

# A Tale of Tails

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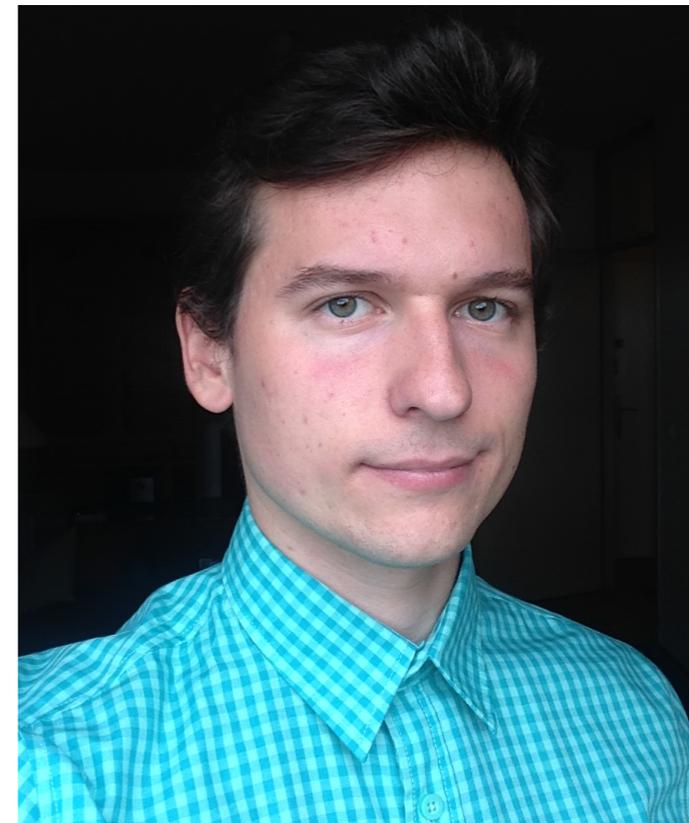


Gauge / Gravity Duality  
Würzburg  
2 August 2018

work (mainly) with



Benjamin Withers



Igor Novak



**SwissMAP**

The Mathematics of Physics  
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introduction

# Non equilibrium holography

## Holography:

quantum gravity = quantum field theory

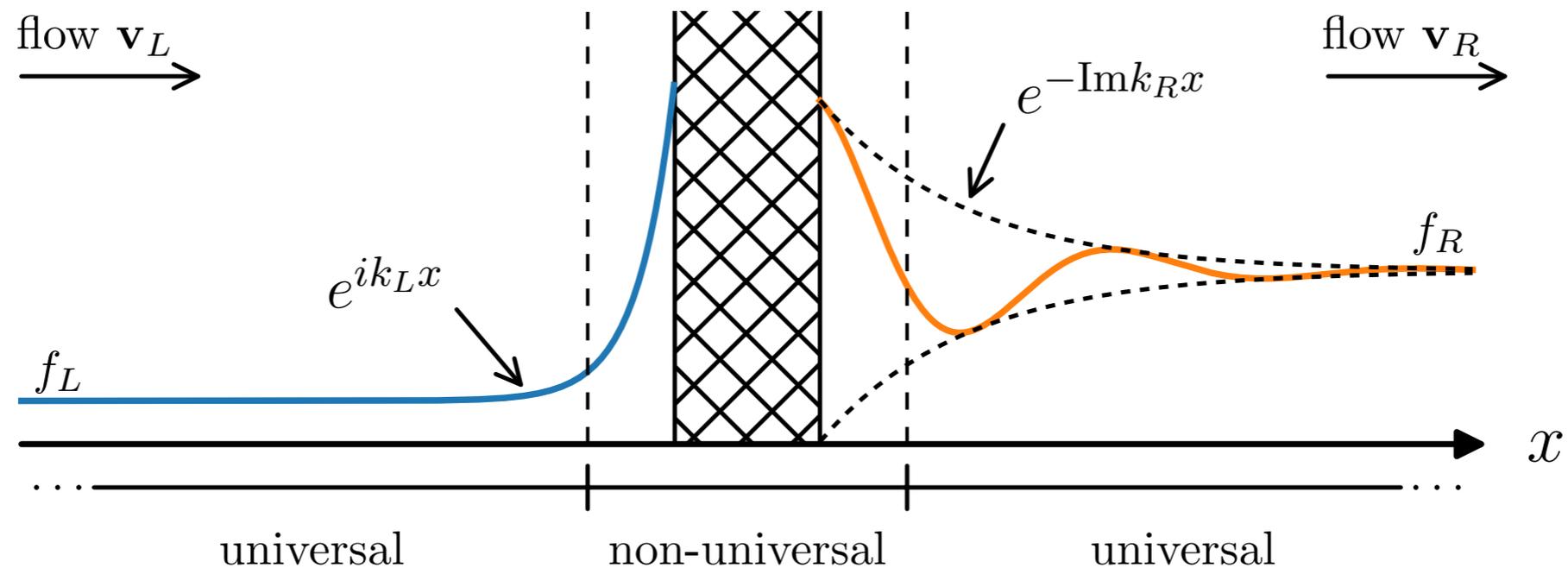
Thermalization/non-equilibrium  $\rightarrow$  BH formation (& evaporation)

