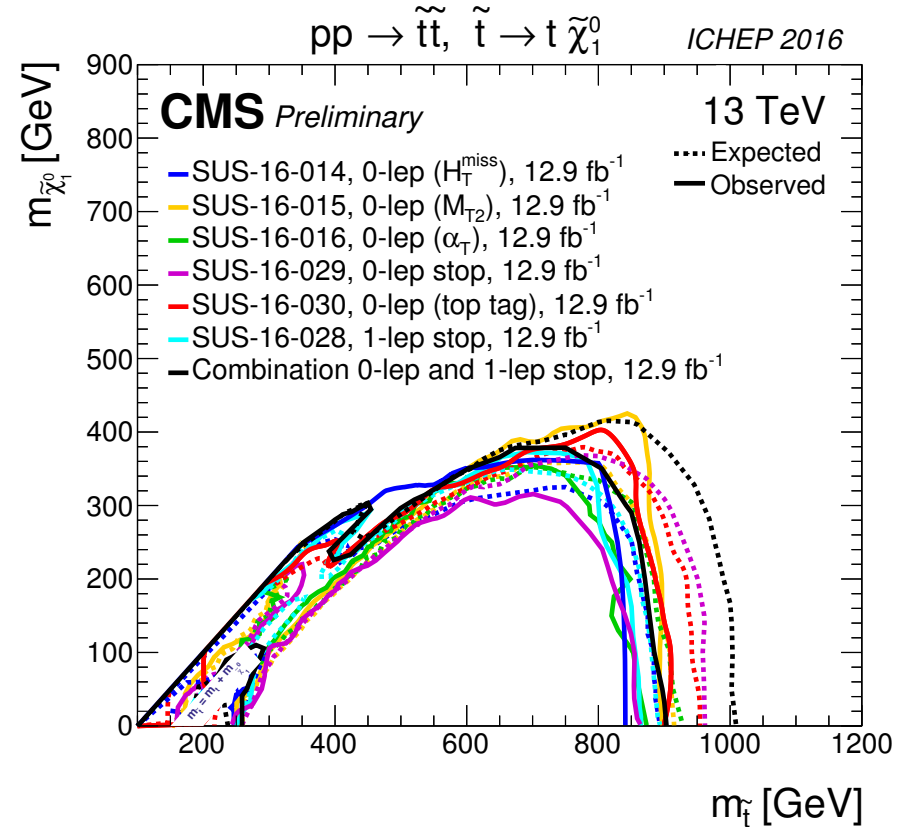
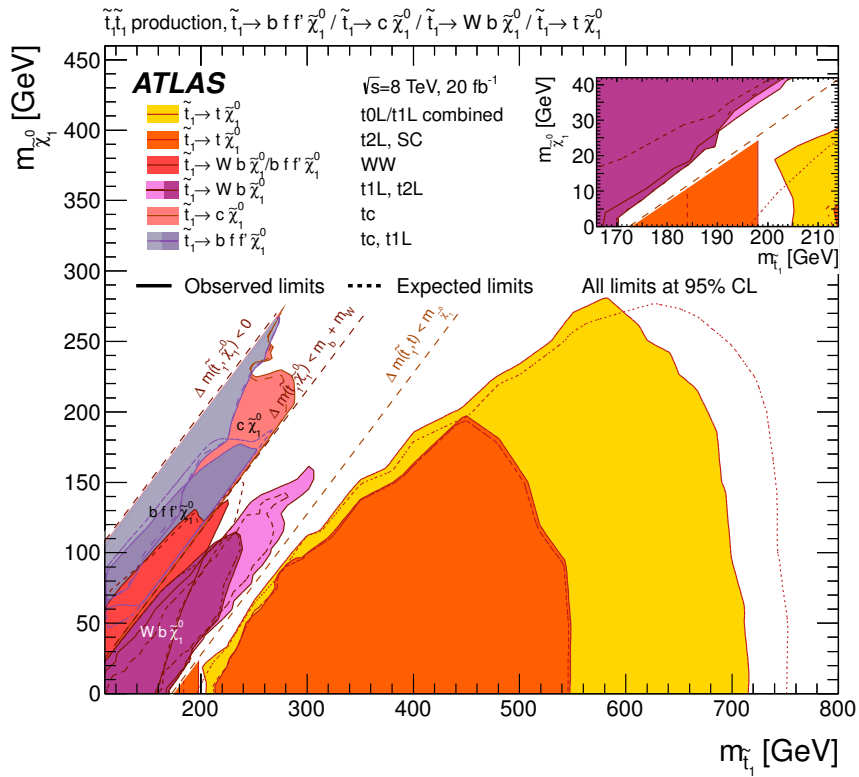
$$\Rightarrow 100 \text{ MeV} \lesssim m_{\tilde{\chi}_1^+} - m_{\tilde{\chi}_1^0} \simeq m_{\tilde{\chi}_2^0} - m_{\tilde{\chi}_1^0} \lesssim 5 - 10 \text{ GeV}$$

$$\begin{aligned} \tilde{g} &\rightarrow \tilde{t}_1 t, \tilde{b}_1 b \\ \tilde{t}_1 &\rightarrow t\tilde{\chi}_{1,2}^0, b\tilde{\chi}_1^+, W^+\tilde{b}_1 \\ \tilde{b}_1 &\rightarrow b\tilde{\chi}_{1,2}^0, t\tilde{\chi}_1^-, W^-\tilde{t}_1 \end{aligned}$$



BRs depend on the nature of  $\tilde{t}_1$  and  $\tilde{b}_1$

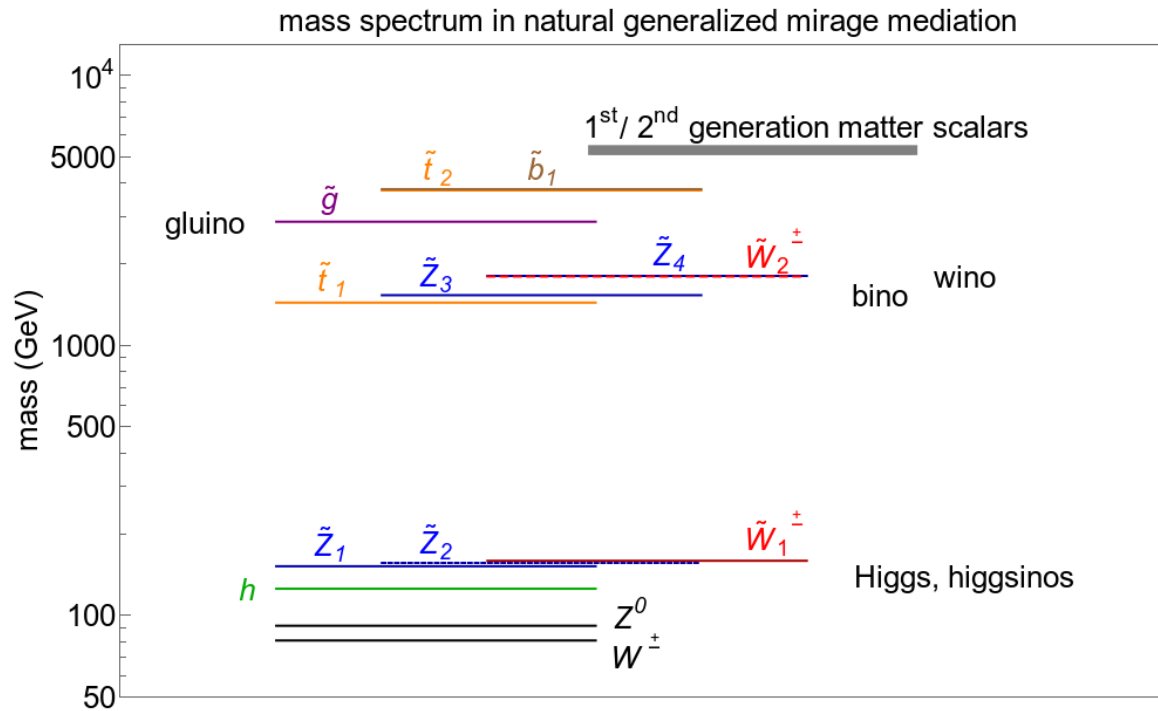




$$\frac{d}{dt} \begin{pmatrix} m_{H_u}^2 \\ m_{\tilde{t}_R}^2 \\ m_{\tilde{Q}_L^3}^2 \end{pmatrix} = -\frac{8\alpha_s}{3\pi} M_3^2 \begin{pmatrix} 0 \\ 1 \\ 1 \end{pmatrix} + \frac{Y_t^2}{8\pi^2} \left( m_{\tilde{Q}_L^3}^2 + m_{\tilde{t}_R}^2 + m_{H_u}^2 + A_t^2 \right) \begin{pmatrix} 3 \\ 2 \\ 1 \end{pmatrix}$$

Different sources for soft SUSY breaking: moduli & AMSB

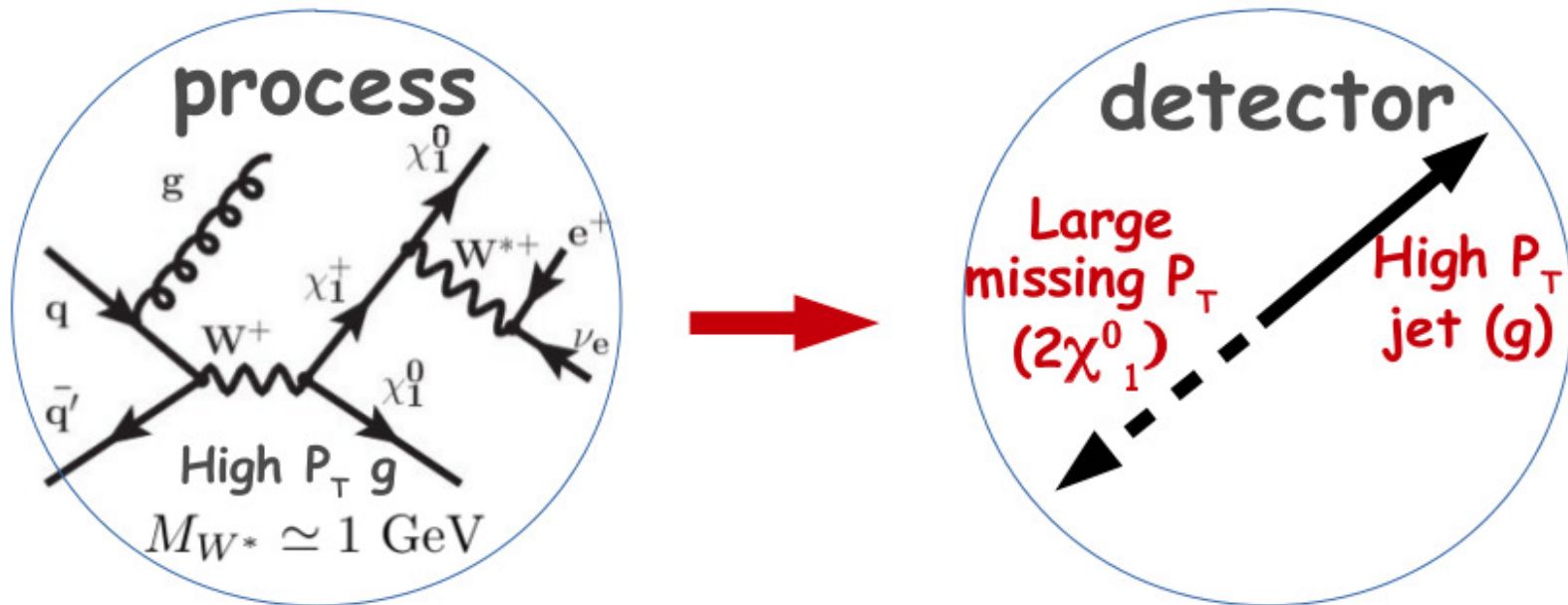
main consequence: gaugino masses unify at a (vastly) different scale than gauge couplings



