

## Einstein Double Field Equations

$$G_{AB} = 8\pi GT_{AB}$$

Hereafter  $A, B$  are  $O(D, D)$  indices

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# Prologue

- GR is based on Riemannian geometry, where the only geometric and gravitational field is the Riemannian metric,  $g_{\mu\nu}$ . Other fields are meant to be extra matter.
- On the other hand, string theory suggests to put a two-form gauge potential,  $B_{\mu\nu}$ , and a scalar dilaton,  $\phi$ , on an equal footing along with the metric:

- They form the closed string massless sector, being ubiquitous in all string theories,

$$\int d^D x \sqrt{-g} e^{-2\phi} \left( R_g + 4\partial_\mu \phi \partial^\mu \phi - \frac{1}{12} H_{\lambda\mu\nu} H^{\lambda\mu\nu} \right) \quad \text{where} \quad H = dB.$$

This action hides  $\mathbf{O}(D, D)$  symmetry of T-duality which transforms  $g, B, \phi$  into one another. Buscher 1987

- T-duality hints at a natural augmentation to General Relativity, in which the entire closed string massless sector constitutes the fundamental gravitational multiplet and the above action corresponds to ‘pure’ gravity.

Double Field Theory (DFT), initiated by Siegel 1993 & Hull, Zwiebach 2009-2010, turns out to provide a concrete realization for this idea of **Stringy Gravity** by manifesting  $\mathbf{O}(D, D)$  T-duality.

- My talk sketches the geometric construction of Stringy Gravity.  
In particular, I will introduce Einstein Double Field Equations,  $G_{AB} = 8\pi G T_{AB}$ , as the unifying single expression for all the equations of motion of the closed string massless sector, as well as Newton-Cartan, Carroll and Gomis-Ooguri gravities (non-Riemannian).

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# DFT as Stringy Gravity

## Notation for $\mathbf{O}(D, D)$ and $\mathbf{Spin}(1, D-1)_L \times \mathbf{Spin}(D-1, 1)_R$ Symmetries

Index	Representation	Metric (raising/lowering indices)
$A, B, \dots, M, N, \dots$	$\mathbf{O}(D, D)$ vector	$\mathcal{J}_{AB} = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}$
$p, q, \dots$	$\mathbf{Spin}(1, D-1)_L$ vector	$\eta_{pq} = \text{diag}(- + + \dots +)$
$\alpha, \beta, \dots$	$\mathbf{Spin}(1, D-1)_L$ spinor	$C_{\alpha\beta}, \quad (\gamma^p)^T = C\gamma^p C^{-1}$
$\bar{p}, \bar{q}, \dots$	$\mathbf{Spin}(D-1, 1)_R$ vector	$\bar{\eta}_{\bar{p}\bar{q}} = \text{diag}(+ - - \dots -)$
$\bar{\alpha}, \bar{\beta}, \dots$	$\mathbf{Spin}(D-1, 1)_R$ spinor	$\bar{C}_{\bar{\alpha}\bar{\beta}}, \quad (\bar{\gamma}^{\bar{p}})^T = \bar{C}\bar{\gamma}^{\bar{p}}\bar{C}^{-1}$

- Each symmetry rotates its own indices *exclusively*: spinors are  $\mathbf{O}(D, D)$  singlet.
- The constant  $\mathbf{O}(D, D)$  metric,  $\mathcal{J}_{AB}$ , decomposes the doubled coordinates into two parts,

$$x^A = (\tilde{x}_\mu, x^\nu), \quad \partial_A = (\tilde{\partial}^\mu, \partial_\nu),$$

where  $\mu, \nu$  are  $D$ -dimensional curved indices.

- The twofold local Lorentz symmetries indicate two distinct locally inertial frames for the left-moving and the right-moving closed string sectors separately  $\Rightarrow$  **Unification of IIA and IIB.**

- **Closed string massless sector as ‘Stringy Graviton Fields’**

The stringy graviton fields consist of the DFT dilaton,  $d$ , and DFT metric,  $\mathcal{H}_{MN}$  :

$$\mathcal{H}_{MN} = \mathcal{H}_{NM}, \quad \mathcal{H}_K{}^L \mathcal{H}_M{}^N \mathcal{J}_{LN} = \mathcal{J}_{KM}.$$

Combining  $\mathcal{J}_{MN}$  and  $\mathcal{H}_{MN}$ , we get a pair of symmetric projection matrices,

$$\begin{aligned} P_{MN} &= P_{NM} = \frac{1}{2}(\mathcal{J}_{MN} + \mathcal{H}_{MN}), & P_L{}^M P_M{}^N &= P_L{}^N, \\ \bar{P}_{MN} &= \bar{P}_{NM} = \frac{1}{2}(\mathcal{J}_{MN} - \mathcal{H}_{MN}), & \bar{P}_L{}^M \bar{P}_M{}^N &= \bar{P}_L{}^N, \end{aligned}$$

which are orthogonal and complete,

$$P_L{}^M \bar{P}_M{}^N = 0, \quad P_M{}^N + \bar{P}_M{}^N = \delta_M{}^N.$$

Further, taking the “square roots” of the projectors,

$$P_{MN} = V_M{}^P V_N{}^q \eta_{pq}, \quad \bar{P}_{MN} = \bar{V}_M{}^{\bar{P}} \bar{V}_N{}^{\bar{Q}} \bar{\eta}_{\bar{P}\bar{Q}},$$

we get a pair of DFT vielbeins:

$$V_{Mp} V^M{}_q = \eta_{pq}, \quad \bar{V}_{M\bar{P}} \bar{V}^M{}_{\bar{Q}} = \bar{\eta}_{\bar{P}\bar{Q}}, \quad V_{Mp} \bar{V}^M{}_{\bar{Q}} = 0, \quad V_M{}^P V_{Np} + \bar{V}_M{}^{\bar{P}} \bar{V}_{N\bar{P}} = \mathcal{J}_{MN}.$$



## Classification of DFT backgrounds, 1707.03713 with Kevin Morand

The most general form of the DFT metric,  $\mathcal{H}_{MN} = \mathcal{H}_{NM}$ ,  $\mathcal{H}_K{}^L \mathcal{H}_M{}^N \mathcal{J}_{LN} = \mathcal{J}_{KM}$ , is characterized by two non-negative integers,  $(n, \bar{n})$ ,  $0 \leq n + \bar{n} \leq D$ :

$$\mathcal{H}_{AB} = \begin{pmatrix} H^{\mu\nu} & -H^{\mu\sigma} B_{\sigma\lambda} + Y_i^\mu X_\lambda^i - \bar{Y}_{\bar{i}}^\mu \bar{X}_{\bar{\lambda}}^{\bar{i}} \\ B_{\kappa\rho} H^{\rho\nu} + X_\kappa^i Y_i^\nu - \bar{X}_{\bar{\kappa}}^{\bar{i}} \bar{Y}_{\bar{i}}^\nu & K_{\kappa\lambda} - B_{\kappa\rho} H^{\rho\sigma} B_{\sigma\lambda} + 2X_{(\kappa}^i B_{\lambda)\rho} Y_i^\rho - 2\bar{X}_{(\bar{\kappa}}^{\bar{i}} B_{\bar{\lambda})\rho} \bar{Y}_{\bar{i}}^\rho \end{pmatrix}$$