**CFT:** late-time quantum physics of black holes

**AdS:** extract universal notions of non equilibrium physics from gravity

# The moral of this tale

Let us consider stationary flows of strongly coupled liquids



Universal features of spatial structure of NESS are determined cleanly by

$$\eta/s$$

Signatures in thermoelectric probes of strongly coupled liquids

# old tails of quenches

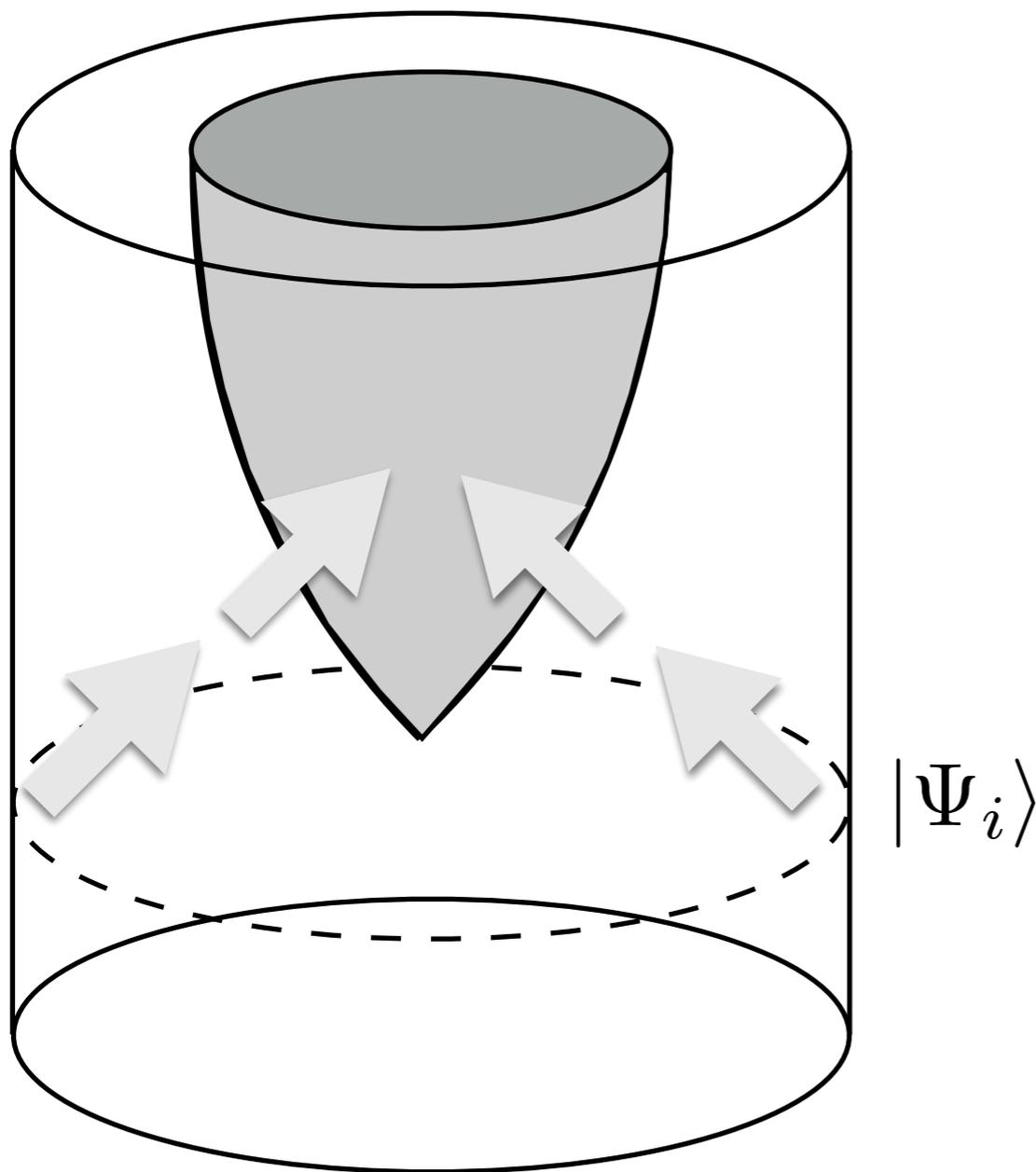
Nonequilibrium dynamics of holographic superfluid

[Bhaseen, Gauntlett, JS, Simons, Wiseman, PRL 110 (2013)]

# Quench dynamics in holographic CFT

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global  $\text{AdS}_{d+1}$



$|\Psi_i\rangle$

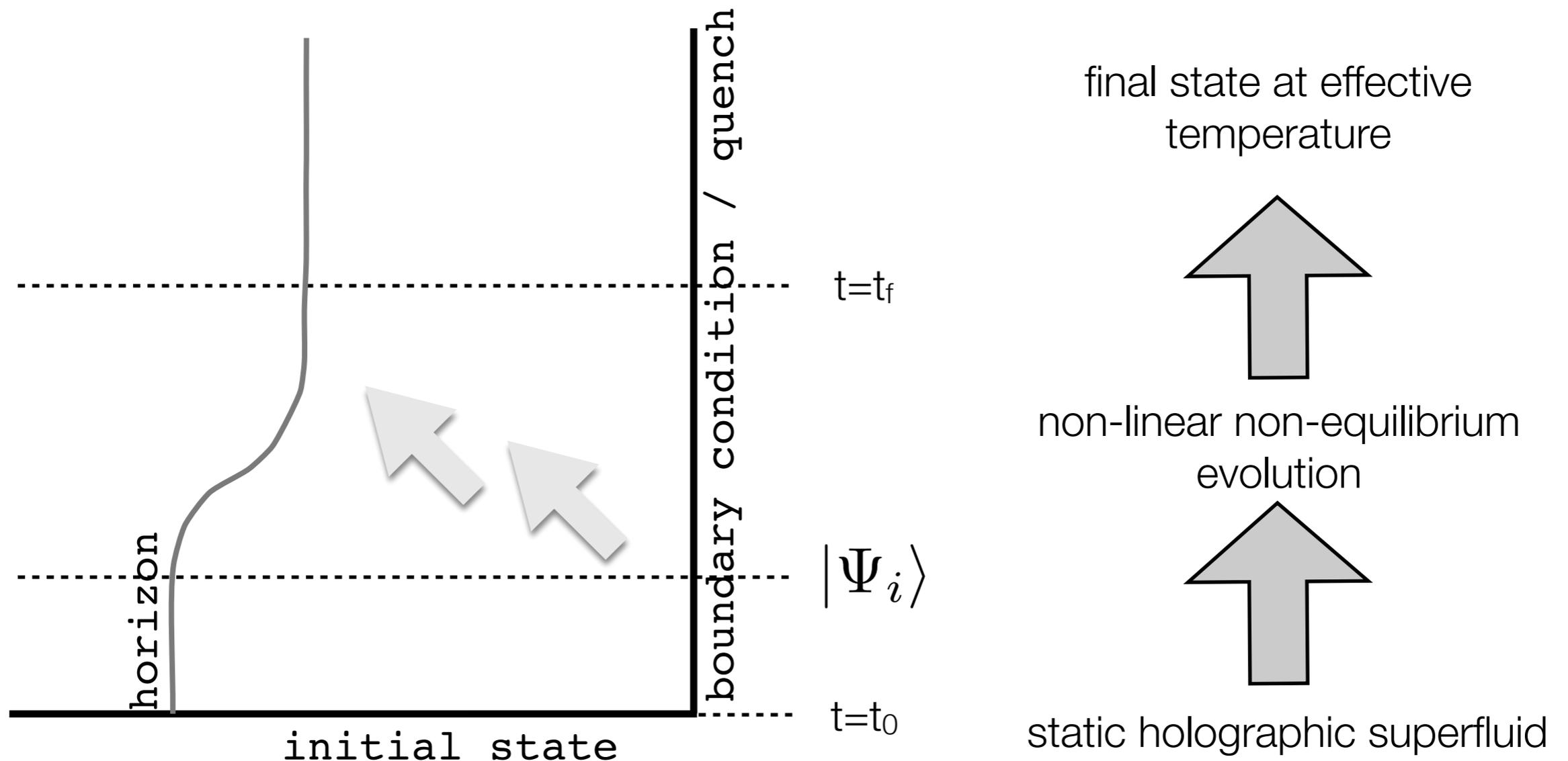
- Initial state: initial condition or result of operator deformation
- Quench is non-linear, but quickly gives way to linear QNM
- Can track n-pt functions, EE post quench
- Closed quantum system, but at late times can think of (large) BH as bath