H. Baer, V. Barger, H. Serce and X. Tata, arXiv:1610.06205



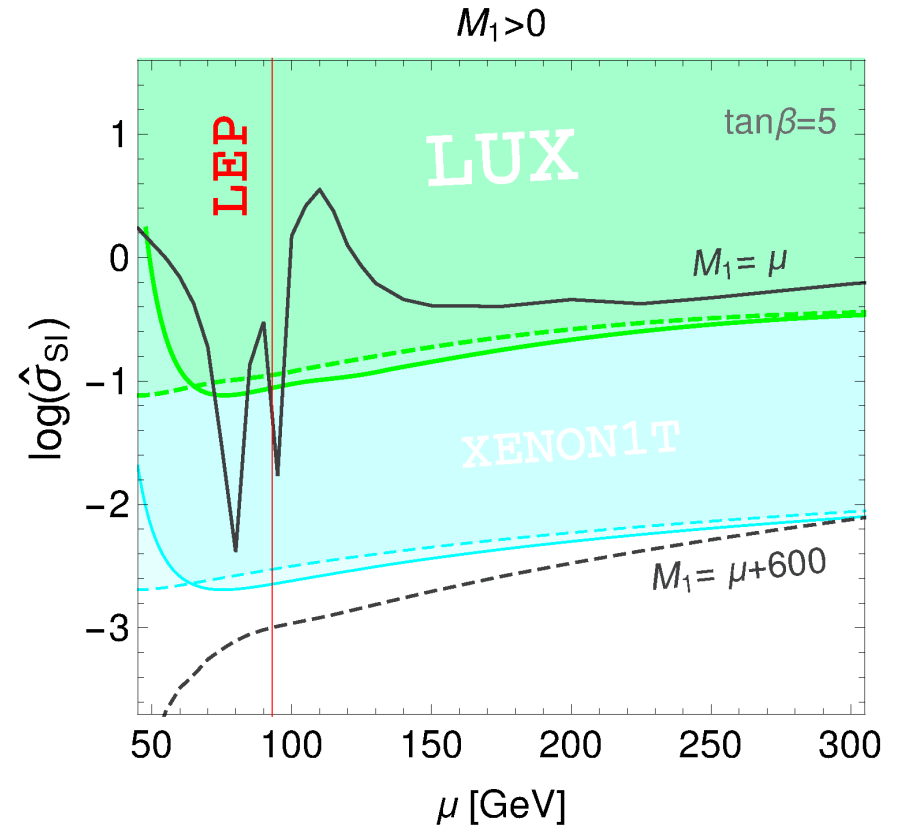
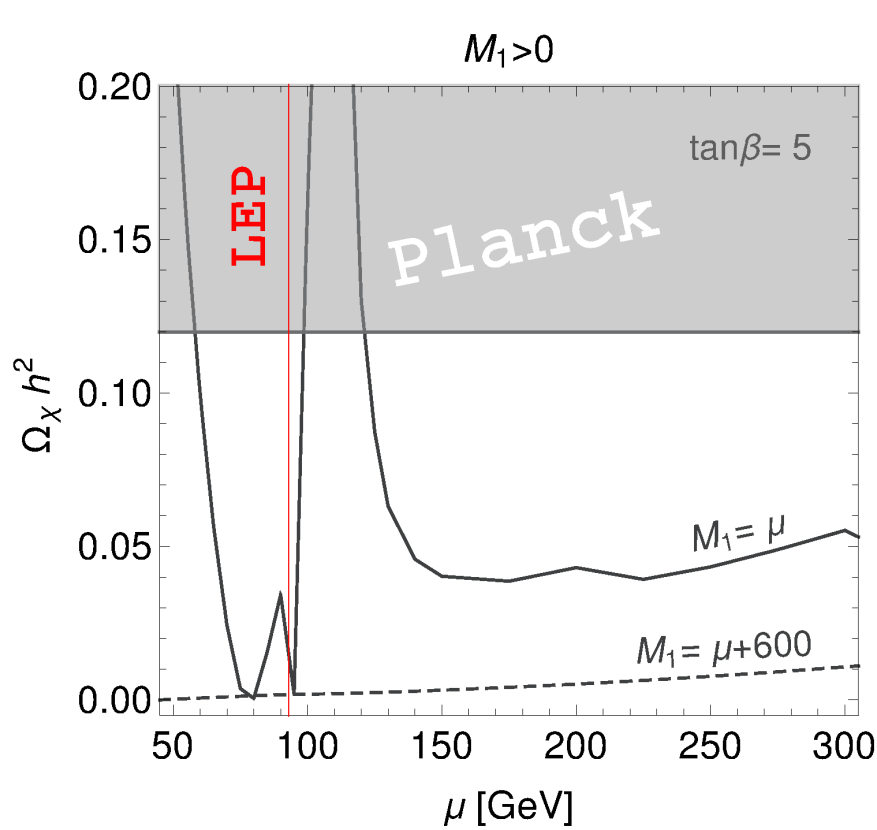
Most challenging case: only higgsinos accessible but nothing else  
and  $\Delta M$  too small for any leptonic signature



The only way to probe compressed higgsinos is a mono-jet signature:  
'Where the Sidewalk Ends? ...' Alves, Izaguirre, Wacker 2011

which has been used in studies on compressed SUSY spectra, e.g. Dreiner, Kramer, Tattersall 2012; Han, Kobakhidze, Liu, Saavedra, Wu 2013; Han, Kribs, Martin, Menon 2014



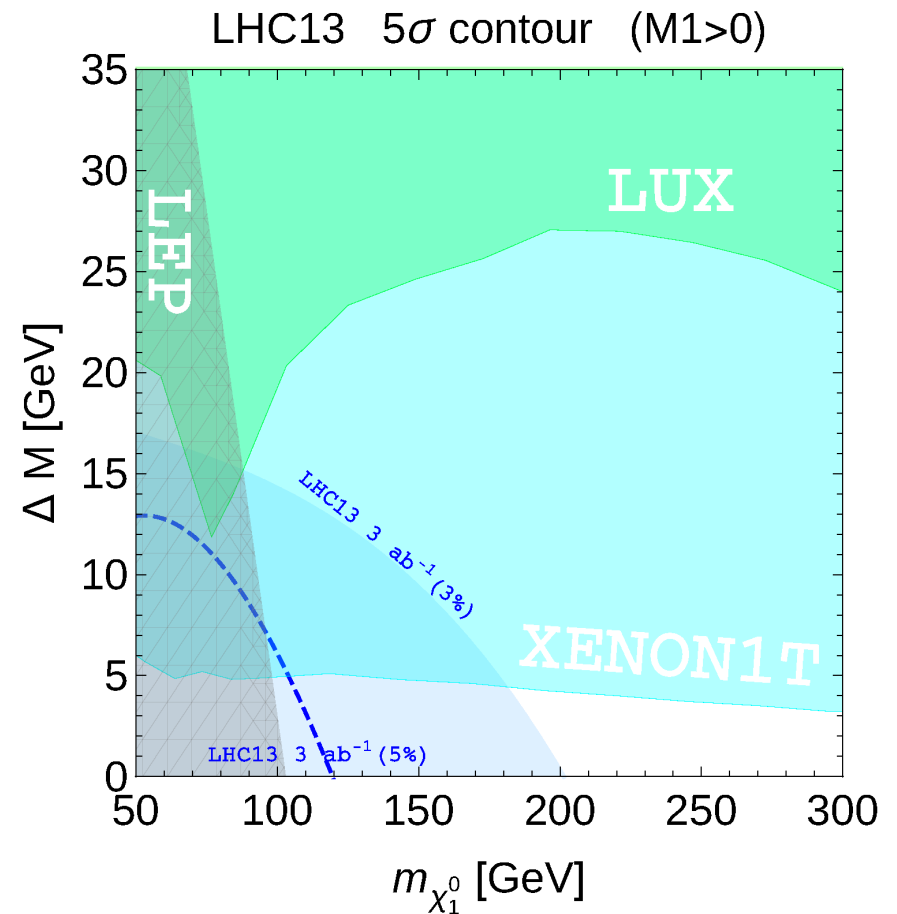
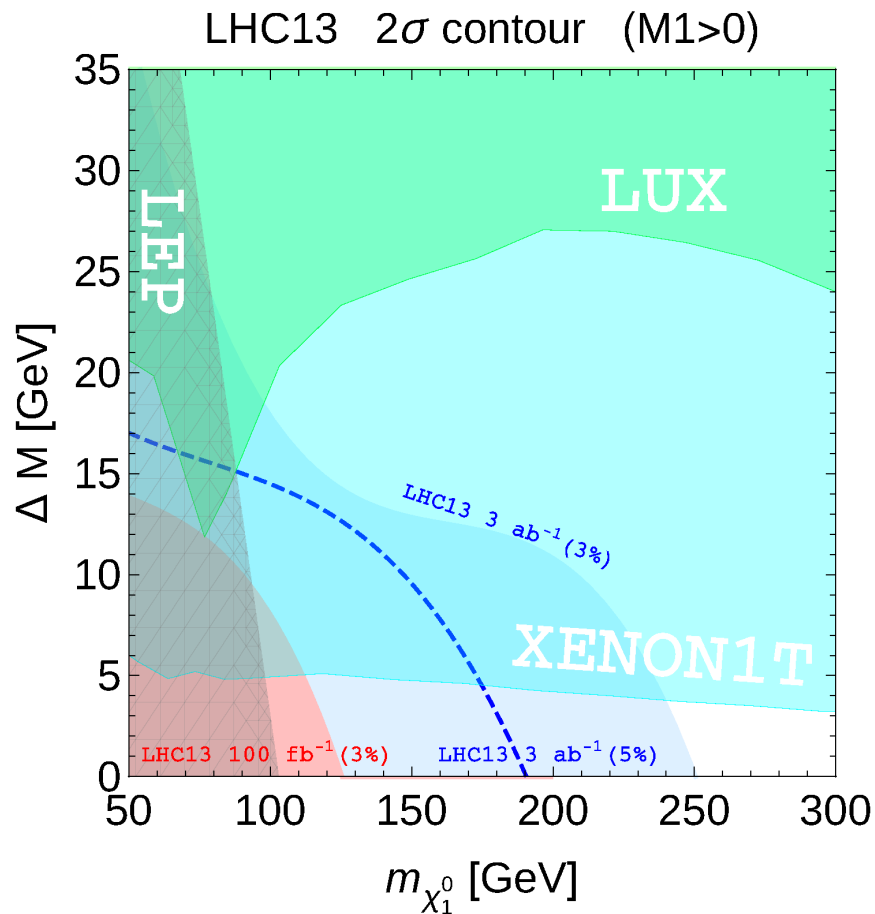


- relic density too low because higgsinos couple 'strongly' to  $W$  and  $Z$
- DD cross section rescaled with relic density  $\rightarrow$  chance for LHC?