i) Symmetric and skew-symmetric fields:  $H^{\mu\nu} = H^{\nu\mu}$ ,  $K_{\mu\nu} = K_{\nu\mu}$ ,  $B_{\mu\nu} = -B_{\nu\mu}$ ;

ii) Two kinds of eigenvectors having zero eigenvalue, with  $i, j = 1, 2, \dots, n$  &  $\bar{i}, \bar{j} = 1, 2, \dots, \bar{n}$ ,

$$H^{\mu\nu} X_\nu^i = 0, \quad H^{\mu\nu} \bar{X}_{\bar{\nu}}^{\bar{i}} = 0, \quad K_{\mu\nu} Y_j^\nu = 0, \quad K_{\mu\nu} \bar{Y}_{\bar{j}}^\nu = 0;$$

iii) Completeness relation:  $H^{\mu\rho} K_{\rho\nu} + Y_i^\mu X_\nu^i + \bar{Y}_{\bar{i}}^\mu \bar{X}_{\bar{\nu}}^{\bar{i}} = \delta^\mu{}_\nu$ .

• Orthonormality follows

$$Y_i^\mu X_\mu^j = \delta_i^j, \quad \bar{Y}_{\bar{i}}^\mu \bar{X}_{\bar{\mu}}^{\bar{j}} = \delta_{\bar{i}}^{\bar{j}}, \quad Y_i^\mu \bar{X}_{\bar{\mu}}^{\bar{j}} = \bar{Y}_{\bar{i}}^\mu X_\mu^j = 0.$$

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- It is instructive to note the  $\mathbf{O}(D, D)$  invariant trace,  $\mathcal{H}_A{}^A = 2(n - \bar{n})$  and

$$\mathcal{H}_{AB} = \begin{pmatrix} 1 & 0 \\ B & 1 \end{pmatrix} \begin{pmatrix} H & Y_i(X^i)^T - \bar{Y}_{\bar{i}}(\bar{X}^{\bar{i}})^T \\ X^i(Y_i)^T - \bar{X}^{\bar{i}}(\bar{Y}_{\bar{i}})^T & K \end{pmatrix} \begin{pmatrix} 1 & -B \\ 0 & 1 \end{pmatrix}.$$

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I.  $(n, \bar{n}) = (0, 0)$  corresponds to the Riemannian case or “Generalized Geometry”:

$$\mathcal{H}_{MN} \equiv \begin{pmatrix} g^{-1} & -g^{-1}B \\ Bg^{-1} & g - Bg^{-1}B \end{pmatrix}, \quad e^{-2d} \equiv \sqrt{|g|}e^{-2\phi} \quad \text{Giveon, Rabinovici, Veneziano '89, Duff '90}$$

II. Generically, string becomes chiral and anti-chiral over the  $n$  and  $\bar{n}$  dimensions:

$$X_{\mu}^i \partial_+ x^{\mu}(\tau, \sigma) \equiv 0, \quad \bar{X}_{\mu}^{\bar{i}} \partial_- x^{\mu}(\tau, \sigma) \equiv 0.$$

– Such non-Riemannian examples include

- $(1, 0)$  Newton-Cartan gravity **c.f. Obers' talk** ( $ds^2 = -c^2 dt^2 + dx^2$ ,  $\lim_{c \rightarrow \infty} g^{-1}$  is finite & degenerate)
- $(1, 1)$  Gomis-Ooguri non-relativistic string **Charles Melby-Thompson, Rene Meyer, Ko, JHP 2015**
- $(D-1, 0)$  ultra-relativistic Carroll gravity

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- **Diffeomorphisms** in Stringy Gravity are given by “generalized Lie derivative”: Siegel 1993

$$\hat{\mathcal{L}}_{\xi} T_{A_1 \dots A_n} := \xi^B \partial_B T_{A_1 \dots A_n} + \omega_T \partial_B \xi^B T_{A_1 \dots A_n} + \sum_{i=1}^n (\partial_{A_i} \xi_B - \partial_B \xi_{A_i}) T_{A_1 \dots A_{i-1} \quad B \quad A_{i+1} \dots A_n},$$

where  $\omega_T$  is the weight, e.g.  $\delta e^{-2d} = \partial_B (\xi^B e^{-2d})$ ,  $\delta V_{Ap} = \xi^B \partial_B V_{Ap} + (\partial_A \xi_B - \partial_B \xi_A) V^B_p$ .

- For consistency, the so-called ‘section condition’ should be imposed:  $\partial_M \partial^M = 0$ .

From  $\partial_M \partial^M = 2\partial_{\mu} \tilde{\delta}^{\mu}$ , the section condition can be easily solved by letting  $\tilde{\delta}^{\mu} = 0$ .

The general solutions are then generated by the  $\mathbf{O}(D, D)$  rotation of it.

- The section condition is mathematically equivalent to a certain translational invariance:

$$\Phi_i(x) = \Phi_i(x + \Delta), \quad \Delta^M = \Phi_j \partial^M \Phi_k,$$

where  $\Phi_i, \Phi_j, \Phi_k \in \{d, \mathcal{H}_{MN}, \xi^M, \partial_N d, \partial_L \mathcal{H}_{MN}, \dots\}$ , arbitrary functions appearing in DFT, and  $\Delta^M$  is said to be derivative-index-valued.

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$$\hat{\mathcal{L}}_\xi T_{A_1 \dots A_n} := \xi^B \partial_B T_{A_1 \dots A_n} + \omega_T \partial_B \xi^B T_{A_1 \dots A_n} + \sum_{i=1}^n (\partial_{A_i} \xi_B - \partial_B \xi_{A_i}) T_{A_1 \dots A_{i-1} \quad A_{i+1} \dots A_n},$$

where  $\omega_T$  is the weight, e.g.  $\delta e^{-2d} = \partial_B (\xi^B e^{-2d})$ ,  $\delta V_{Ap} = \xi^B \partial_B V_{Ap} + (\partial_A \xi_B - \partial_B \xi_A) V^B_p$ .

- For consistency, the so-called ‘section condition’ should be imposed:  $\partial_M \partial^M = 0$ .

From  $\partial_M \partial^M = 2\partial_\mu \tilde{\partial}^\mu$ , the section condition can be easily solved by letting  $\tilde{\partial}^\mu = 0$ .

The general solutions are then generated by the  $\mathbf{O}(D, D)$  rotation of it.

- The section condition is mathematically equivalent to a certain translational invariance:

$$\Phi_i(x) = \Phi_i(x + \Delta), \quad \Delta^M = \Phi_j \partial^M \Phi_k,$$

where  $\Phi_i, \Phi_j, \Phi_k \in \{d, \mathcal{H}_{MN}, \xi^M, \partial_N d, \partial_L \mathcal{H}_{MN}, \dots\}$ , arbitrary functions appearing in DFT,  
and  $\Delta^M$  is said to be derivative-index-valued.