# Example: Order Parameter Quench

Holographic superconductor quench in Poincaré AdS

conjugate to  
 $\langle \mathcal{O}(t, x) \rangle$

$$J_\psi(t) = \delta e^{-(t/\bar{\tau})^2}$$



# Universal Relaxation Dynamics

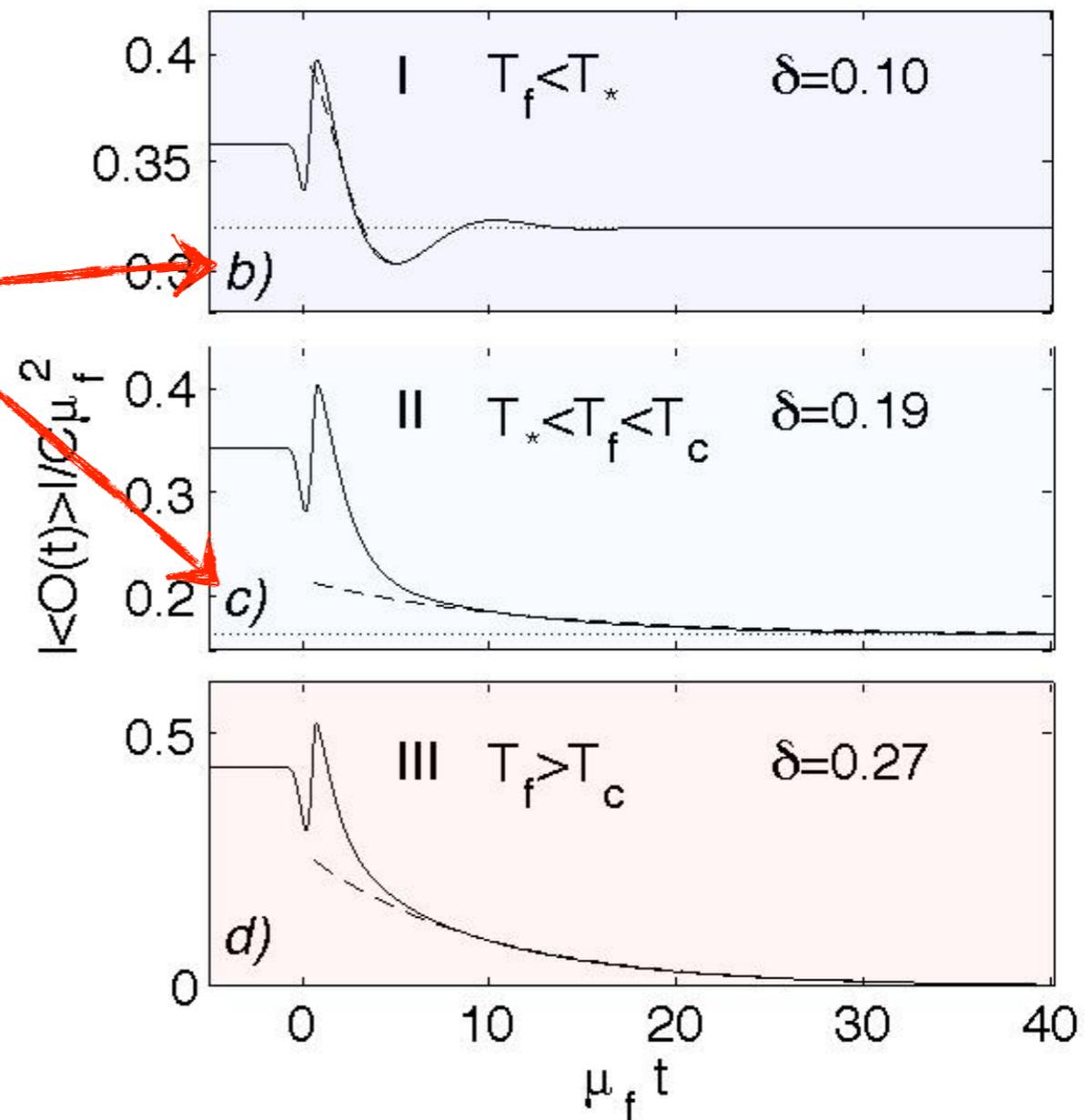
The dynamics of this quench give rise to three distinct regimes

I. Oscillation

**dynamical phase transition**

II. Decay to finite gap

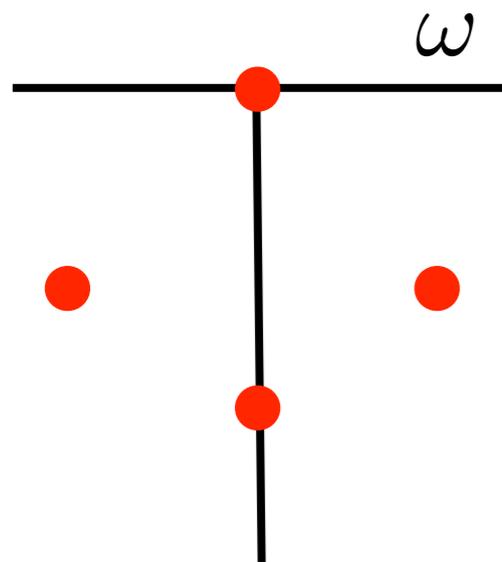
III. Decay to zero gap



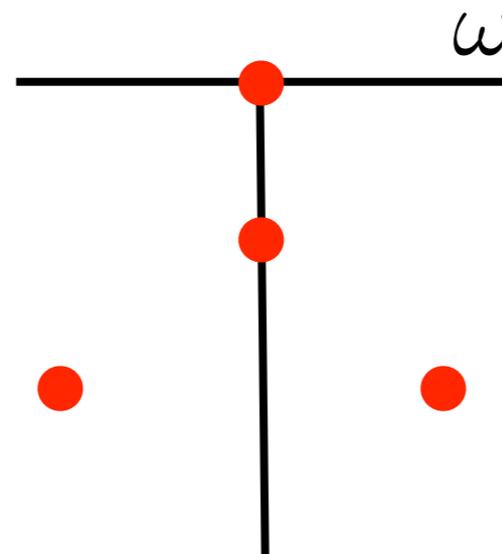
# Universal Relaxation Dynamics

Universal explanation in terms of QNM: poles in correlation functions at real momentum and complex frequency