D. Barducci, A. Belyaev, A. Bharucha, WP, V. Sanz, arXiv:1504.02472

exclusion reach

discovery reach

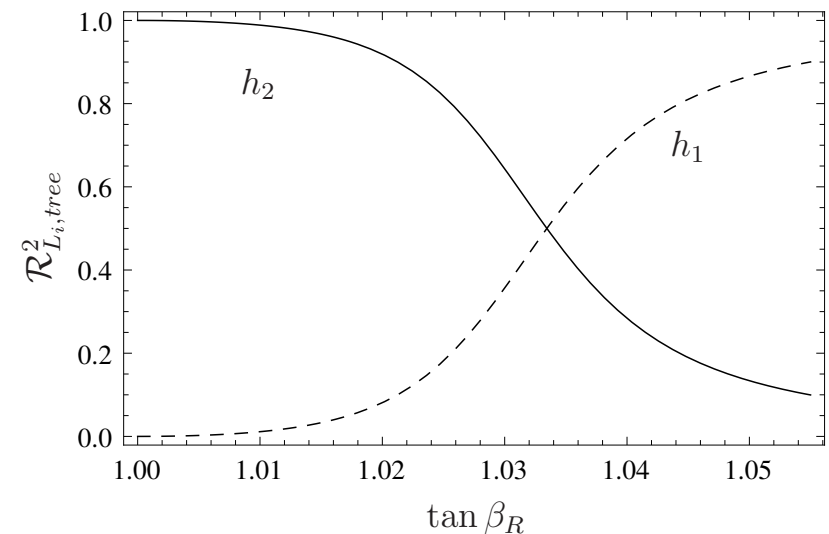
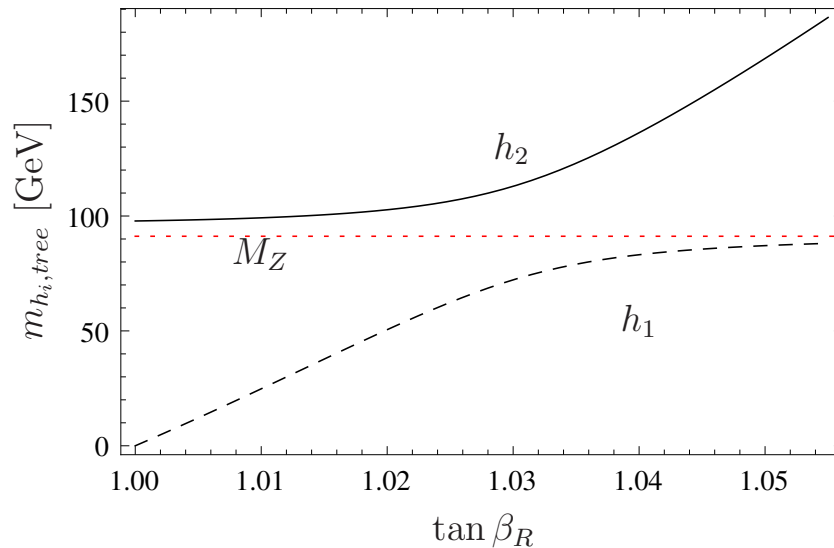


D. Barducci, A. Belyaev, A. Bharucha, WP, V. Sanz, arXiv:1504.02472

- additional D-term contributions to  $m_h$  at tree-level
- Origin of  $R$ -parity  $R_P = (-1)^{2s+3(B-L)}$ 
  - $\Rightarrow SO(10) \rightarrow SU(3)_C \times SU(2)_L \times SU(2)_R \times U(1)_{B-L}$
  - $\rightarrow SU(3)_C \times SU(2)_L \times U(1)_R \times U(1)_{B-L}$
  - $\cong SU(3)_C \times SU(2)_L \times U(1)_Y \times U(1)_X$
  - or  $E(8) \times E(8) \rightarrow SU(3)_C \times SU(2)_L \times U(1)_Y \times U(1)_{B-L}$
- Neutrino masses
  - $B - L$  anomaly free  $\Rightarrow \nu_R$
  - usual seesaw, inverse seesaw

extra  $U(1)_\chi$  with new D-term contributions at tree-level:  $m_{h_i,tree}^2 \leq m_Z^2 + \frac{1}{4}g_\chi^2 v^2$

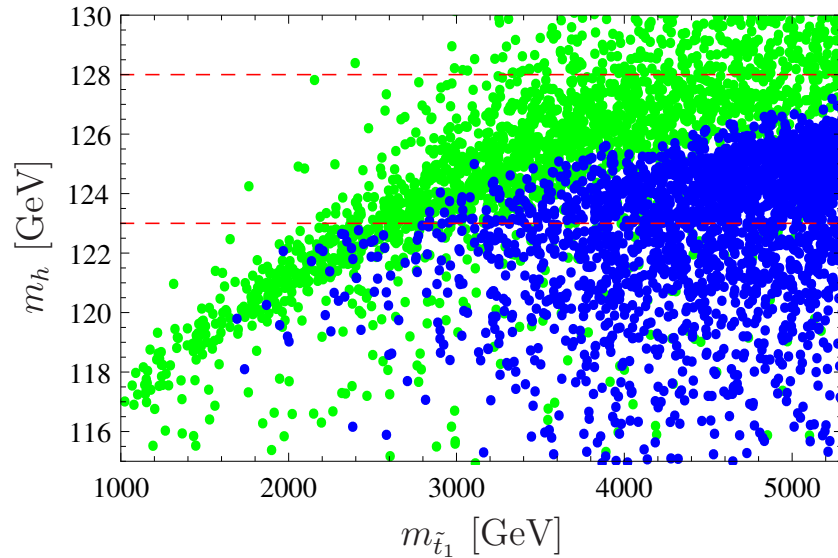
H.E. Haber, M. Sher, PRD 35 (1987) 2206; M. Drees, PRD 35 (1987) 2910; M. Cvetič et al., hep-ph/9703317; E. Ma, arXiv:1108.4029; M. Hirsch et al., arXiv:1110.3037



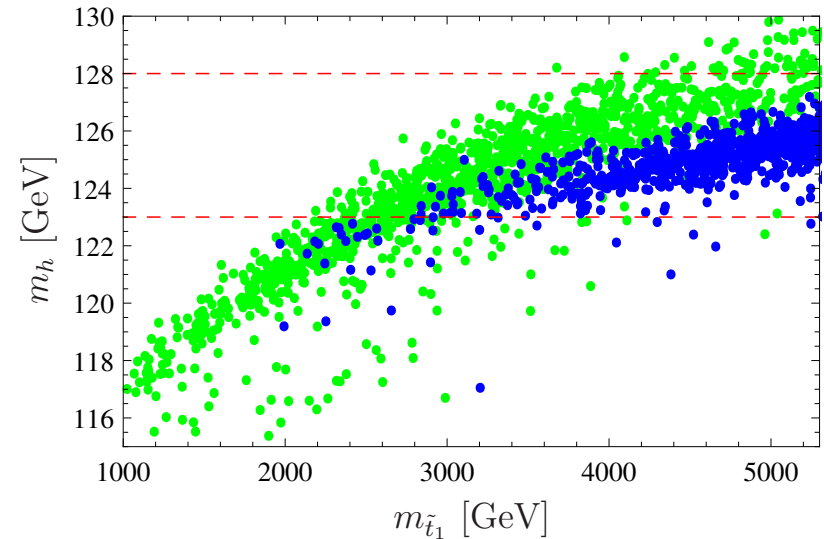
$n = 1$ ,  $\Lambda = 5 \cdot 10^5$  GeV,  $M = 10^{11}$  GeV,  $\tan \beta = 30$ ,  $\text{sign}(\mu_R) = -$ ,  $\text{diag}(Y_S) = (0.7, 0.6, 0.6)$ ,  $Y_\nu^{ii} = 0.01$ ,  $v_R = 7$  TeV

M.E. Krauss, W.P., F. Staub, arXiv:1304.0769

$$R_{h \rightarrow \gamma\gamma} \geq 0.5$$



$$R_{h \rightarrow \gamma\gamma} \geq 0.9$$



scan over GMSB parameters:  $1 \leq n \leq 4$ ,  $10^5 \leq M \leq 10^{12}$  GeV,  $10^5 \leq \sqrt{n}\Lambda \leq 10^6$  GeV,  
 $1.5 \leq \tan \beta \leq 40$ ,  $1 < \tan \beta_R \leq 1.15$ ,  $\text{sign}(\mu_R) \pm 1$ ,  $\text{sign}(\mu) = 1$ ,  $6.5 \leq v_R \leq 10$  TeV,  
 $0.01 \leq Y_S^{ii} \leq 0.8$ ,  $10^{-5} \leq Y_\nu^{ii} \leq 0.5$   
 blue points:  $h_1 \simeq h$ , green points:  $h_2 \simeq h$

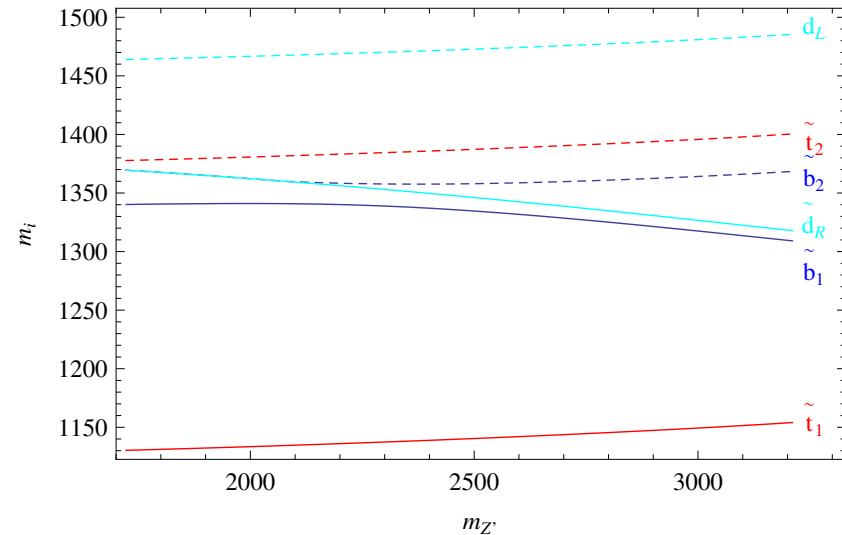
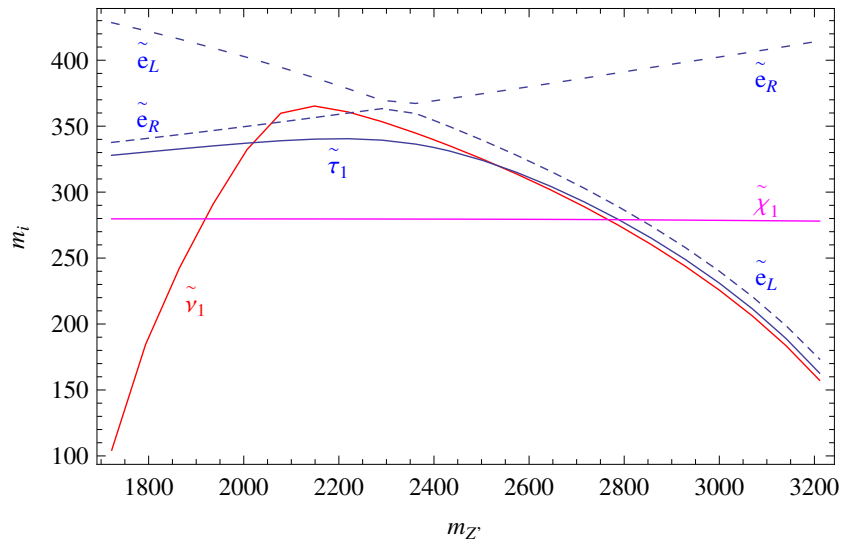
$$R_{h \rightarrow \gamma\gamma} = \frac{[\sigma(pp \rightarrow h) \times BR(h \rightarrow \gamma\gamma)]_{BLR}}{[\sigma(pp \rightarrow h) \times BR(h \rightarrow \gamma\gamma)]_{SM}}$$