- ▶ ‘Physics’ should be invariant under such shifts of the doubled coordinates in Stringy Gravity.

Doubled coordinates,  $x^M = (\tilde{x}_\mu, x^\nu)$ , are gauged through an equivalence relation,

$$x^M \sim x^M + \Delta^M(x),$$

where  $\Delta^M$  is derivative-index-valued.



Each equivalence class, or gauge orbit in  $\mathbb{R}^{D+D}$ , corresponds to a single physical point in  $\mathbb{R}^D$ .

- If we solve the section condition by letting  $\tilde{\partial}^\mu \equiv 0$ , and further choose  $\Delta^M = c_\mu \partial^M x^\mu$ , we note

$$(\tilde{x}_\mu, x^\nu) \sim (\tilde{x}_\mu + c_\mu, x^\nu) : \tilde{x}_\mu \text{'s are gauged and } x^\nu \text{'s form a section.}$$

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- Further, if we 'gauge'  $dx^M$  explicitly by introducing a derivative-index-valued gauge potential,

$$dx^M \longrightarrow Dx^M = dx^M - \mathcal{A}^M, \quad \mathcal{A}^M \partial_M = 0,$$

it is possible to define  $\mathbf{O}(D, D)$  & diffeomorphism covariant 'proper length' through a path integral,

$$\text{Proper Length} := -\ln \left[ \int \mathcal{D}\mathcal{A} \exp \left( - \int \sqrt{Dx^M Dx^N \mathcal{H}_{MN}} \right) \right].$$

Under  $\delta x^M = \xi^M$ ,  $\delta(dx^M) = dx^N \partial_N \xi^M \neq dx^N (\partial_N \xi^M - \partial^M \xi_N)$  versus  $\delta(Dx^M) = Dx^N (\partial_N \xi^M - \partial^M \xi_N)$ .

- For the  $(0, 0)$  Riemannian DFT-metric, with  $\tilde{\partial}^\mu = 0$ ,  $\mathcal{A}^M = A_\mu \partial^M x^\mu = (A_\mu, 0)$ , from

$$Dx^M Dx^N \mathcal{H}_{MN} \equiv dx^\mu dx^\nu g_{\mu\nu} + (d\tilde{x}_\mu - A_\mu + dx^\rho B_{\rho\mu}) (d\tilde{x}_\nu - A_\nu + dx^\sigma B_{\sigma\nu}) g^{\mu\nu}$$

after integrating out  $A_\mu$ , the proper length reduces to the conventional one,

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# Doubled-yet-gauged sigma models

The definition of the proper length readily leads to ‘covariant’ actions:

## I. Particle action

Ko-JHP-Suh 2016

$$S_{\text{particle}} = \int d\tau e^{-1} D_\tau x^M D_\tau x^N \mathcal{H}_{MN}(x) - \frac{1}{4} m^2 e$$

## II. String action

Lee-JHP 2013, c.f. Hull 2006

$$S_{\text{string}} = \frac{1}{4\pi\alpha'} \int d^2\sigma - \frac{1}{2} \sqrt{-hh^{ij}} D_i x^M D_j x^N \mathcal{H}_{MN}(x) - \epsilon^{ij} D_i x^M A_{jM}$$

With the (0, 0) Riemannian DFT-metric plugged, after integrating out the auxiliary fields, the above actions reduce to the conventional ones:

$$S_{\text{particle}} \Rightarrow \int d\tau e^{-1} \dot{x}^\mu \dot{x}^\nu g_{\mu\nu} - \frac{1}{4} m^2 e,$$

$$S_{\text{string}} \Rightarrow \frac{1}{2\pi\alpha'} \int d^2\sigma - \frac{1}{2} \sqrt{-hh^{ij}} \partial_i x^\mu \partial_j x^\nu g_{\mu\nu} + \frac{1}{2} \epsilon^{ij} \partial_i x^\mu \partial_j x^\nu B_{\mu\nu} + \frac{1}{2} \epsilon^{ij} \partial_i \tilde{x}_\mu \partial_j x^\mu.$$

## III. $\kappa$ -symmetric doubled-yet-gauged Green-Schwarz superstring, unifying IIA & IIB JHP 2016

$$S_{\text{GS}} = \frac{1}{2\pi\alpha'} \int d^2\sigma - \frac{1}{2} \sqrt{-hh^{ij}} \Pi_i^M \Pi_j^N \mathcal{H}_{MN} - \epsilon^{ij} D_i x^M (A_{jM} - \Pi_{jM}),$$

$$\text{where } \Pi_i^M = D_i x^M - \Pi_i^M \text{ and } \Sigma_i^M = \tilde{\theta} \gamma^M \partial_i \theta + \tilde{\psi} \gamma^M \partial_i \psi.$$



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On the other hand, upon the generic  $(n, \bar{n})$  DFT backgrounds, the auxiliary gauge potential decomposes into three parts:

$$A_\mu = K_{\mu\rho} H^{\rho\nu} A_\nu + X_\mu^i Y_i^\nu A_\nu + \bar{X}_\mu^{\bar{i}} \bar{Y}_{\bar{i}}^\nu A_\nu .$$

- The first part appears quadratically, which leads to Gaussian integral.
- The second and third parts appear linearly, as Lagrange multipliers, to prescribe

i) Particle freezes over the  $(n + \bar{n})$  dimensions

$$X_\mu^i \dot{x}^\mu \equiv 0, \quad \bar{X}_\mu^{\bar{i}} \dot{x}^\mu \equiv 0 .$$

Remaining orthogonal directions are described by a reduced action:

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Covariant derivatives and curvatures in Stringy Gravity feature two stages:  
**'semi-covariance'** and **'complete covariantization'**.

- **Semi-covariant derivative :**

Jeon-Lee-JHP 2010, 2011

$$\nabla_C T_{A_1 A_2 \dots A_n} := \partial_C T_{A_1 A_2 \dots A_n} - \omega_T \Gamma^B_{BC} T_{A_1 A_2 \dots A_n} + \sum_{i=1}^n \Gamma_{CA_i}{}^B T_{A_1 \dots A_{i-1} B A_{i+1} \dots A_n},$$

for which the stringy Christoffel connection can be uniquely fixed,

$$\Gamma_{CAB} = 2(P\partial_C P\bar{P})_{[AB]} + 2(\bar{P}_{[A}{}^D \bar{P}_{B]}{}^E - P_{[A}{}^D P_{B]}{}^E) \partial_D P_{EC} - \frac{4}{D-1} (\bar{P}_{C[A} \bar{P}_{B]}{}^D + P_{C[A} P_{B]}{}^D) (\partial_D d + (P\partial^E P\bar{P})_{[ED]})$$

by demanding the compatibility,  $\nabla_A P_{BC} = \nabla_A \bar{P}_{BC} = \nabla_A d = 0$ , and some torsionless conditions.

\* There are no normal coordinates where  $\Gamma_{CAB}$  would vanish point-wise: Equivalence Principle is broken for string (*i.e.* extended object) but recoverable for point particle.