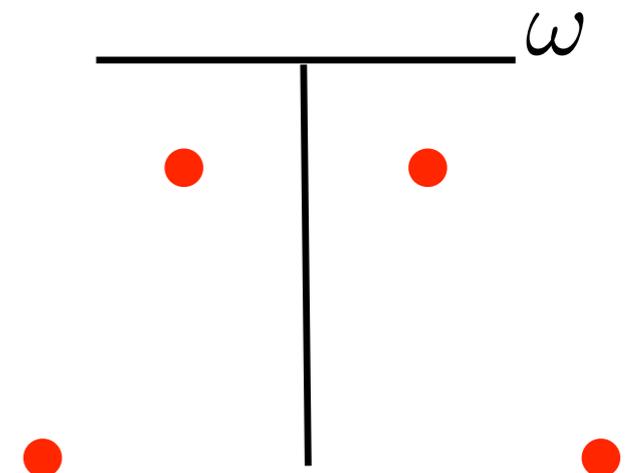
I. Oscillation



II. Pure Decay



III. Destruction of order (with pure decay)



strength of quench

Boundary between I & II is a dynamical phase transition. Also observed in e.g. BCS . [Barankov, Levitov, Spivak]

# tails of steady states

Universal structure of non equilibrium steady states

[JS, Benjamin Withers, PRL 119 (2017), Igor Novak, JS, Benjamin Withers (2018)]

# Nonequilibrium Steady States

NESS: out of equilibrium, but independent of time

**Example I:** current-driven steady states (apply E field)

- Non-linear conductivity [A. Karch, A. O'Bannon; A. Karch, S. Sondhi]
- out-of-equilibrium fluctuation theorem [JS, Andrew Green]

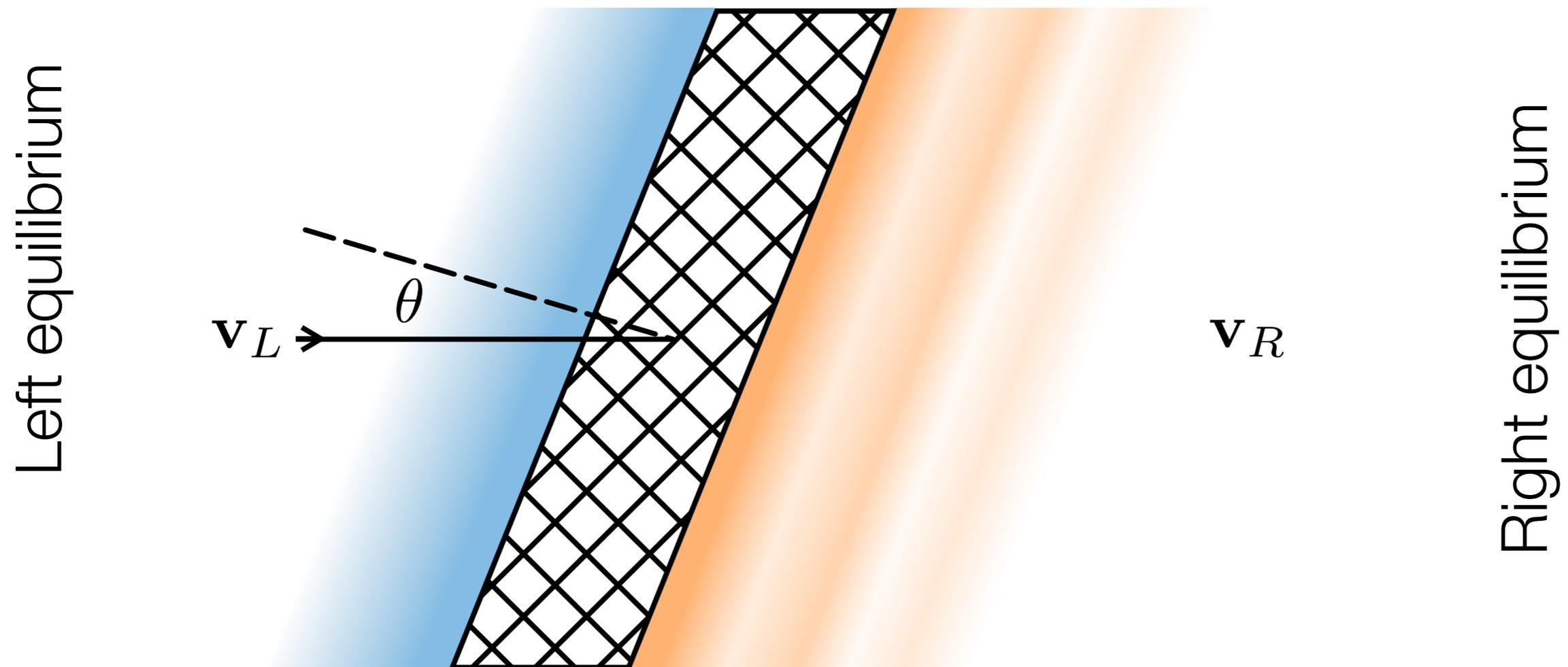
**Example II:** heat-driven steady states (local thermal quench)  
[J. Bhaseen, Doyon, A. Lucas, K. Schalm;...]