M.E. Krauss, W.P., F. Staub, arXiv:1304.0769

$$M_{\tilde{l}}^2 = \begin{pmatrix} M_{\tilde{L}}^2 + D_L + m_l^2 & \frac{1}{\sqrt{2}} (v_d T_l - \mu Y_l v_u) \\ \frac{1}{\sqrt{2}} (v_d T_l - \mu Y_l v_u) & M_{\tilde{E}}^2 + D_R + m_l^2 \end{pmatrix},$$

$$D_L \simeq \left(-\frac{1}{2} + \sin^2_{\theta_W}\right) m_Z^2 c_{2\beta} - \frac{5}{4} m_{Z'}^2 c_{2\beta_R} \quad \text{and} \quad D_R \simeq -\sin^2_{\theta_W} m_Z^2 c_{2\beta} + \frac{5}{4} m_{Z'}^2 c_{2\beta_R}$$

neglecting gauge kinetic effects; similarly for squarks



$$m_0 = 100 \text{ GeV}, m_{1/2} = 700 \text{ GeV}, A_0 = 0, \tan \beta = 10, \mu > 0$$

$$\tan \beta_R = 0.94, m_{A_R} = 2 \text{ TeV}, \mu_R = -800 \text{ GeV}$$

	BLRSP1	BLRSP2	BLRSP3	BLRSP4	BLRSP5
$m_{\tilde{\nu}_1}$	105.0	797.	91.6	542.	921.
$m_{\tilde{\nu}_{2/3}}$	215.0	797.	92.6	542.	924.
$m_{\tilde{\nu}_4}$	604.	1120.	253.	585.	940.
$m_{\tilde{e}_1}$	524.	1014.	255.	263.	693.
$m_{\tilde{e}_{2,3}}$	557.	1055.	266.	271.	706.
$m_{\tilde{u}_1}$	1436.	1185.	1247.	1111.	1545.
$m_{\tilde{u}_2}$	1721.	1852.	1527.	1361.	1905.
$m_{\tilde{u}_{3,4}}$	1799.	2155.	1566.	1392.	2008.
$m_{\chi_1^0}$	367.	417.	313.	259. $\tilde{h}_R$	412.
$m_{\chi_2^0}$	718.	780. $\tilde{h}_R$	615.	280.	739. $\tilde{h}_R$
$m_{\chi_3^0}$	1047.	818.	1087.	549.	804.
$m_{\chi_5^0}$	1348. ( $\tilde{B}_\perp$ )	1866.	1232. ( $\tilde{B}_\perp$ )	857.	1294.
$m_{\chi_6^0}$	1802. $\tilde{h}_R$	2018. ( $\tilde{B}_\perp$ )	1811. ( $\tilde{B}_\perp$ )	1639. ( $\tilde{B}_\perp$ )	1688. ( $\tilde{B}_\perp$ )

B. O'Leary, W.P., F. Staub, arXiv:1112.4600



CMSSM, GMSB:  $\tilde{q}_R \rightarrow q\tilde{\chi}_1^0$

BLRSP1:  $\tilde{\nu}$  LSP,  $m_{\nu_h} \simeq 100$  GeV

$$\tilde{q}_R \rightarrow q\tilde{\chi}_1^0 \rightarrow q\nu_k\tilde{\nu}_1 \rightarrow q\nu_j Z\tilde{\nu}_1 \quad (k = 4, \dots, 9, j = 1, 2, 3)$$

$$\tilde{q}_R \rightarrow q\tilde{\chi}_1^0 \rightarrow q\nu_k\tilde{\nu}_1 \rightarrow ql^\pm W^\mp \tilde{\nu}_1$$

$$\tilde{q}_R \rightarrow q\tilde{\chi}_1^0 \rightarrow q\nu_k\tilde{\nu}_3 \rightarrow ql^\pm W^\mp l'^+ l'^- \tilde{\nu}_1$$

$$\tilde{d}_R \rightarrow d\tilde{\chi}_5^0 \rightarrow dl^\pm \tilde{l}_i^\mp \rightarrow dl^\pm l^\mp \tilde{\chi}_1^0 \rightarrow dl^\pm l^\mp \nu_k \tilde{\nu}_1 \rightarrow dl^\pm l^\mp l'^\pm W^\mp \tilde{\nu}_1$$

BLRSP3: usual cascades similar to CMSSM, but

$$\tilde{\chi}_1^0 \rightarrow l^\pm \tilde{l}^\mp \rightarrow l^\pm W^\mp \tilde{\nu}_1$$

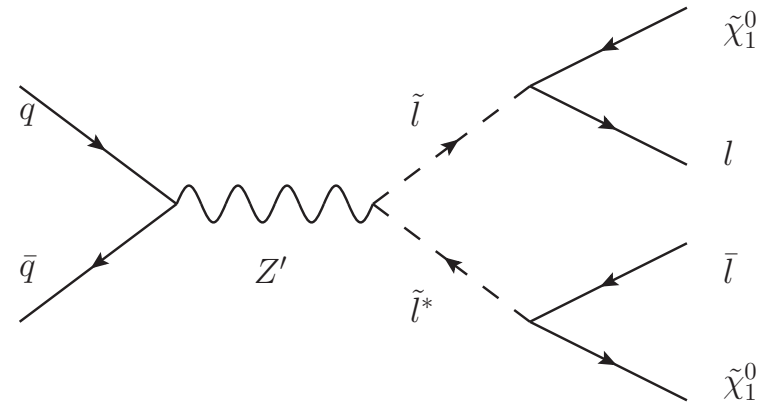
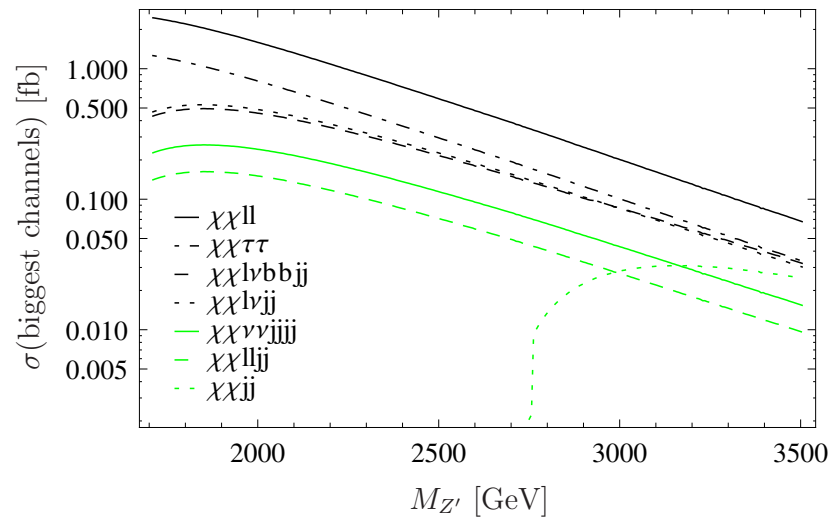
$$\tilde{\chi}_1^0 \rightarrow l^\pm \tilde{l}^\mp \rightarrow l^\pm W^\mp \tilde{\nu}_{2,3} \rightarrow l^\pm W^\mp f\bar{f}\tilde{\nu}_1$$

$$\tilde{\chi}_1^0 \rightarrow \nu_j \tilde{\nu}_{2,3} \rightarrow \nu_{1,2,3} f\bar{f}\tilde{\nu}_1$$

$$\tilde{\chi}_1^0 \rightarrow \nu_j \tilde{\nu}_k \rightarrow \nu_j h_{1,2} \tilde{\nu}_1 \quad (j, k = 1, 2, 3)$$

$$\tilde{\chi}_1^0 \rightarrow \nu_j \tilde{\nu}_k \rightarrow \nu_j h_{1,2} f\bar{f}\tilde{\nu}_1$$

⇒ enhanced jet and lepton multiplicities, study of  $\nu_R$



M. Krauss, B. O'Leary, W.P., F. Staub, arXiv:1206.3513

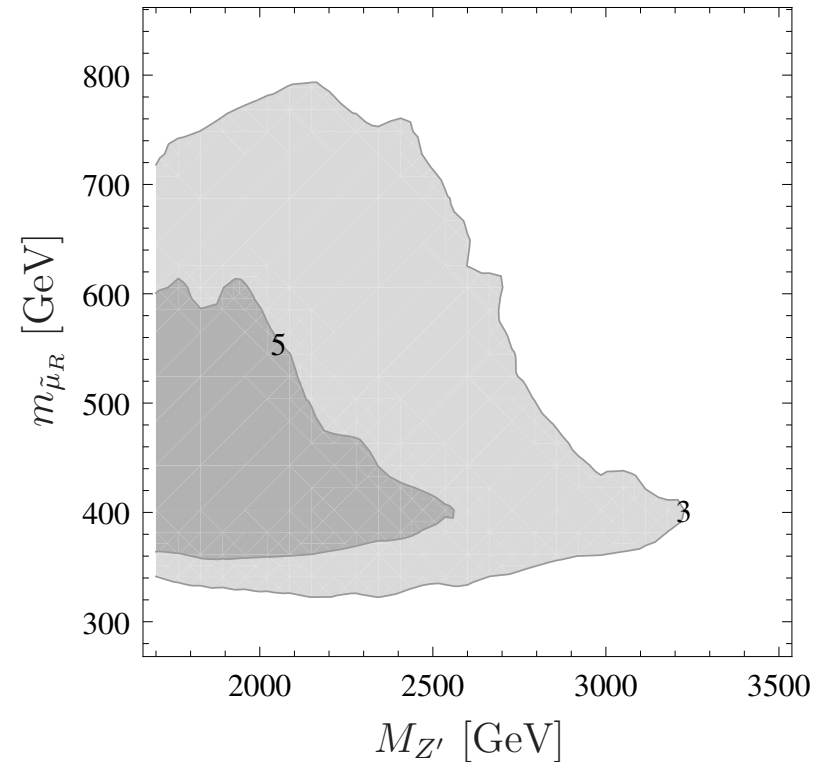
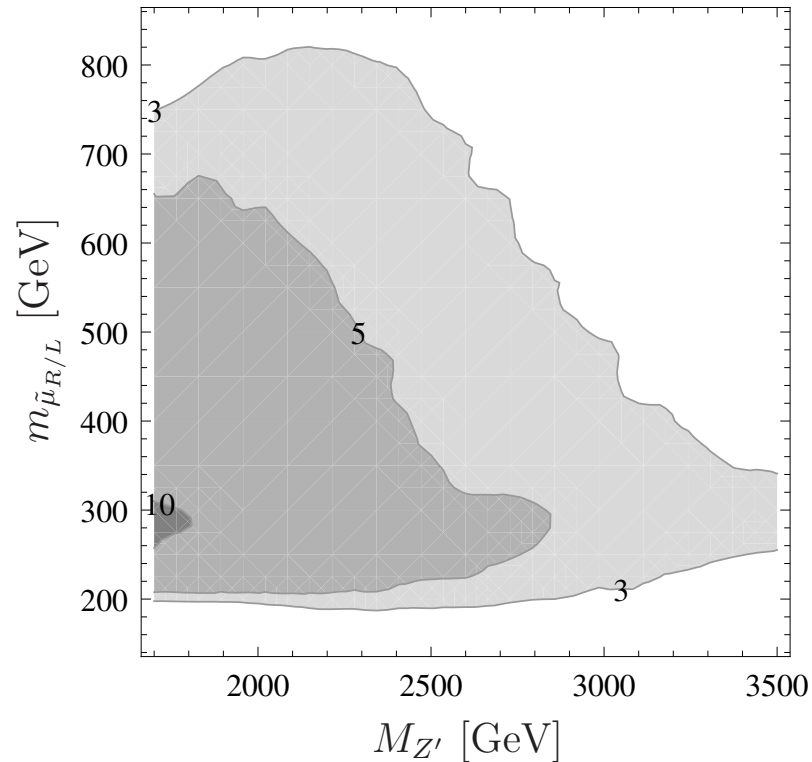
see also: J. Kang and P. Langacker, PRD 71 (2005) 035014; M. Baumgart, T. Hartman, C. Kilic, and L.-T. Wang, JHEP 0711 (2007) 084; C.-F. Chang, K. Cheung, and T.-C. Yuan, JHEP 1109 (2011) 058; G. Corcella and S. Gentile, arXiv:1205.5780

main dependence on masses  $\Rightarrow$  vary  $m_{\tilde{l}}$  and  $m_{Z'}$ ,  $M_L = 1.2M_E$