- **Semi-covariant Riemann curvature :**

$$S_{ABCD} = S_{[AB][CD]} = S_{CDAB} := \frac{1}{2} (R_{ABCD} + R_{CDAB} - \Gamma^E_{AB} \Gamma_{ECD}), \quad S_{[ABC]D} = 0,$$

where  $R_{ABCD}$  denotes the ordinary "field strength":  $R_{CDAB} = \partial_A \Gamma_{BCD} - \partial_B \Gamma_{ACD} + \Gamma_{AC}{}^E \Gamma_{BED} - \Gamma_{BC}{}^E \Gamma_{AED}$ .

By construction, it varies as 'total derivative':  $\delta S_{ABCD} = \nabla_{[A} \delta \Gamma_{B]CD} + \nabla_{[C} \delta \Gamma_{D]AB}$ .

- **Semi-covariant 'Master' derivative :**

$$\mathcal{D}_A := \partial_A + \Gamma_A + \Phi_A + \bar{\Phi}_A = \nabla_A + \Phi_A + \bar{\Phi}_A.$$

The two spin connections for the  $\mathbf{Spin}(1, D-1)_L \times \mathbf{Spin}(D-1, 1)_R$  local Lorentz symmetries are determined in terms of the stringy Christoffel connection by requiring the compatibility with DFT vielbeins,

$$\mathcal{D}_A V_{Bp} = \nabla_A V_{Bp} + \Phi_{Ap}{}^q V_{Bq} = 0, \quad \mathcal{D}_A \bar{V}_{B\bar{p}} = \nabla_A \bar{V}_{B\bar{p}} + \bar{\Phi}_{A\bar{p}}{}^{\bar{q}} \bar{V}_{B\bar{q}} = 0.$$

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## • Complete covariantization

– Tensors,

$$P_C{}^D \bar{P}_{A_1}{}^{B_1} \dots \bar{P}_{A_n}{}^{B_n} \nabla_D T_{B_1 \dots B_n} \implies \mathcal{D}_\rho T_{\bar{q}_1 \bar{q}_2 \dots \bar{q}_n},$$

$$\bar{P}_C{}^D P_{A_1}{}^{B_1} \dots P_{A_n}{}^{B_n} \nabla_D T_{B_1 \dots B_n} \implies \mathcal{D}_{\bar{\rho}} T_{q_1 q_2 \dots q_n},$$

$$\mathcal{D}^\rho T_{\rho \bar{q}_1 \bar{q}_2 \dots \bar{q}_n}, \quad \mathcal{D}^{\bar{\rho}} T_{\bar{\rho} q_1 q_2 \dots q_n}; \quad \mathcal{D}_\rho \mathcal{D}^\rho T_{\bar{q}_1 \bar{q}_2 \dots \bar{q}_n}, \quad \mathcal{D}_{\bar{\rho}} \mathcal{D}^{\bar{\rho}} T_{q_1 q_2 \dots q_n}.$$

– Spinors,  $\rho^\alpha, \rho'^{\bar{\alpha}}, \psi_{\bar{\rho}}^\alpha, \psi'_{\rho}{}^{\bar{\alpha}},$

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– RR sector,  $C^\alpha{}_{\bar{\alpha}}$   $\mathbf{O}(D, D)$  covariant nilpotent operators

$$\mathcal{D}_\pm C := \gamma^\rho \mathcal{D}_\rho C \pm \gamma^{(D+1)} \mathcal{D}_{\bar{\rho}} C \bar{\gamma}^{\bar{\rho}}, \quad (\mathcal{D}_\pm)^2 = 0 \implies \mathcal{F} := \mathcal{D}_+ C \quad (\text{RR flux}).$$

– Yang-Mills,

$$\mathcal{F}_{\rho\bar{q}} := \mathcal{F}_{AB} V^A{}_\rho \bar{V}^B{}_{\bar{q}} \quad \text{where} \quad \mathcal{F}_{AB} := \nabla_A \mathcal{V}_B - \nabla_B \mathcal{V}_A - i[\mathcal{V}_A, \mathcal{V}_B].$$

– Curvatures,

$$S_{\rho\bar{q}} := S_{AB} V^A{}_\rho \bar{V}^B{}_{\bar{q}} \quad (\text{Ricci}), \quad S_{(0)} := (P^{AC} P^{BD} - \bar{P}^{AC} \bar{P}^{BD}) S_{ABCD} \quad (\text{scalar}).$$

## • Complete covariantization

– Tensors,

$$P_C{}^D \bar{P}_{A_1}{}^{B_1} \dots \bar{P}_{A_n}{}^{B_n} \nabla_D T_{B_1 \dots B_n} \implies \mathcal{D}_\rho T_{\bar{q}_1 \bar{q}_2 \dots \bar{q}_n},$$

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**Assuming  $(0, 0)$  Riemannian background,  $\{e_{\mu}{}^{\rho}, \bar{e}_{\mu}{}^{\bar{\rho}}, B, \phi\}$ , they reduce e.g. to**

- Generalized Geometry Hitchin 2002, Gualtieri 2004, Waldram *et al.* 2008, 2011 :

$$\begin{aligned} \mathcal{D}_{\bar{\rho}} T_q &= \frac{1}{\sqrt{2}} \left( \partial_{\bar{\rho}} T_q + \omega_{\bar{\rho}qr} T^r + \frac{1}{2} H_{\bar{\rho}qr} T^r \right), \\ \gamma^{\rho} \mathcal{D}_{\rho} \rho &= \frac{1}{\sqrt{2}} \gamma^m \left( \partial_m \rho + \frac{1}{4} \omega_{mnp} \gamma^{np} \rho + \frac{1}{24} H_{mnp} \gamma^{np} \rho - \partial_m \phi \right). \end{aligned}$$

- With  $e_{\mu}{}^{\rho} \equiv \bar{e}_{\mu}{}^{\bar{\rho}}$ ,  $H$ -twisted & democratic RR Bergshoeff, Kallosh, Ortín, Roest, Van Proeyen 2001 :

$$\mathcal{D}_+ \Rightarrow d + H \wedge, \quad \mathcal{D}_- \Rightarrow \star (d + H \wedge) \star.$$

- The scalar curvature gives the closed string effective action :

$$\int e^{-2d} S_{(0)} = \int \sqrt{|g|} e^{-2\phi} \left( R_g + 4 \partial_{\mu} \phi \partial^{\mu} \phi - \frac{1}{12} H_{\lambda\mu\nu} H^{\lambda\mu\nu} \right).$$

**These results show how closed string massless sector,  $\{g_{\mu\nu}, B_{\mu\nu}, \phi\}$ , should couple to extra matters minimally, while forming (pure) Stringy Gravity.**

Equipped with the semi-covariant derivatives, one can construct, e.g.