- Characterisation of steady state region
- full out-of-equilibrium fluctuation relations

# Fluid NESS

Consider a stationary flow of a strongly coupled liquid over an obstacle

$$J_g(x) = \delta e^{-(x/\bar{x})^2}$$



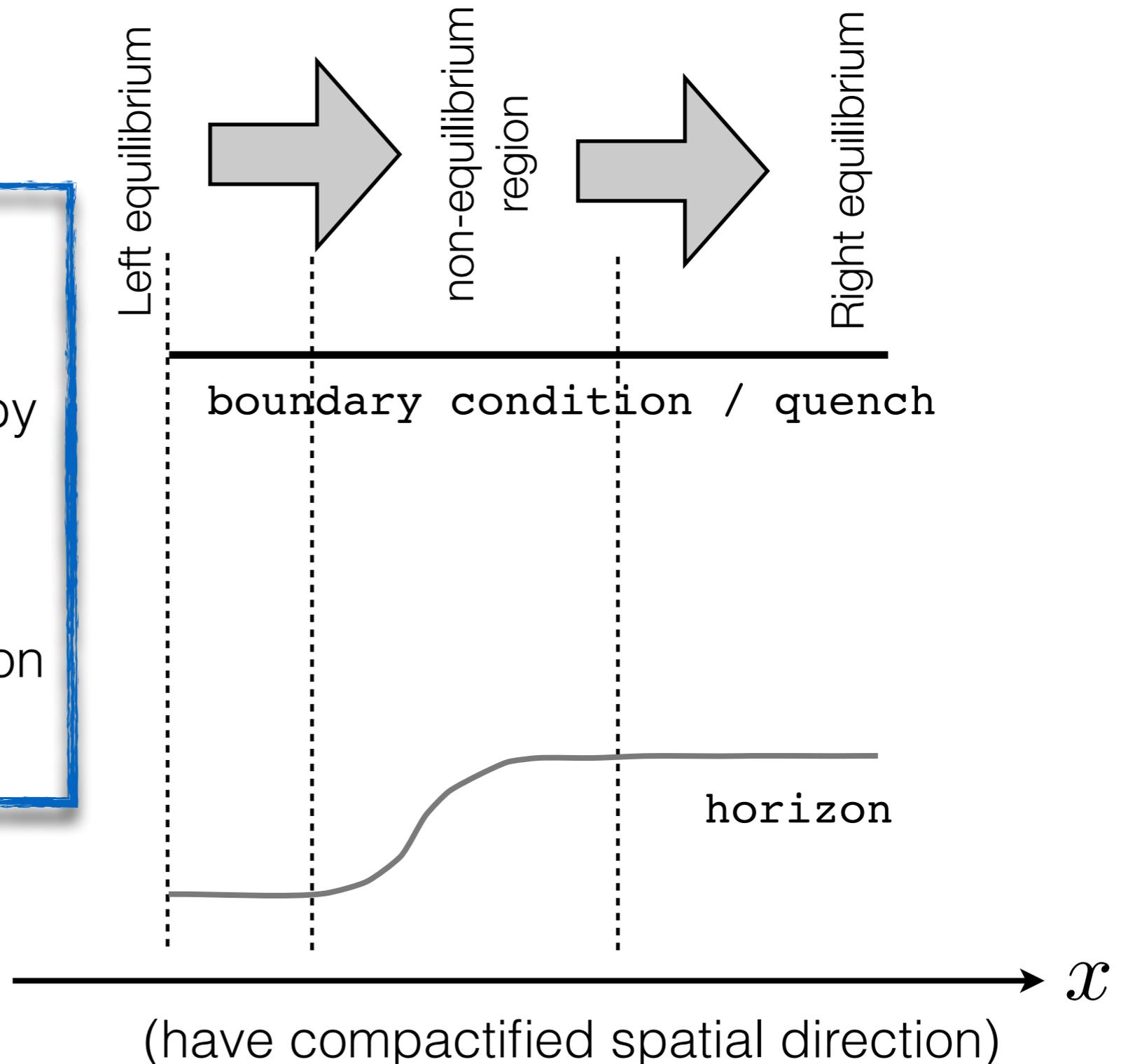
Compressible flow non-linearly disturbed by the obstacle

# Perspective: 'Spatial Quench' [Figueras & Wiseman (2013)]

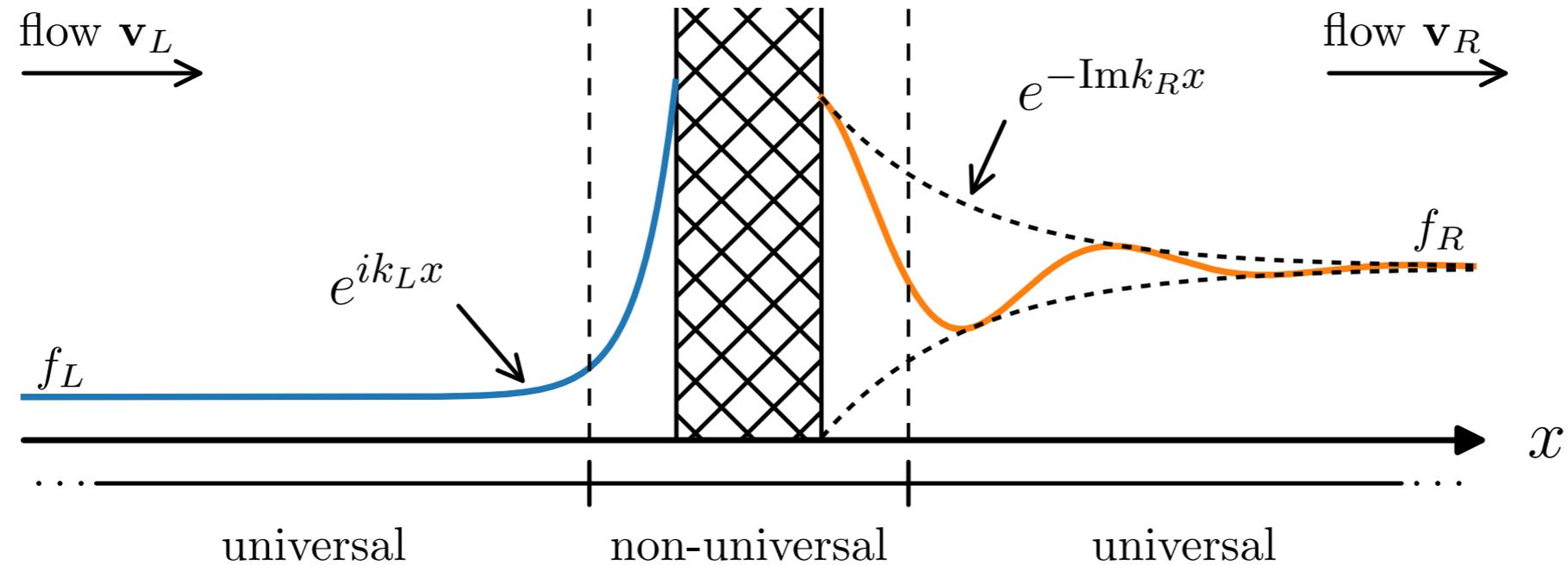
## Holographic dual:

Turn a 'standard' quench by 90 degrees

Duals are stationary black holes with nonKilling horizon



# Tails of NESS



How is asymptotic equilibrium reached (e.g. energy density)?

$$\epsilon(t, \mathbf{x}) = \epsilon_{L,R} + e^{i\mathbf{k} \cdot \mathbf{x}} \delta\epsilon$$

Seek time independent mode on background with finite velocity

$$\text{SCM:} \quad (\omega \in \mathbb{R}) \quad k \in \mathbb{C}$$

# Spatial Collective Modes (SCM)

Depending on sign of  $\text{Im } k$  can have diverging mode. Regularity:

$\text{Im } k > 0 \implies$  right mode

$\text{Im } k < 0 \implies$  left mode

Approach of expectation value of some operator at position  $x$ :

$$\delta \langle \Phi(t, x) \rangle = \int_{-\infty}^{\infty} F(x') \underbrace{(-i)\theta(x - x') \langle [\Phi(t, x), \Phi(t', x')] \rangle}_{\text{decay to right correlator}}$$

‘decay to right correlator’  $G^{[\searrow]}(x - x')$

This object has SCM poles in upper-half complex momentum plane

[see also: Amado, Hoyos, Landsteiner, Montero]

“The QNM of breaking spatial translation symmetry”

## Aside: relation to QNM

But aren't these just the usual quasinormal modes?

Let's look at BTZ for simplicity

$$\omega = \pm q - 4\pi iT \left( \frac{\Delta}{2} + n \right)$$

Boosting SCM into the fluid rest frame  $\rightarrow \omega = -\gamma kv$ ,  $q = \gamma k$

$$k = \frac{4\pi iT}{\gamma(v \pm 1)} \left( \frac{\Delta}{2} + n \right)$$

SCM is boosted QNM and then analytically continued to complex  $k$ .  
This also applies in higher dimensions

But there is **no** such argument in a nonrelativistic fluid!

# Example: hydro

Such modes can be constructed in hydro

$$\epsilon(t, \mathbf{x}) = \epsilon_{L,R} + e^{i\mathbf{k}\cdot\mathbf{x}} \delta\epsilon$$

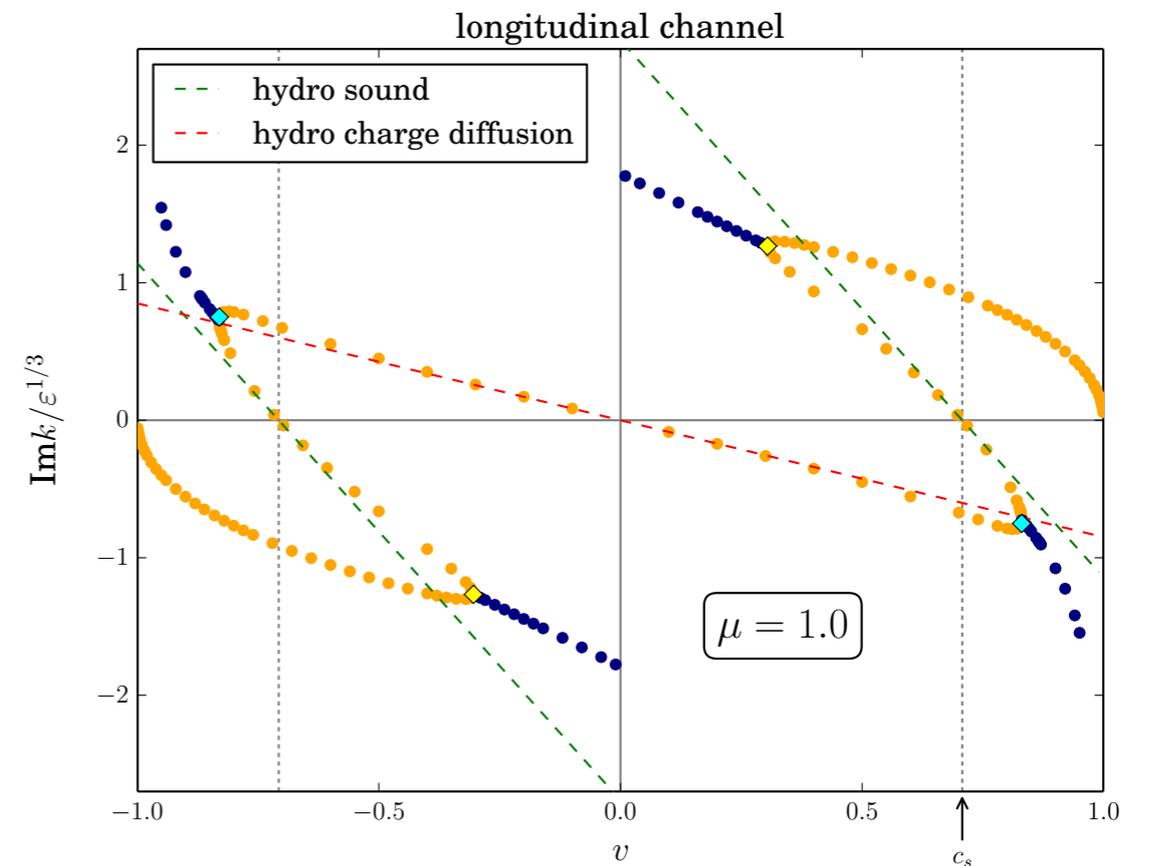
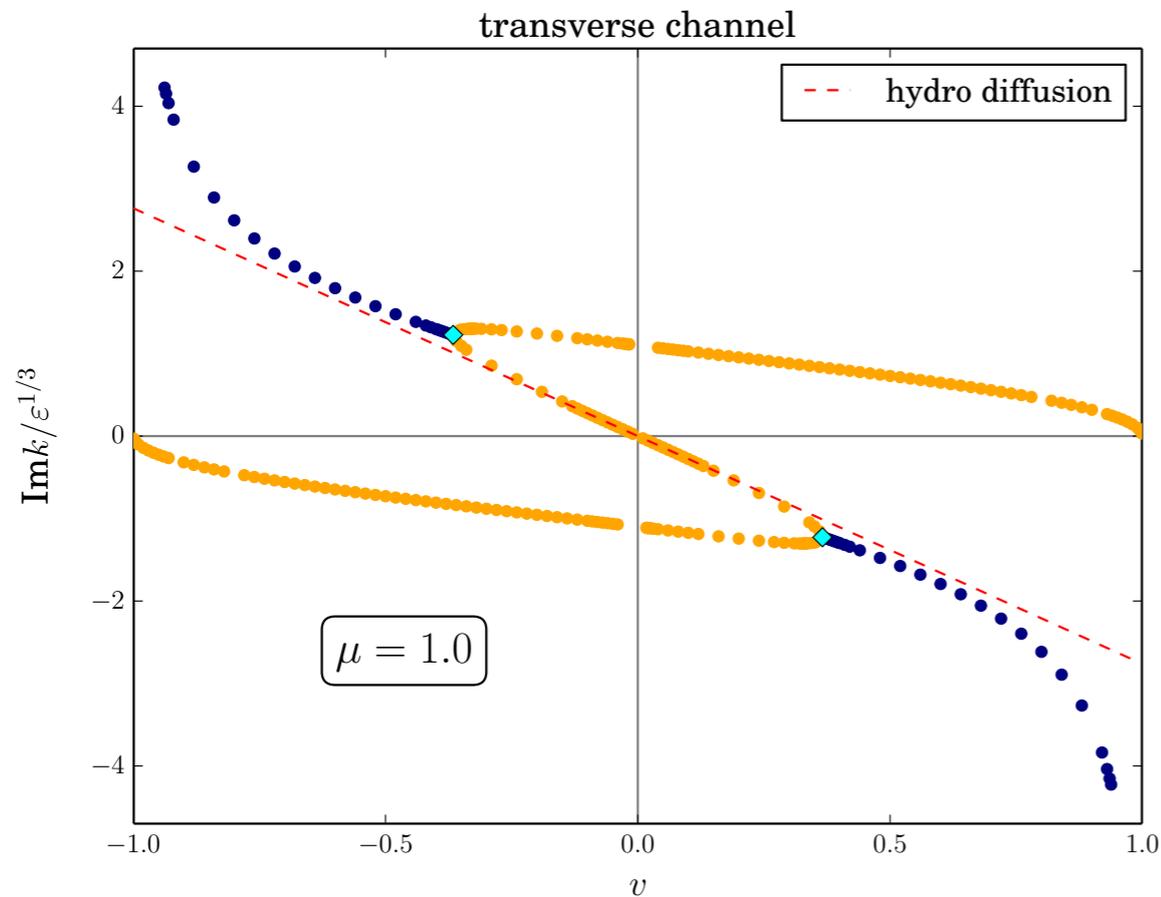
Seek time independent mode on background with finite velocity