$100 \text{ fb}^{-1}$ ,  $\sqrt{s} = 14 \text{ TeV}$

$m_{\tilde{\chi}_1^0} = 140 \text{ GeV}$

$m_{\tilde{\chi}_1^0} = 280 \text{ GeV}$



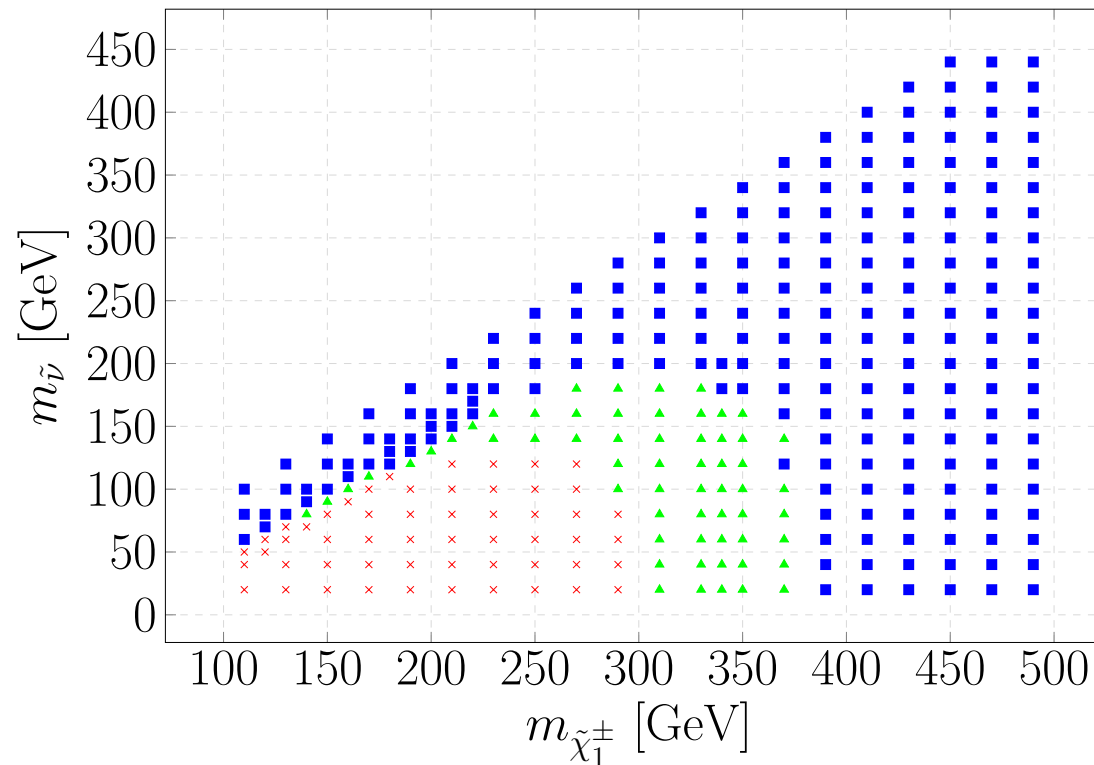
M. Krauss, B. O'Leary, W.P., F. Staub, arXiv:1206.3513

- $m_{\tilde{t}_1}$  in GeV: 300, 400, 500, 600, 700, 800, 900, 1000
- $m_{\tilde{b}_1}$  in GeV: 300, 400, 500, 600, 700, 800, 900, 1000
- $m_{\tilde{\nu}_R}$  in GeV : 60, 100, 200, 300, 400, 500
- $\mu$  in GeV: 110, 190, 290, 390, 490, 590 and require  $m_{\tilde{\nu}_R} < \mu$
- $\tan \beta$ : 10, 50
- $\theta_{\tilde{t}}$ :  $0^\circ, 45^\circ, 90^\circ$
- $\theta_{\tilde{b}}$ :  $0^\circ, 45^\circ, 90^\circ$
- $M_1 = M_2 = 1$  TeV
- everything else, including  $\tilde{t}_2, \tilde{b}_2$  and  $m_{\tilde{g}}$ : 2 TeV  
The exception is potentially  $m_{\tilde{b}_2}$  in case of  $\theta_{\tilde{t}} = 0$

$$m_W^2 \cos 2\beta = m_{\tilde{t}_1}^2 - m_{\tilde{b}_1}^2 \cos^2 \theta_{\tilde{b}} - m_{\tilde{b}_2}^2 \sin^2 \theta_{\tilde{b}} - m_t^2 + m_b^2$$

$\Rightarrow m_{\tilde{b}_2} \leftrightarrow m_{\tilde{b}_1}$  if necessary

$$pp \rightarrow \tilde{\chi}_1^+ \tilde{\chi}_1^- \rightarrow \ell^+ \ell^- \tilde{\nu}_R \tilde{\nu}_R^*$$



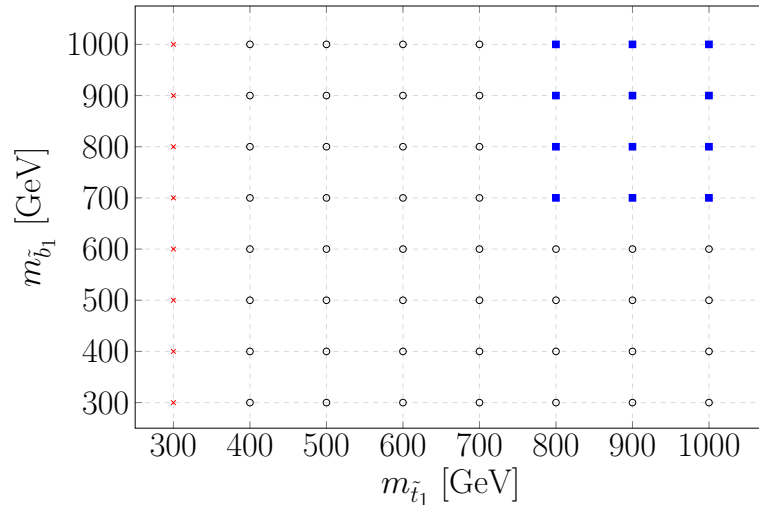
× excluded, ▲ ambiguous, ■ allowed

8 TeV data using CheckMATE

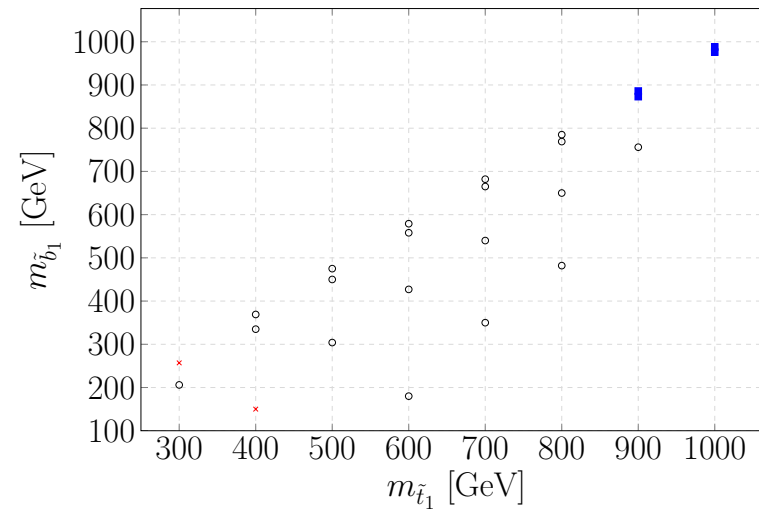
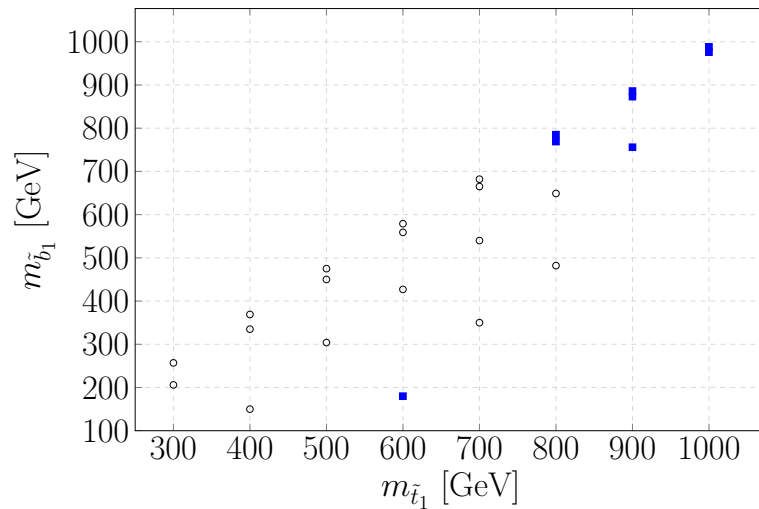
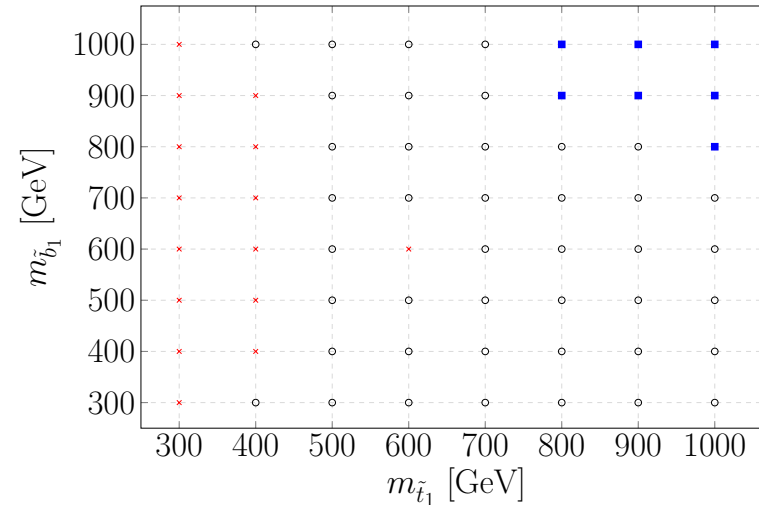
L. Mitzka, WP arXiv:1603.06130

13 TeV update: ongoing work with N. Cerna Velezco, T. Faber, J. Jones

ambiguous as allowed



ambiguous as forbidden



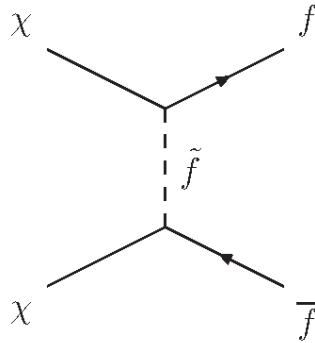
× excluded for all parameters, ○ exclusion depends on parameters, ■ allowed for all parameters