- $D = 10$  Maximally Supersymmetric Double Field Theory,

Jeon-Lee-JHP-Suh 2012

$$\mathcal{L}_{\text{type II}} = e^{-2d} \left[ \frac{1}{8} S_{(0)} + \frac{1}{2} \text{Tr}(\mathcal{F}\bar{\mathcal{F}}) + i\bar{\rho}\mathcal{F}\rho' + i\bar{\psi}_{\bar{p}}\gamma_q\mathcal{F}\bar{\gamma}^{\bar{p}}\psi'^q + i\frac{1}{2}\bar{\rho}\gamma^{\rho}\mathcal{D}_{\rho}\rho - i\frac{1}{2}\bar{\rho}'\bar{\gamma}^{\bar{p}}\mathcal{D}_{\bar{p}}\rho' \right. \\ \left. - i\bar{\psi}^{\bar{p}}\mathcal{D}_{\bar{p}}\rho - i\frac{1}{2}\bar{\psi}^{\bar{p}}\gamma^q\mathcal{D}_q\psi_{\bar{p}} + i\bar{\psi}'^{\rho}\mathcal{D}_{\rho}\rho' + i\frac{1}{2}\bar{\psi}'^{\rho}\bar{\gamma}^{\bar{q}}\mathcal{D}_{\bar{q}}\psi'_{\rho} \right]$$

which unifies IIA and IIB SUGRAs, thanks to the twofold spin groups.

- Minimal coupling to the  $D = 4$  Standard Model,

Kangsin Choi & JHP 2015 [PRL]

$$\mathcal{L}_{\text{SM}} = e^{-2d} \left[ \frac{1}{16\pi G_N} S_{(0)} \right. \\ \left. + \sum_{\mathcal{V}} P^{AB}\bar{P}^{CD}\text{Tr}(\mathcal{F}_{AC}\mathcal{F}_{BD}) + \sum_{\psi} \bar{\psi}\gamma^a\mathcal{D}_a\psi + \sum_{\psi'} \bar{\psi}'\bar{\gamma}^{\bar{a}}\mathcal{D}_{\bar{a}}\psi' \right. \\ \left. - \mathcal{H}^{AB}(\mathcal{D}_A\phi)^\dagger\mathcal{D}_B\phi - V(\phi) + y_d\bar{q}\cdot\phi\mathbf{d} + y_u\bar{q}\cdot\tilde{\phi}\mathbf{u} + y_e\bar{l}'\cdot\phi\mathbf{e}' \right]$$

Every single term above is completely covariant, w.r.t.  $O(D, D)$ , diffeomorphisms, and twofold local Lorentz symmetries,  $\text{Spin}(1, D-1)_L \times \text{Spin}(D-1, 1)_R$ .

## Derivation of the Einstein Double Field Equations

Henceforth, we consider a general action for Stringy Gravity coupled to matter fields,  $\Upsilon_a$ ,

$$\int_{\Sigma} e^{-2d} \left[ \frac{1}{16\pi G} S_{(0)} + L_{\text{matter}} \right],$$

where  $S_{(0)}$  is the stringy scalar curvature and  $L_{\text{matter}}$  is the matter Lagrangian equipped with the completely covariantized master derivatives,  $\mathcal{D}_M$ . The integral is taken over a section,  $\Sigma$ .

We seek the variation of the action induced by all the fields,  $d$ ,  $V_{Ap}$ ,  $\bar{V}_{A\bar{p}}$ ,  $\Upsilon_a$ .

– Firstly, the pure Stringy Gravity term transforms, up to total derivatives ( $\simeq$ ), as

$$\delta \left( e^{-2d} S_{(0)} \right) \simeq 4e^{-2d} \left( \bar{V}^{B\bar{q}} \delta V_B^p S_{p\bar{q}} - \frac{1}{2} \delta d S_{(0)} \right)$$

– Secondly, the matter Lagrangian transforms as

$$\delta \left( e^{-2d} L_{\text{matter}} \right) \simeq e^{-2d} \left( -2\bar{V}^{A\bar{q}} \delta V_A^p K_{p\bar{q}} + \delta d T_{(0)} + \delta \Upsilon_a \frac{\delta L_{\text{matter}}}{\delta \Upsilon_a} \right)$$

where we have been naturally led to define

$$K_{p\bar{q}} := \frac{1}{2} \left( V_{Ap} \frac{\delta L_{\text{matter}}}{\delta \bar{V}_A^{\bar{q}}} - \bar{V}_{A\bar{q}} \frac{\delta L_{\text{matter}}}{\delta V_A^p} \right), \quad T_{(0)} := e^{2d} \times \frac{\delta \left( e^{-2d} L_{\text{matter}} \right)}{\delta d}.$$

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- Combining the two results, the variation of the action reads

$$\begin{aligned} & \delta \int_{\Sigma} e^{-2d} \left[ \frac{1}{16\pi G} \mathcal{S}_{(0)} + L_{\text{matter}} \right] \\ &= \int_{\Sigma} e^{-2d} \left[ \frac{1}{4\pi G} \bar{V}^{A\bar{q}} \delta V_A{}^p (S_{p\bar{q}} - 8\pi G K_{p\bar{q}}) - \frac{1}{8\pi G} \delta d (\mathcal{S}_{(0)} - 8\pi G T_{(0)}) + \delta \Upsilon_a \frac{\delta L_{\text{matter}}}{\delta \Upsilon_a} \right] \end{aligned}$$

Hence, the equations of motion are for now exhaustively,

$$S_{p\bar{q}} = 8\pi G K_{p\bar{q}}, \quad \mathcal{S}_{(0)} = 8\pi G T_{(0)}, \quad \frac{\delta L_{\text{matter}}}{\delta \Upsilon_a} = 0.$$

- Specifically when the variation is generated by diffeomorphisms, we have  $\delta_{\xi} \Upsilon_a = \hat{\mathcal{L}}_{\xi} \Upsilon_a$  and

$$\delta_{\xi} d = -\frac{1}{2} e^{2d} \hat{\mathcal{L}}_{\xi} (e^{-2d}) = -\frac{1}{2} \mathcal{D}_A \xi^A, \quad \bar{V}^{A\bar{q}} \delta_{\xi} V_A{}^p = \bar{V}^{A\bar{q}} \hat{\mathcal{L}}_{\xi} V_A{}^p = 2\mathcal{D}_{[A} \xi_{B]} \bar{V}^{A\bar{q}} V^{Bp}.$$

Substituting these, the diffeomorphic invariance of the action implies

$$0 = \int_{\Sigma} e^{-2d} \left[ \frac{1}{8\pi G} \xi^B \mathcal{D}^A \left\{ 4V_{[A}{}^p \bar{V}_{B]}{}^{\bar{q}} (S_{p\bar{q}} - 8\pi G K_{p\bar{q}}) - \frac{1}{2} \mathcal{J}_{AB} (\mathcal{S}_{(0)} - 8\pi G T_{(0)}) \right\} + \delta_{\xi} \Upsilon_a \frac{\delta L_{\text{matter}}}{\delta \Upsilon_a} \right]$$

which leads to the definitions of the off-shell conserved **stringy Einstein curvature**,

$$G_{AB} := 4V_{[A}{}^p \bar{V}_{B]}{}^{\bar{q}} S_{p\bar{q}} - \frac{1}{2} \mathcal{J}_{AB} \mathcal{S}_{(0)}, \quad \mathcal{D}_A G^{AB} = 0 \quad (\text{off-shell}),$$

and the on-shell conserved **stringy Energy-Momentum tensor**,

JHP-Rey-Rim-Sakatani 2015

$$T_{AB} := 4V_{[A}{}^p \bar{V}_{B]}{}^{\bar{q}} K_{p\bar{q}} - \frac{1}{2} \mathcal{J}_{AB} T_{(0)}, \quad \mathcal{D}_A T^{AB} = 0 \quad (\text{on-shell}).$$