<b>2 longitudinal</b>	$k = -iT \frac{s}{\frac{d-2}{d-1}\eta + \frac{1}{2}\zeta} \sqrt{1 - c_s^2} (v \pm c_s) + O(k)^2$	
<b>1 transverse</b>	$k = -iT \frac{s}{\eta} v + O(k)^2$	(angles suppressed)

(in charged case also have a mode analogous to charge diffusion)

# Beyond hydro

SCM of charged planar AdS<sub>4</sub> black hole

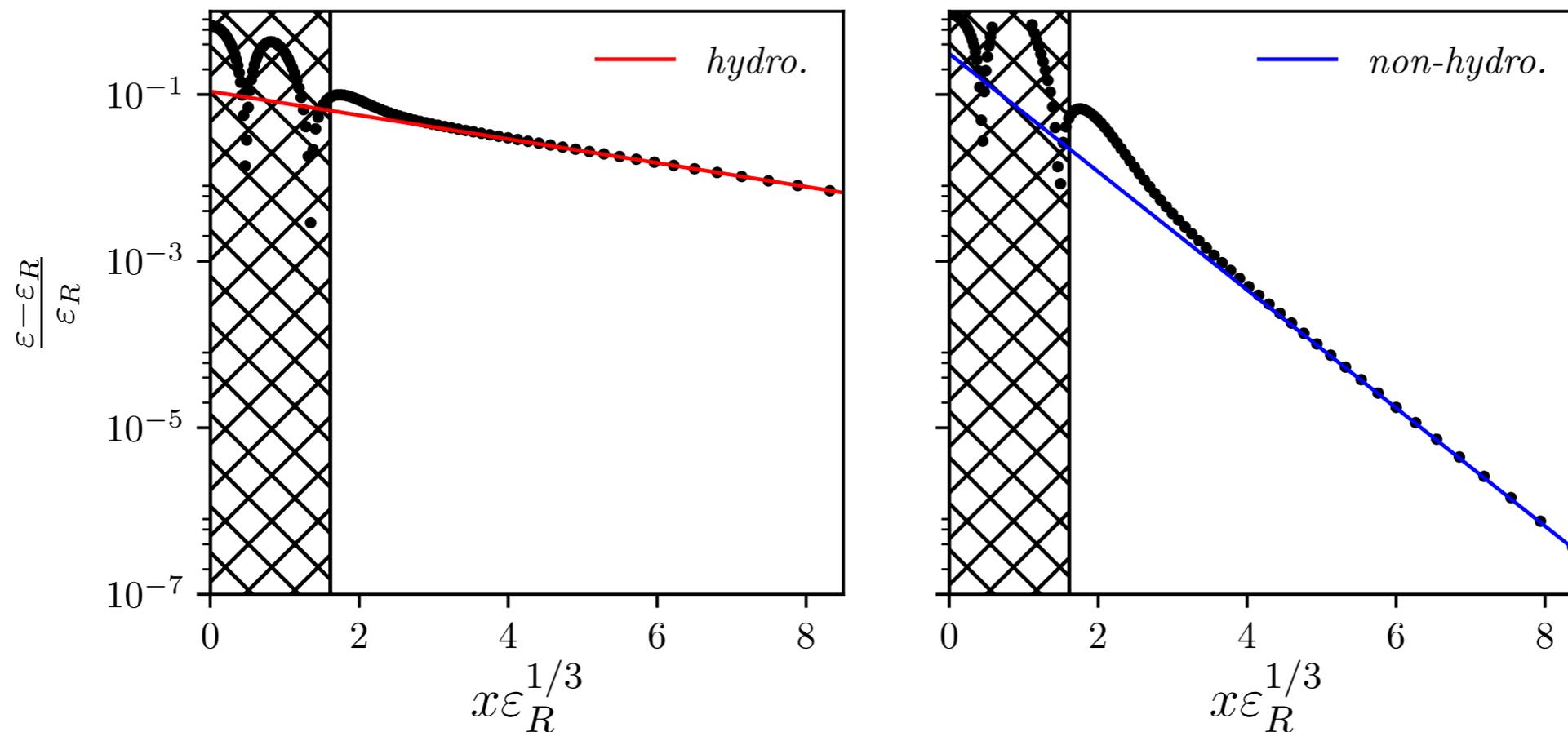


We get 'hydro SCM' and, additionally, more rapidly decaying non-hydro ones

Pole motion leads to interesting hydro-to-nonhydro phase transitions of NESS!