L. Mitzka, WP arXiv:1603.06130

- $m_h = 125.1$  GeV & SUSY: either large radiative corrections or additional tree-level contributions in models beyond MSSM
- MSSM particle content
  - GMSB: beyond LHC reach (minimal version)
  - CMSSM: expect  $m_{\tilde{g}}, m_{\tilde{q}} \gtrsim 2$  GeV, excluded with 90% CL if all data combined
  - ‘Natural SUSY’: take only those states light which contribute to EWSB:  
 $\tilde{h}^{0,\pm}, \tilde{t}_1, \tilde{g}, \tilde{b}_i$
  - extreme case with higgsinos only:
    - very challenging: DM direkt detection and LHC probe complementary parameter space regions
    - LHC: can discover higgsinos up to  $|\mu| \simeq 120$  GeV (200 GeV) for  $\mathcal{L}=3$  ab<sup>-1</sup>  
Clear need for  $e^+e^-$  collider
  - light stop still consistent with data
  - general MSSM: predictions depend strongly on details
  - models with large  $A_t, A_b$ : problems with charge/color breaking minima

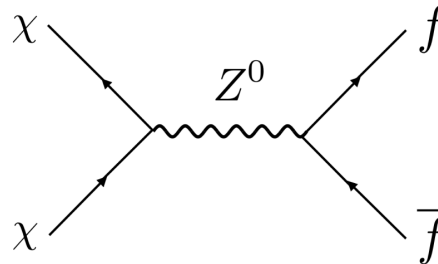


- extended gauge groups
  - also motivated by  $\nu$ -physics  $\Rightarrow$  extended (s)neutrino sector
  - GMSB-like realisation: testable at LHC but heavy  $\tilde{g}, \tilde{q}$
  - CMSSM-like realisation: different spectrum compared to CMSSM  $\Rightarrow$  substantial changes of cascade decays
  - $W'$  and  $Z'$  might look differently than expected & might even serve as SUSY discovery channel
  - potentially indications of SUSY in  $Z'$  decays
  - $\tilde{\nu}_R$  LSP: compatible with DM, no direct DM constraint apply
  - findings based on 8 TeV
    - $m_{\tilde{H}^\pm} \lesssim 290$  GeV excluded if  $m_{\tilde{H}^\pm} - m_{\tilde{\nu}_R} \gtrsim 150$  GeV  
13 TeV update in coll. with Nhell and Joel
    - independent of other parameters:  $m_{\tilde{t}_1} \lesssim 300$  GeV excluded
    - for  $300$  GeV  $\lesssim m_{\tilde{t}_1} \lesssim 800$  GeV: exclusion depends on parameters, in particular on  $\cos \theta_{\tilde{t}}$



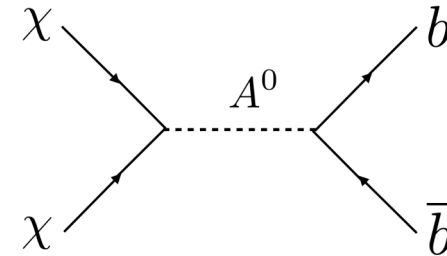
bino

bulk region

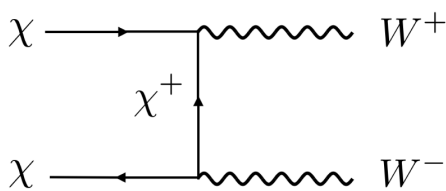


wino, higgsino

focus-point region

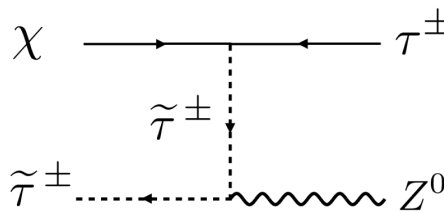


funnel region

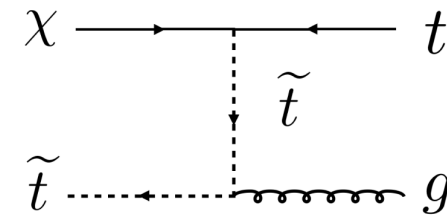


wino, higgsino

focus-point region



stau co-annihilation



stop co-annihilation

Constraints from  $Z$ -width:  $m_{\nu_h} \gtrsim m_Z$

invisible width

$$\left| 1 - \sum_{ij=1, i \leq j}^3 \left| \sum_{k=1}^3 U_{ik}^\nu U_{jk}^{\nu,*} \right|^2 \right| < 0.009$$

dominant decays

$$\nu_j \rightarrow W^\pm l^\mp$$

$$\nu_j \rightarrow Z \nu_i$$

$$\nu_j \rightarrow h_k \nu_i$$

roughly

$$BR(\nu_j \rightarrow W^\pm l^\mp) : BR(\nu_j \rightarrow Z \nu_i) : BR(\nu_j \rightarrow h_k \nu_i) \simeq 0.5 : 0.25 : 0.25$$

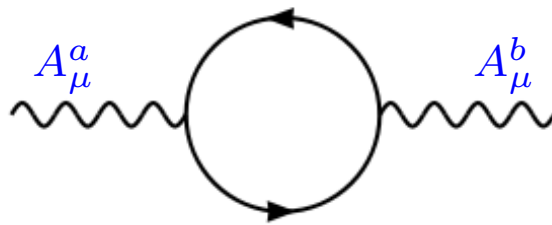
in BLRSP4

$$BR(\nu_k \rightarrow \tilde{\nu}_i \tilde{\chi}_1^0) \simeq 0.03 \quad , (k = 4, 5, 6) \text{ and } (i = 1, 2, 3)$$

$U(1)_a \times U(1)_b$  models allow for

(B. Holdom, PLB 166m0 = 250 (1986) 196)

$$\mathcal{L} \supset -\chi_{ab} \hat{F}^{a,\mu\nu} \hat{F}_{\mu\nu}^b$$



$$\iff \gamma_{ab} = \frac{1}{16\pi^2} \text{Tr}(Q_a Q_b)$$

equivalent

$$D_\mu = \partial_\mu - i(Q_a, Q_b) \underbrace{\begin{pmatrix} g_{aa} & g_{ab} \\ g_{ba} & g_{bb} \end{pmatrix}}_{NG} \begin{pmatrix} A_\mu^a \\ A_\mu^b \end{pmatrix}$$

both  $U(1)$  unbroken  $\Rightarrow$  chose basis with e.g.  $g_{ba} = 0$

affects also RGE running of soft SUSY parameters:

R. Fonseca, M. Malinsky, W.P., F. Staub, NPB 854 (2012) 28

basis  $(W^0, B_Y, B_\chi)$

$$M_{VV}^2 = \frac{1}{4} \begin{pmatrix} g_2^2 v^2 & -g_2 g' v^2 & g_2 \tilde{g}_\chi v^2 \\ -g_2 g' v^2 & g'^2 v^2 & -g' \tilde{g}_\chi v^2 \\ g_2 \tilde{g}_\chi v^2 & -g' \tilde{g}_\chi v^2 & \frac{25}{4} g_\chi^2 v_R^2 + \tilde{g}_\chi^2 v^2 \end{pmatrix}$$

$$\tilde{g}_\chi = g_\chi - g_{Y\chi}$$

$$v^2 = v_d^2 + v_u^2, \quad v_R^2 = v_{\chi R}^2 + v_{\tilde{\chi} R}^2$$

expanding in  $v^2/v_R^2$

$$m_Z^2 \simeq \frac{1}{4} (g'^2 + g_2^2) v^2 \left( 1 - \frac{4}{25} \left( 1 - \frac{g_{Y\chi}}{g_\chi} \right)^2 \frac{v^2}{v_R^2} \right)$$

$$m_{Z'}^2 \simeq \left( \frac{5}{4} g_\chi v_R \right)^2$$

M. Hirsch, W.P., L. Reichert, F. Staub, arXiv:1206:3516;

M.E. Krauss, W.P., F. Staub, arXiv:1304.0769

basis  $(\lambda_{BL}, \lambda_L^0, \tilde{h}_d^0, \tilde{h}_u^0, \lambda_R, \tilde{\chi}_R, \tilde{\chi}_R)$

$M_{\tilde{\chi}^0} =$

$$\begin{pmatrix} M_{BL} & 0 & -\frac{1}{2}g_{RBL}v_d & \frac{1}{2}g_{RBL}v_u & \frac{M_{BLR}}{2} & \frac{1}{2}v_{\tilde{\chi}_R}\tilde{g}_{BL} & -\frac{1}{2}v_{\chi_R}\tilde{g}_{BL} \\ 0 & M_2 & \frac{1}{2}g_2v_d & -\frac{1}{2}g_2v_u & 0 & 0 & 0 \\ -\frac{1}{2}g_{RBL}v_d & \frac{1}{2}g_2v_d & 0 & -\mu & -\frac{1}{2}g_Rv_d & 0 & 0 \\ \frac{1}{2}g_{RBL}v_u & -\frac{1}{2}g_2v_u & -\mu & 0 & \frac{1}{2}g_Rv_u & 0 & 0 \\ \frac{M_{BLR}}{2} & 0 & -\frac{1}{2}g_Rv_d & \frac{1}{2}g_Rv_u & M_R & -\frac{1}{2}v_{\tilde{\chi}_R}\tilde{g}_R & \frac{1}{2}v_{\chi_R}\tilde{g}_R \\ \frac{1}{2}v_{\tilde{\chi}_R}\tilde{g}_{BL} & 0 & 0 & 0 & -\frac{1}{2}v_{\tilde{\chi}_R}\tilde{g}_R & 0 & -\mu_R \\ -\frac{1}{2}v_{\chi_R}\tilde{g}_{BL} & 0 & 0 & 0 & \frac{1}{2}v_{\chi_R}\tilde{g}_R & -\mu_R & 0 \end{pmatrix}$$

$$\chi_R = \frac{1}{\sqrt{2}} (\sigma_R + i\varphi_R + v_{\chi_R}) , \quad \bar{\chi}_R = \frac{1}{\sqrt{2}} (\bar{\sigma}_R + i\bar{\varphi}_R + v_{\bar{\chi}_R})$$

$$H_d^0 = \frac{1}{\sqrt{2}} (\sigma_d + i\varphi_d + v_d) , \quad H_u^0 = \frac{1}{\sqrt{2}} (\sigma_u + i\varphi_u + v_u)$$

pseudo scalars, basis  $(\varphi_d, \varphi_u, \bar{\varphi}_R, \varphi_R)$

$$M_{AA}^2 = \begin{pmatrix} M_{AA,L}^2 & 0 \\ 0 & M_{AA,R}^2 \end{pmatrix}$$

$$M_{AA,L}^2 = B_\mu \begin{pmatrix} \tan \beta & 1 \\ 1 & \cot \beta \end{pmatrix} , \quad M_{AA,R}^2 = B_{\mu_R} \begin{pmatrix} \tan \beta_R & 1 \\ 1 & \cot \beta_R \end{pmatrix}$$

$\tan \beta = v_u/v_d$  and  $\tan \beta_R = v_{\chi_R}/v_{\bar{\chi}_R}$   
two physical states

$$m_A^2 = B_\mu (\tan \beta + \cot \beta) , \quad m_{A_R}^2 = B_{\mu_R} (\tan \beta_R + \cot \beta_R)$$

independent of gauge kinetic mixing!