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- Since  $G_{AB}$  and  $T_{AB}$  each have  $D^2 + 1$  components decomposable as

$$V^A{}_\rho \bar{V}^B{}_{\bar{q}} G_{AB} = 2S_{\rho\bar{q}}, \quad G^A{}_A = -DS_{(0)}, \quad V^A{}_\rho \bar{V}^B{}_{\bar{q}} T_{AB} = 2K_{\rho\bar{q}}, \quad T^A{}_A = -DT_{(0)},$$

the equations of motion of the DFT vielbeins and dilaton can be unified into a single expression:

Einstein Double Field Equations

$$G_{AB} = 8\pi G T_{AB}$$

which is naturally consistent with the central idea that Stringy Gravity treats the entire closed string massless sector as geometrical stringy graviton fields.

## Einstein Double Field Equations

$$G_{AB} = 8\pi GT_{AB}$$

- Restricting to the  $(0, 0)$  Riemannian backgrounds, the EDFE decompose into

$$R_{\mu\nu} + 2\nabla_{\mu}(\partial_{\nu}\phi) - \frac{1}{4}H_{\mu\rho\sigma}H_{\nu}{}^{\rho\sigma} = 8\pi GK_{(\mu\nu)},$$

$$\nabla^{\rho}\left(e^{-2\phi}H_{\rho\mu\nu}\right) = 16\pi Ge^{-2\phi}K_{[\mu\nu]},$$

$$R + 4\Box\phi - 4\partial_{\mu}\phi\partial^{\mu}\phi - \frac{1}{12}H_{\lambda\mu\nu}H^{\lambda\mu\nu} = 8\pi GT_{(0)}.$$



- For other non-Riemannian cases,  $(n, \bar{n}) \neq (0, 0)$ , EDFE govern the dynamics of the 'chiral' gravities, such as Newton-Cartan, Carroll, and Gomis-Ooguri, etc.

## Einstein Double Field Equations

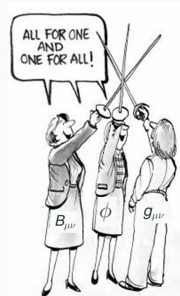
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**Examples:**  $T_{AB} := 4V_{[A}{}^{\rho} \bar{V}_{B]}{}^{\bar{q}} K_{\rho\bar{q}} - \frac{1}{2} \mathcal{J}_{AB} T_{(0)}$

- RR sector,

$$L_{\text{RR}} = \frac{1}{2} \text{Tr}(\mathcal{F}\bar{\mathcal{F}}), \quad K_{p\bar{q}} = -\frac{1}{4} \text{Tr}(\gamma_p \mathcal{F} \bar{\gamma}_{\bar{q}} \bar{\mathcal{F}}), \quad T_{(0)} = 0.$$

- Spinor field,

$$L_{\psi} = \bar{\psi} \gamma^{\rho} \mathcal{D}_{\rho} \psi + m_{\psi} \bar{\psi} \psi, \quad K_{p\bar{q}} = -\frac{1}{4} (\bar{\psi} \gamma_p \mathcal{D}_{\bar{q}} \psi - \mathcal{D}_{\bar{q}} \bar{\psi} \gamma_p \psi), \quad T_{(0)} \equiv 0.$$

- Scalar field,

$$L_{\varphi} = -\frac{1}{2} \mathcal{H}^{MN} \partial_M \varphi \partial_N \varphi - V(\varphi), \quad K_{p\bar{q}} = \partial_p \varphi \partial_{\bar{q}} \varphi, \quad T_{(0)} = -2L_{\varphi}.$$

- Fundamental string: with  $D_i y^M = \partial_i y^M - \mathcal{A}_i^M$  (doubled-yet-gauged),

$$e^{-2d} L_{\text{string}} = \frac{1}{4\pi\alpha'} \int d^2\sigma \left[ -\frac{1}{2} \sqrt{-h} h^{ij} D_i y^M D_j y^N \mathcal{H}_{MN}(y) - \epsilon^{ij} D_i y^M \mathcal{A}_{jM} \right] \delta^D(x - y(\sigma)),$$

$$K_{p\bar{q}} = \frac{1}{4\pi\alpha'} \int d^2\sigma \sqrt{-h} h^{ij} (D_i y)_{\rho} (D_j y)_{\bar{q}} e^{2d(x)} \delta^D(x - y(\sigma)), \quad T_{(0)} = 0.$$

– More examples in our paper include Yang-Mills, point particle, superstring, etc.

# Conclusion

- String theory may predict its own gravity, *i.e.* Stringy Gravity (DFT), rather than GR.
- Stringy Gravity may be formulated in 'doubled-yet-gauged' spacetime, unifies Riemannian SUGRA and non-Riemannian Newton-Cartan, Carroll, Gomis-Ooguri, *etc.* and deserves further explorations, *e.g.*

$O(D, D)$  covariant holographic dual of Stringy Gravity?



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$O(D, D)$  covariant holographic dual of Stringy Gravity?

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- String theory may predict its own gravity, *i.e.* Stringy Gravity (DFT), rather than GR.
- Stringy Gravity may be formulated in ‘doubled-yet-gauged’ spacetime, unifies Riemannian SUGRA and non-Riemannian Newton-Cartan, Carroll, Gomis-Ooguri, *etc.* and deserves further explorations, *e.g.*

$O(D, D)$  covariant holographic dual of Stringy Gravity?

**Thank you**

*One must be prepared to follow up the consequence of theory, and feel that one just has to accept the consequences no matter where they lead.*

*– Paul Dirac –*

# Einstein Double Field Theories

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## Core idea: string theory predicts its own gravity rather than GR

In General Relativity the metric  $g_{\mu\nu}$  is the *only* geometric and gravitational field, whereas in string theory the closed-string massless sector comprises a two-form potential  $B_{\mu\nu}$  and the string dilaton  $\Phi$  in addition to the metric. Furthermore, these three fields transform like each other under T-duality. This hints at a natural augmentation of GR: upon treating the whole closed string massless sector as string gravitation fields, Double Field Theory [1, 2] now evolves into “String Gravity”. Equipped with an O(D, D) constant differential geometry beyond Riemann [3], we spell out the definitions of the string Einstein curvature tensor and the string Energy-Momentum tensor. Equating them, all the equations of motion of the closed string massless sector are spelled into a single expression [4].

$$G_{AB} = -3\partial^2 T_{AB} \quad (1)$$

## Which we dub the Einstein Double Field Equations.

### Double Field Theory as String Gravity

#### • Built in symmetries & Notation:

- O(D, D) T-duality

- DFT diffeomorphisms (ordinary diffeomorphisms plus D-field gauge symmetry)

- Twisted local Lorentz symmetry,  $\text{Spin}(D, D-1)$

- Two locally inertial frames: spacetime, separately for the left and the right nodes.