# Connection to non-linear NESS

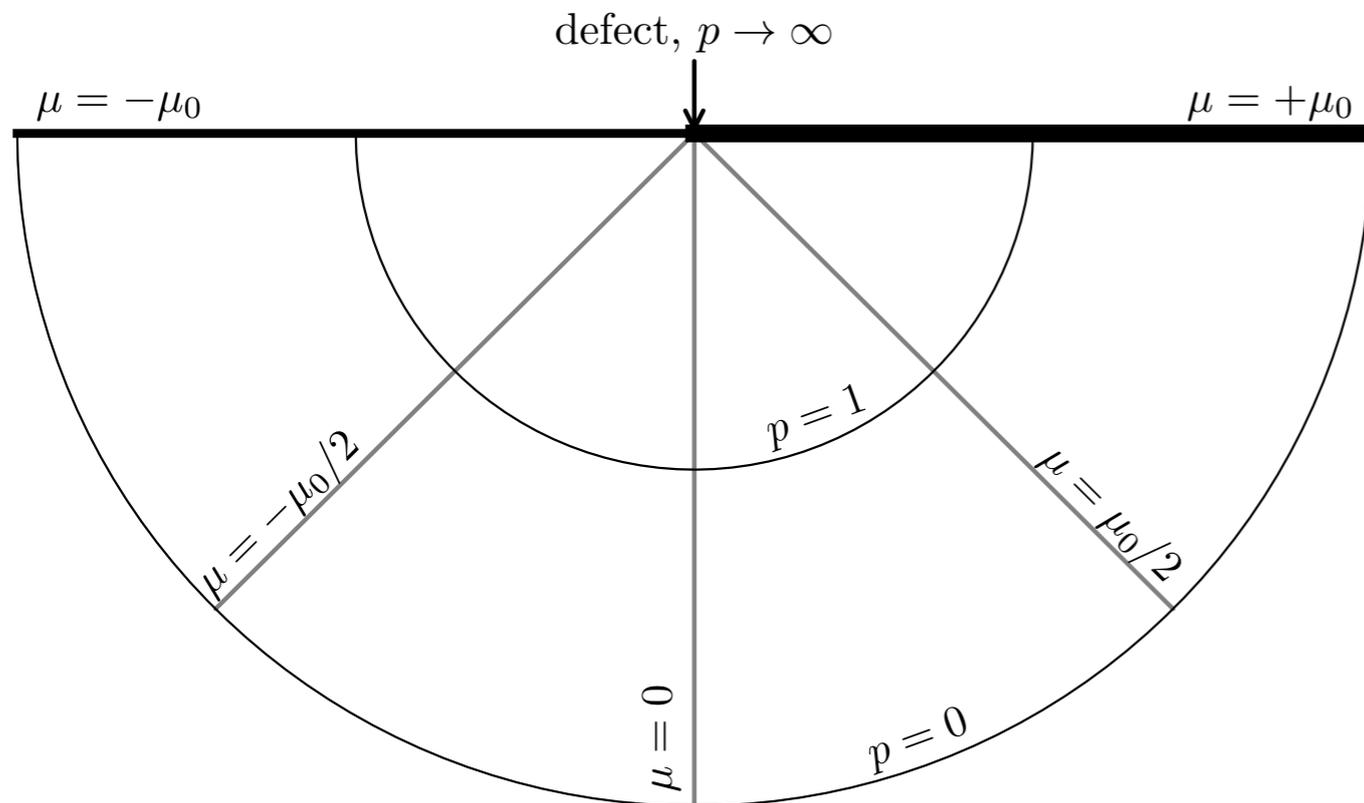
Can construct dual geometry to full non-linear NESS: nonKilling black holes in  $\text{AdS}_4$  using ‘generalized harmonic’ method (deTurck)



In all cases the strongly non-linear behaviour around the obstacle transitions rapidly to fall-offs predicted by SCM analysis

# A fully analytical example

In AdS<sub>3</sub> can construct black Janus, a finite-temperature defect solution  
 [Bak, Gutperle, Janik]



Contains scalar field with

$$\langle O_\phi \rangle = \mp 2 \frac{\sqrt{m}(2\pi T)^2}{1+m} \operatorname{csch}^2(2\pi T x)$$

Inverse

Laplace

$$k_n^\pm = \pm i 4\pi T (1 + n)$$

Matches precisely SCM spectrum of BTZ black hole

Fun corollary: can understand emergence of branch cut in  $T \rightarrow 0$  limit  
 which gives power law spatial decay

# Accessing $\eta/s$

Decay length of transverse mode depends cleanly on  $\eta/s$

$$k = -iT \frac{s}{\eta} \cos \theta + \mathcal{O}(k^2)$$

Setup NESS in strongly coupled material, use spatial structure to measure  $\eta/s$  (See also [Falkovich & Levitov, Crossno et al.] and poster on Poiseuille flow at this conference)

Decay lengths at in real life (graphene near charge neutrality):

Estimate using Geim group parameters:

$$|\text{Im}k|^{-1} = 0.7 \mu m$$

at room temperature for normal incidence

# Conclusions

A ubiquitous notion throughout holography: collective phenomena manifest as damped poles on complex  $\omega/k$  plane

**Time dependent case (QNM):** hydro and beyond + transitions between them

→ nonequilibrium phase transitions, critical phenomena

**Stationary case (SCM):** hydrostatic and beyond + transitions between them

→ excellent numerical and analytical evidence of relevance of SCM

→ nonequilibrium phase transitions, critical phenomena

→ stationary manifestation of shear viscosity

**Nonlinear:** both time dependent and NESS case under full control at large  $N$ , thanks to holography → new nonKilling horizons

# Outlook

**Broadly speaking:** a key task to my mind is to understand QNM/SCM beyond holography

microscopic approach?

resurgent gradient expansion [Heller & Spalinski, Withers]?

....

## **More focused:**

Large space of NESS to be explored

(spatial oscillations, relation to black funnels & droplets)

nonrelativistic case (SCM independently defined)

How far can we push this idea in graphene or other strongly correlated materials exhibiting hydro behavior? Relevant to experiment?