$$M_{hh}^2 = \begin{pmatrix} m_{LL}^2 & m_{LR}^2 \\ m_{LR}^{2,T} & m_{RR}^2 \end{pmatrix}$$

$$m_{LL}^2 = \begin{pmatrix} g_\Sigma^2 v^2 c_\beta^2 + m_A^2 s_\beta^2 & -\frac{1}{2} (m_A^2 + g_\Sigma^2 v^2) s_{2\beta} \\ -\frac{1}{2} (m_A^2 + g_\Sigma^2 v^2) s_{2\beta} & g_\Sigma^2 v^2 s_\beta^2 + m_A^2 c_\beta^2 \end{pmatrix},$$

$$m_{LR}^2 = \frac{5}{8} g_\chi \tilde{g}_\chi v v_R \begin{pmatrix} c_\beta c_{\beta_R} & -c_\beta s_{\beta_R} \\ -s_\beta c_{\beta_R} & s_\beta s_{\beta_R} \end{pmatrix},$$

$$m_{RR}^2 = \begin{pmatrix} g_{Z_R}^2 v_R^2 c_{\beta_R}^2 + m_{A_R}^2 s_{\beta_R}^2 & -\frac{1}{2} (m_{A_R}^2 + g_{\Sigma_R}^2 v_R^2) s_{2\beta_R} \\ -\frac{1}{2} (m_{A_R}^2 + g_{\Sigma_R}^2 v_R^2) s_{2\beta_R} & g_{\Sigma_R}^2 v_R^2 s_{\beta_R}^2 + m_{A_R}^2 c_{\beta_R}^2 \end{pmatrix}$$

$$v_R^2 = v_{\chi_R}^2 + v_{\bar{\chi}_R}^2, \quad v^2 = v_d^2 + v_u^2, \quad s_x = \sin(x), \quad c_x = \cos(x)$$

$$g_\Sigma^2 = \frac{1}{4} (g_2^2 + g'^2 + \tilde{g}_\chi^2), \quad g_{\Sigma_R}^2 = \frac{25}{16} g_\chi^2, \quad \tilde{g}_\chi = g_\chi - g_{Y_\chi}$$

⇒ new D-term contributions at tree-level:  $m_{h^0, tree}^2 \leq m_Z^2 + \frac{1}{4} \tilde{g}_\chi^2 v^2$

H.E. Haber, M. Sher, PRD 35 (1987) 2206; M. Drees, PRD 35 (1987) 2910; M. Cvetič et al., PRD 56 (1997) 2861; E. Ma, arXiv:1108.4029; M. Hirsch et al., arXiv:1110.3037, arXiv:1206.3516

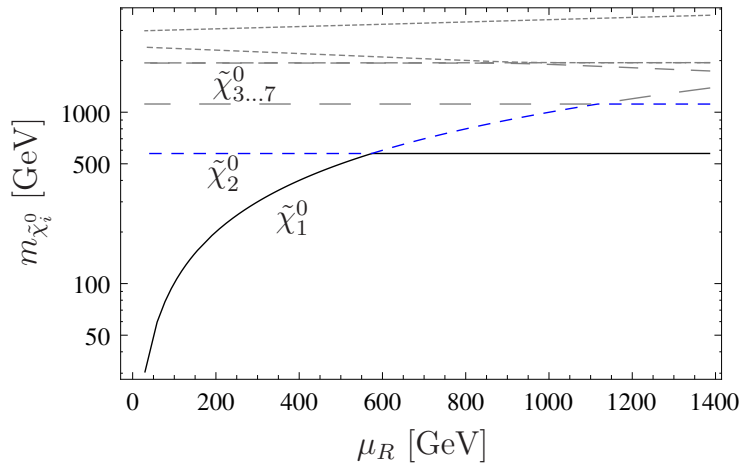
basis  $(\lambda_Y, \lambda_{W^3}, \tilde{h}_d^0, \tilde{h}_u^0, \lambda_\chi, \tilde{\chi}_R, \tilde{\chi}_R)$

$M_{\tilde{\chi}^0} =$

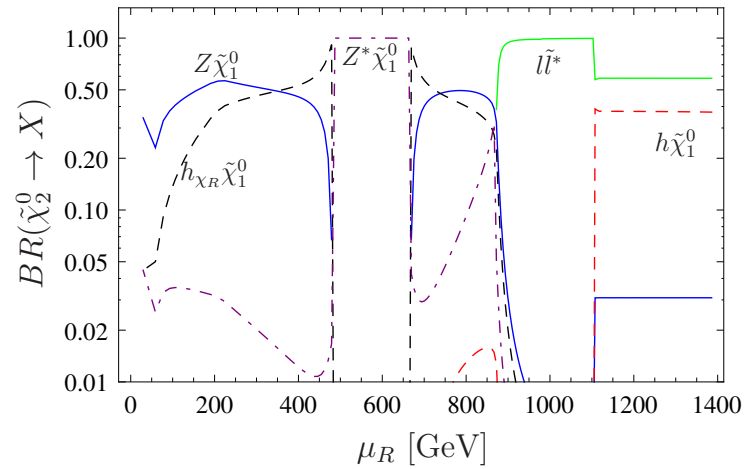
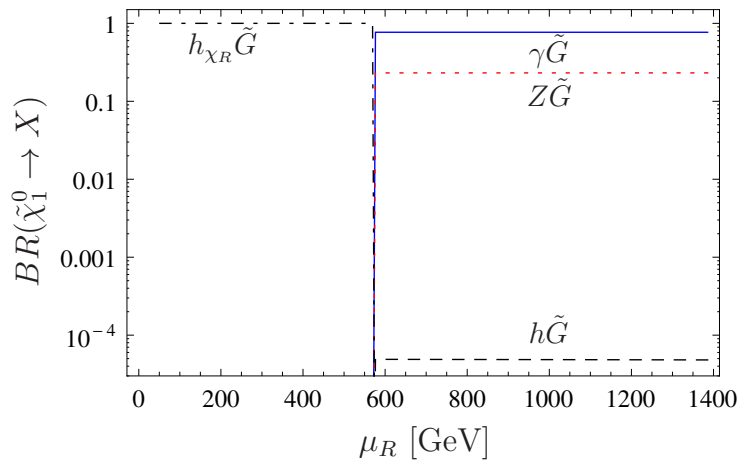
$$\begin{pmatrix} M_1 & 0 & -\frac{g'v_d}{2} & \frac{g'v_u}{2} & \frac{M_{Y\chi}}{2} & 0 & 0 \\ 0 & M_2 & \frac{g_2v_d}{2} & -\frac{g_2v_u}{2} & 0 & 0 & 0 \\ -\frac{g'v_d}{2} & \frac{g_2v_d}{2} & 0 & -\mu & \frac{(g_\chi - g_{Y\chi})v_d}{2} & 0 & 0 \\ \frac{g'v_u}{2} & -\frac{g_2v_u}{2} & -\mu & 0 & -\frac{(g_\chi - g_{Y\chi})v_u}{2} & 0 & 0 \\ \frac{M_{Y\chi}}{2} & 0 & \frac{(g_\chi - g_{Y\chi})v_d}{2} & -\frac{(g_\chi - g_{Y\chi})v_u}{2} & M_\chi & \frac{5g_\chi v_{\tilde{\chi}_R}}{4} & -\frac{5g_\chi v_{\chi_R}}{4} \\ 0 & 0 & 0 & 0 & \frac{5g_\chi v_{\tilde{\chi}_R}}{4} & 0 & -\mu_R \\ 0 & 0 & 0 & 0 & -\frac{5g_\chi v_{\chi_R}}{4} & -\mu_R & 0 \end{pmatrix}$$

neglecting the mixing between the two sectors and setting  $\tan \beta_R = 1$

$$m_i : \mu_R, \frac{1}{2} \left( M_\chi + \mu_R \pm \sqrt{\frac{1}{4}m_{Z'}^2 + (M_\chi - \mu_R)^2} \right)$$



M.E. Krauss, W.P., F. Staub, arXiv:1304.0769



$n = 1, \Lambda = 3.8 \cdot 10^5 \text{ GeV}, M = 9 \cdot 10^{11} \text{ GeV}, \tan \beta = 30, v_R = 6.7 \text{ GeV}, \tan \beta_R \text{ varied}$

$$M_\nu = \begin{pmatrix} 0 & \frac{1}{\sqrt{2}}v_u Y_\nu^T & 0 \\ \frac{1}{\sqrt{2}}v_u Y_\nu & 0 & \frac{1}{\sqrt{2}}v_{\chi_R} Y_s \\ 0 & \frac{1}{\sqrt{2}}v_{\chi_R} Y_s & \mu_S \end{pmatrix} \xrightarrow{1\text{gen}, \mu_S=0} m_\nu = \begin{pmatrix} 0 \\ -\sqrt{\frac{|Y_\nu|^2 v_u^2 + |Y_s|^2 v_{\chi_R}^2}{|Y_\nu|^2 v_u^2 + |Y_s|^2 v_{\chi_R}^2}} \\ \sqrt{\frac{|Y_\nu|^2 v_u^2 + |Y_s|^2 v_{\chi_R}^2}{|Y_\nu|^2 v_u^2 + |Y_s|^2 v_{\chi_R}^2}} \end{pmatrix}$$

setting  $\mu_S = 0$  and  $B_{\mu_S} = 0$

$$M_{\tilde{\nu}}^2 = \begin{pmatrix} m_L^2 + \frac{v_u^2}{2} Y_\nu^\dagger Y_\nu + D_L & \frac{1}{\sqrt{2}}v_u (T_\nu^\dagger - Y_\nu^\dagger \cot \beta\mu) & \frac{1}{2}v_u v_{\chi_R} Y_\nu^\dagger Y_s \\ \frac{1}{\sqrt{2}}v_u (T_\nu - Y_\nu \cot \beta\mu^*) & m_\nu^2 + \frac{v_u^2}{2} Y_\nu Y_\nu^\dagger + \frac{v_{\chi_R}^2}{2} Y_s^\dagger Y_s + D_R & \frac{1}{\sqrt{2}}v_{\chi_R} (T_s - Y_s \cot \beta_R \mu_R^*) \\ \frac{1}{2}v_u v_{\chi_R} Y_s^\dagger Y_\nu & \frac{1}{\sqrt{2}}v_{\chi_R} (T_s^\dagger - Y_s^\dagger \cot \beta_R \mu_R) & m_S^2 + \frac{v_{\chi_R}^2}{2} Y_s^\dagger Y_s \end{pmatrix}$$