Index	Representation	Metric (raising/lowering indices)
A, B, ...	M, N, ...	$\eta_{AB} = \begin{pmatrix} 1 & \\ & -1 \end{pmatrix}$
O(D, D) vector	$\mathcal{X}^{AB}$	$\mathcal{X}^{AB} = \begin{pmatrix} 1 & \\ & -1 \end{pmatrix}$
$\mu, \nu, \dots$	Spin(D, D-1) vector	$\eta_{\mu\nu} = \text{diag}(-, +, \dots, +)$
$\alpha, \beta, \dots$	Spin(D, D-1) spinor	$C_{\alpha\beta} = (C\gamma^{\mu\nu}C)^{-1}$
$\hat{\mu}, \hat{\nu}, \dots$	Spin(D-1, 1) vector	$\eta_{\hat{\mu}\hat{\nu}} = \text{diag}(+, \dots, +, -)$
$\hat{\alpha}, \hat{\beta}, \dots$	Spin(D-1, 1) spinor	$C_{\hat{\alpha}\hat{\beta}} = (C\gamma^{\mu\nu}C)^{-1}$

The O(D, D) metric  $\mathcal{X}^{AB}$  defines doubled coordinates into twin sets:  $x^{\mu} = (x^{\mu}, \tilde{x}_\mu) = (P^{\mu}, \theta_\mu)$ .

#### • Doubled-yet-gauged spacetime:

The doubled coordinates are “gauged” through a certain equivalence relation,  $x^{\mu} = x^{\nu} = x^{\mu}$ , such that each equivalence class, or gauge orbit as  $\mathbb{R}^D$ , corresponds to a single physical point in  $\mathbb{R}^D$ . This implies a section condition,  $\partial_\mu \tilde{x}^\mu = 0$ , which can be consistently solved by setting  $\tilde{P}^\mu = 0$ .

#### • String graviton fields (closed-string massless sector): $\{A_{\mu\nu}, \Phi, \mathcal{H}_\mu\}$

Defining properties of the DFT-metric:

$$\eta_{AB} = \eta_{NM}, \quad \eta_A^{\ \mu} = \eta_{\mu}^{\ A}, \quad \eta_{\mu}^{\ \mu} = \delta_{\mu}^{\ \mu}, \quad \eta_{\tilde{\mu}}^{\ \tilde{\mu}} = -\delta_{\tilde{\mu}}^{\ \tilde{\mu}}, \quad (2)$$

set of pair symmetric and orthogonal projectors:

$$P_{\mu\nu} = \frac{1}{2}(\delta_{\mu\nu} + \eta_{\mu\nu}), \quad P_{\tilde{\mu}\tilde{\nu}} = \frac{1}{2}(\delta_{\tilde{\mu}\tilde{\nu}} - \eta_{\tilde{\mu}\tilde{\nu}}), \\ P_{\mu\tilde{\nu}} = \frac{1}{2}(\delta_{\mu\tilde{\nu}} - \eta_{\mu\tilde{\nu}}), \quad P_{\tilde{\mu}\nu} = \frac{1}{2}(\delta_{\tilde{\mu}\nu} + \eta_{\tilde{\mu}\nu}).$$

Further, taking the “square root” of the projectors, we acquire a pair of DFT vielbeins,

$$P_{\mu\alpha} = V_{\mu}^{\ \alpha} \gamma_{\alpha}^{\ \mu}, \quad P_{\tilde{\mu}\hat{\alpha}} = V_{\tilde{\mu}}^{\ \hat{\alpha}} \gamma_{\hat{\alpha}}^{\ \tilde{\mu}},$$

satisfying their own defining properties,

$$V_{\mu\alpha} V_{\nu}^{\ \alpha} = \eta_{\mu\nu}, \quad V_{\tilde{\mu}\hat{\alpha}} V_{\tilde{\nu}}^{\ \hat{\alpha}} = \eta_{\tilde{\mu}\tilde{\nu}}, \quad V_{\mu\alpha} V_{\tilde{\mu}\hat{\alpha}} = 0, \quad V_{\tilde{\mu}\hat{\alpha}} V_{\mu\alpha} = \eta_{\tilde{\mu}\mu},$$

The more general solutions to (2) can be classified by two non-negative integers (n, k) [5],

$$\eta_{AB} = \begin{pmatrix} \mathbb{I}^{n \times n} & & & & \\ & \mathbb{I}^{k \times k} & & & \\ & & -\mathbb{I}^{(D-n-k) \times (D-n-k)} & & \\ & & & \mathbb{I}^{(D-n-k) \times (D-n-k)} & \\ & & & & \mathbb{I}^{(D-n-k) \times (D-n-k)} \end{pmatrix}$$

where  $1 \leq n, k, n+k \leq D$  and

$$\mathbb{I}^{n \times n} \mathbb{I}^{k \times k} = 0, \quad \mathbb{I}^{n \times n} \mathbb{I}^{(D-n-k) \times (D-n-k)} = 0, \quad \mathbb{I}^{k \times k} \mathbb{I}^{(D-n-k) \times (D-n-k)} = 0,$$

String being chiral and anti-chiral even and a direction:  $\mathbb{I}^{n \times n} \mathbb{I}^{(D-n-k) \times (D-n-k)} = \mathbb{I}^{(D-n-k) \times (D-n-k)}$ . Examples include (0, k) Riemannian geometry as  $n=0$ ,  $\eta_{\mu\nu} = \mathbb{I}^{D \times D}$ ,  $\mathbb{I}^{1 \times 1}$  General-Relativity non-solution; background, (1, 0) Newton-Cartan gravity, and (D-1, 0) Cahill gravity.

#### • Covariant derivative:

The “metric” constant derivative,  $\partial_A = \partial_\mu + \partial_{\tilde{\mu}}$  is characterized by compatibility:

$$\partial_A \mathcal{X}^{BC} = \partial_\mu \mathcal{X}^{BC} = \partial_{\tilde{\mu}} \mathcal{X}^{BC} = 0, \quad \partial_\mu \eta_{\alpha\beta} = \partial_{\tilde{\mu}} \eta_{\alpha\beta} = \partial_\mu \eta_{\hat{\alpha}\hat{\beta}} = \partial_{\tilde{\mu}} \eta_{\hat{\alpha}\hat{\beta}} = 0.$$

The string Christoffel symbols [6]

$$\Gamma_{AB}^C = \frac{1}{2} (\partial_A \mathcal{X}^{BC} + \partial_B \mathcal{X}^{AC} - \partial_C \mathcal{X}^{AB}), \quad \partial_{\tilde{\mu}} \mathcal{X}^{AB} = -\frac{1}{2} (\partial_{\tilde{\mu}} \mathcal{X}^{AB} + \partial_{\tilde{\nu}} \mathcal{X}^{A\tilde{\nu}} + \partial_{\tilde{\nu}} \mathcal{X}^{A\tilde{\nu}}) \partial_{\tilde{\nu}} \mathcal{X}^{BC}$$

and the connections are  $\partial_\mu = \partial_{\tilde{\mu}} \mathcal{X}^{AB} \partial_A \partial_B$ ,  $\partial_{\tilde{\mu}} = \partial_{\tilde{\nu}} \mathcal{X}^{AB} \partial_A \partial_B$ ,  $\partial_{\tilde{\mu}} \partial_{\tilde{\nu}} = \partial_{\tilde{\nu}} \partial_{\tilde{\mu}}$ .