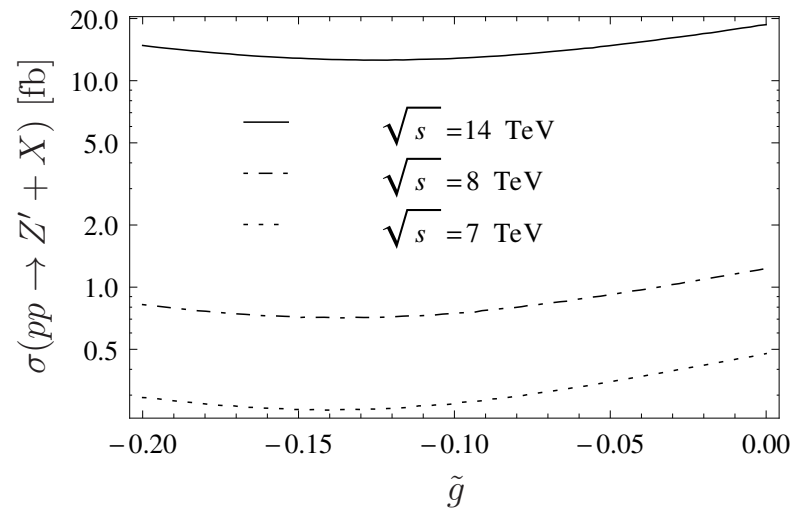
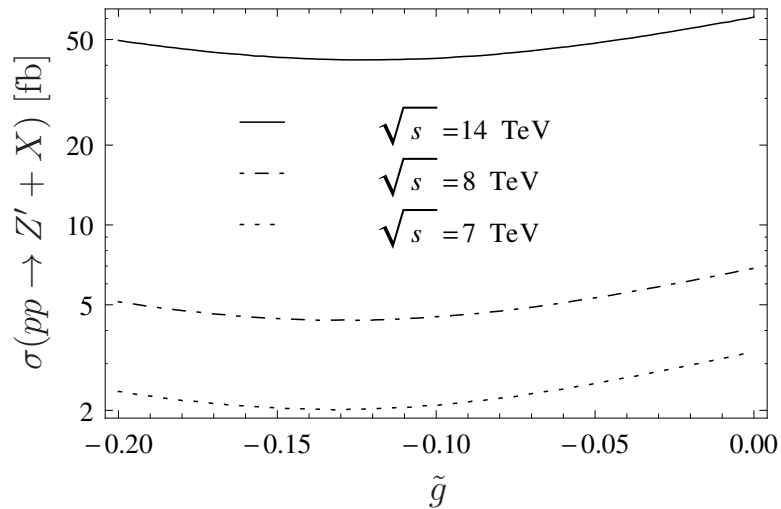
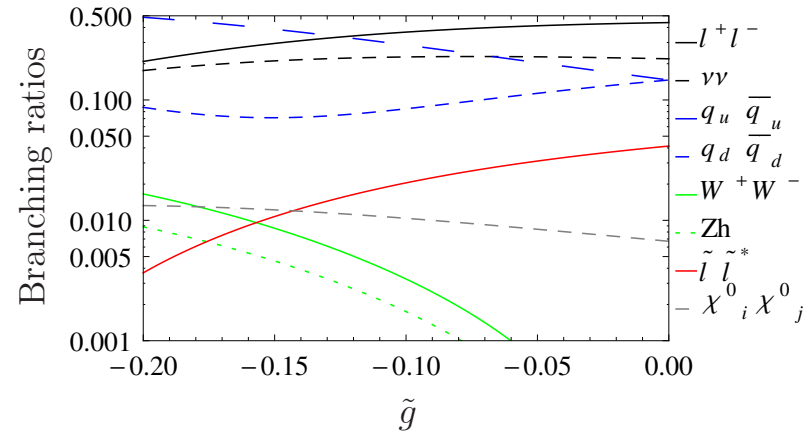
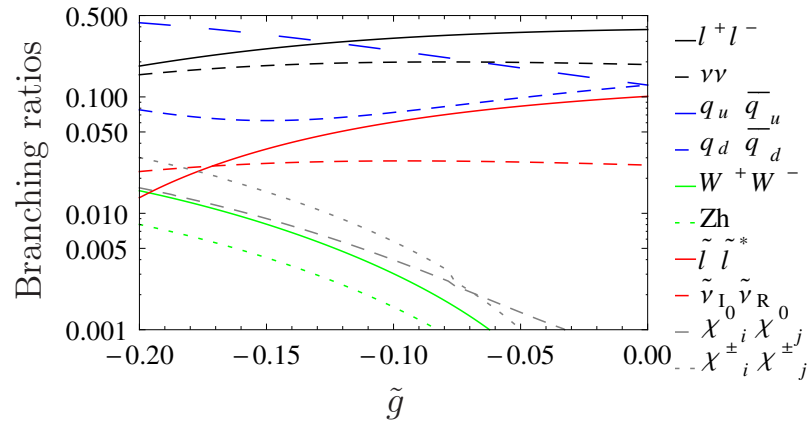
$$D_L = \frac{1}{32} \left( 2(-3g_\chi^2 + g_\chi g_{Y_\chi} + 2(g_2^2 + g'^2 + g_{Y_\chi}^2))v^2 c_{2\beta} - 5g_\chi(3g_\chi + 2g_{Y_\chi})v_R^2 c_{2\beta_R} \right) \mathbf{1}$$

$$D_R = \frac{5g_\chi}{32} \left( 2(g_\chi - g_{Y_\chi})v^2 c_{2\beta} + 5g_\chi v_R^2 c_{2\beta_R} \right) \mathbf{1}$$

$Z'$  couplings:  $Q_{B-L} \cdot g_{B-L} \rightarrow Q_{B-L} \cdot g_{B-L} + Q_Y \cdot \tilde{g}$

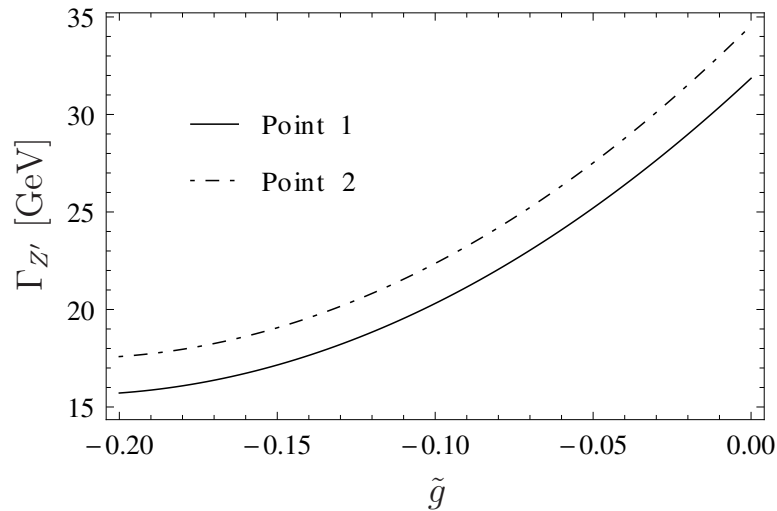
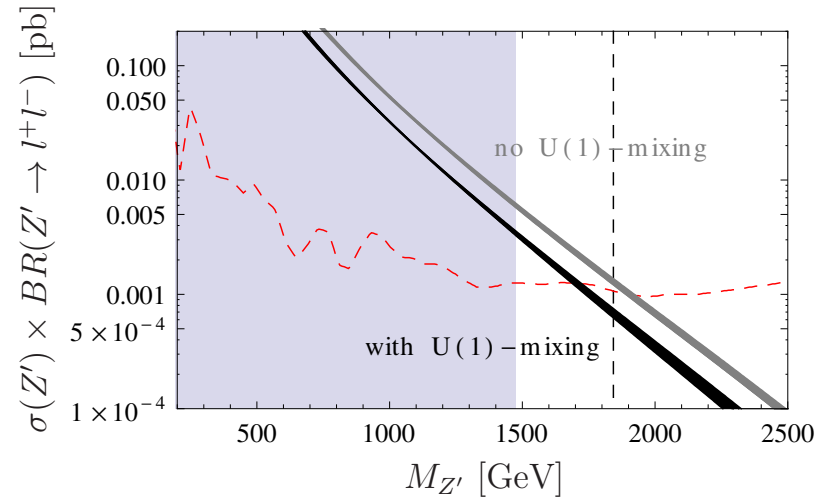
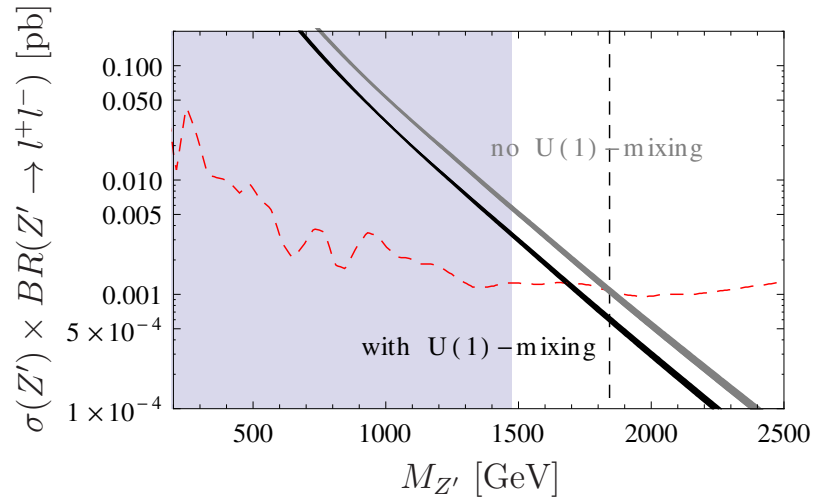
**BL1**

**BL2**



**BL1**

**BL2**



$Z'$  couplings:

$$Q_{B-L} \cdot g_{B-L} \rightarrow Q_{B-L} \cdot g_{B-L} + Q_Y \cdot \tilde{g}$$

No.	$\tilde{g} \neq 0$	$\tilde{g} = 0$
BL1	1680 GeV	1840 GeV
BL2	1700 GeV	1910 GeV

- invariant mass of the muon pair:  $M_{\mu\mu} > 200 \text{ GeV}$
- missing transverse momentum:  $p_T(\cancel{E}) > 200 \text{ GeV}$
- transverse cluster mass

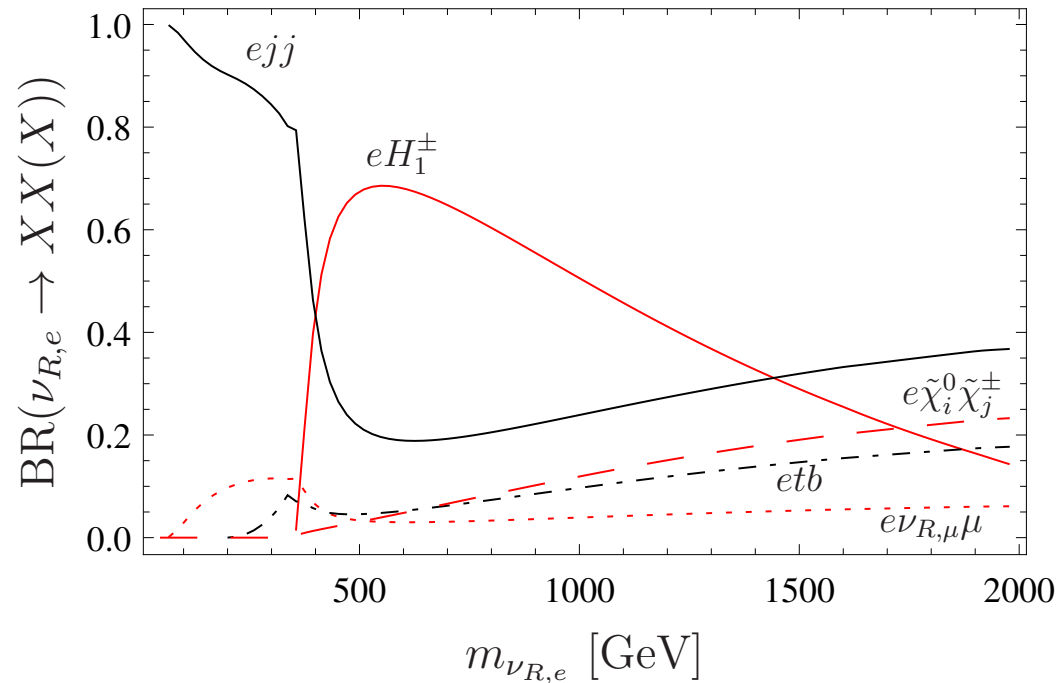
$$M_T = \sqrt{\left( \sqrt{p_T^2(\mu^+\mu^-) + M_{\mu\mu}^2} + p_T(\cancel{E}) \right)^2 - \left( \vec{p}_T(\mu^+\mu^-) + \vec{p}_T(\cancel{E}) \right)^2}$$

$$M_T > 800 \text{ GeV}$$

- for  $t\bar{t}$  suppression and squark/gluino cascade decays:

$$p_{T,\text{hardest jet}} < 40 \text{ GeV}$$





$m_{W'} = 2.2$  TeV,  $\tan \beta_R = 1.02$  and  $\mu_{\text{eff}} = 150$  GeV

M. Krauss, W.P., arXiv:1507.04349