In String Gravity, there are no normal coordinates where  $\Gamma_{ABC} = 0$ ,  $\partial_{\tilde{\mu}} \eta_{\alpha\beta} = 0$ ,  $\partial_{\tilde{\nu}} \eta_{\hat{\alpha}\hat{\beta}} = 0$  and the Equivalence Principle holds for point particles but is generally broken for strings (i.e. extended objects).

#### • Scalar and “Ricci” curvature:

The one-constant Riemann curvature in String Gravity is defined by

$$R_{ABCD} = \frac{1}{2} (\partial_A \partial_C \mathcal{X}^{BD} - \partial_B \partial_C \mathcal{X}^{AD} - \partial_A \partial_D \mathcal{X}^{BC} + \partial_B \partial_D \mathcal{X}^{AC})$$

where  $\partial_A = \partial_\mu \mathcal{X}^{AB} \partial_B$ ,  $\partial_{\tilde{\mu}} = \partial_{\tilde{\nu}} \mathcal{X}^{AB} \partial_B$ ,  $\partial_A \partial_{\tilde{\mu}} = \partial_{\tilde{\mu}} \partial_A$  the “half strength” of  $\Gamma_{ABC}$ .

The completely covariant “Ricci” and scalar curvatures are, with  $\mathcal{X}^{AB} = \delta_{AB}$ ,

$$S_{AB} = \mathcal{X}^{CD} R_{CABD}, \quad S_{\tilde{\mu}\tilde{\nu}} = \mathcal{X}^{CD} R_{\tilde{\mu}C\tilde{\nu}D}, \quad S_{AB} = \mathcal{X}^{CD} R_{ACBD}.$$

While  $\partial_{\tilde{\mu}} S_{AB}$  corresponds to the original DFT Laplacian density [1, 2], or the “pure” String Gravity, the mixed curvature derivative from  $\partial_{\tilde{\mu}} \mathcal{X}^{AB}$  is related to a zero tensor field, e.g. type II mutually supersymmetric DFT [7] of the Standard Model [8].

## Derivation of Einstein Double Field Equations

Variation of the action for String Gravity coupled to generic matter fields,  $T_{\mu\nu}$ , gives

$$\delta \int d^D x \sqrt{-g} \left[ \frac{1}{2\kappa^2} (R - \frac{1}{2} \partial_\mu \Phi \partial^\mu \Phi - \frac{1}{4} \partial_\mu B_{\nu\rho} \partial^\mu B^{\nu\rho}) - \frac{1}{2} \partial_\mu \Phi \partial^\mu \Phi - \frac{1}{4} \partial_\mu B_{\nu\rho} \partial^\mu B^{\nu\rho} + \mathcal{L}_M \right]$$

where the second line is for generic variations and the third line is specifically for diffeomorphic transformations. We are naturally led to define

$$K_{\mu\nu} = \frac{1}{2} \left( V_\mu^\alpha \frac{\delta \mathcal{L}_{\text{matter}}}{\delta V^\alpha} - V_\mu^{\tilde{\alpha}} \frac{\delta \mathcal{L}_{\text{matter}}}{\delta V^{\tilde{\alpha}}} \right), \quad T_{\mu\nu} = \kappa^2 \delta \left( \frac{\delta \mathcal{L}_{\text{matter}}}{\delta \mathcal{X}^{AB}} \right)$$

and subsequently the string Einstein curvature,  $G_{AB}$ , and Energy-Momentum tensor,  $T_{AB}$ ,

$$G_{AB} = \partial_A \mathcal{X}^{CD} \partial_B \mathcal{X}_{CD} - \frac{1}{2} \mathcal{X}^{AB} S_{CD} \mathcal{X}^{CD}, \quad \mathcal{D}_A \mathcal{X}^{AB} = 0 \quad (\text{off-shell}),$$

$$T_{AB} = \partial_A \mathcal{X}^{CD} \partial_B \mathcal{X}_{CD} - \frac{1}{2} \mathcal{X}^{AB} T_{CD} \mathcal{X}^{CD}, \quad \mathcal{D}_A T^A{}_{\tilde{\mu}} = 0 \quad (\text{on-shell}).$$

The equations of motion of the string graviton fields are then spelled into a single expression, the Einstein Double Field Equation (1). Note that  $G_{AB} = -D_{\tilde{\mu}} \mathcal{X}^{\tilde{\mu}A} - D^{\tilde{\mu}A} \mathcal{X}_{\tilde{\mu}B}$ .

Restricting to the (0, k) Riemannian background, the Einstein Double Field Equations reduce to

$$R_{\mu\nu} + 2\partial_{\tilde{\mu}} \mathcal{X}^{\tilde{\mu}\alpha} \partial_{\tilde{\nu}} \mathcal{X}_{\alpha\beta} = -\text{Ric}(\mathcal{X}^{AB}),$$

$$\nabla^{\tilde{\mu}} \mathcal{X}^{\tilde{\mu}\alpha} = 0, \quad \partial_{\tilde{\mu}} \mathcal{X}^{\tilde{\mu}\alpha} = 0,$$

$$R_{\tilde{\mu}\tilde{\nu}} + \partial_{\tilde{\mu}} \mathcal{X}^{\tilde{\mu}\alpha} \partial_{\tilde{\nu}} \mathcal{X}_{\alpha\beta} = -\text{Ric}(\eta_{\hat{\alpha}\hat{\beta}}),$$

$$R_{\tilde{\mu}\nu} + \partial_{\tilde{\mu}} \mathcal{X}^{\tilde{\mu}\alpha} \partial_{\tilde{\nu}} \mathcal{X}_{\alpha\beta} = -\text{Ric}(\mathcal{X}^{AB}),$$

which imply the conservation law,  $\mathcal{D}_A T^{AB} = 0$ , given explicitly by

$$\nabla_{\tilde{\mu}} K_{\mu\nu} - 2\partial_{\tilde{\mu}} \mathcal{X}^{\tilde{\mu}\alpha} K_{\alpha\nu} + 4\partial_{\tilde{\mu}} \mathcal{X}^{\tilde{\mu}\alpha} \partial_{\tilde{\nu}} \mathcal{X}_{\alpha\beta} = 0, \quad \nabla^{\tilde{\mu}} (-\partial_{\tilde{\mu}} \mathcal{X}^{\tilde{\mu}\alpha}) = 0.$$

The Einstein Double Field Equation also governs the dynamics of other non-Riemannian cases, (n, k)  $\neq$  (0, k), where the Riemannian metric,  $\eta_{\mu\nu}$ , cannot be defined.

## Examples

- *Pure String Gravity with cosmological constant:*

$$\frac{1}{2\kappa^2} (R_{\mu\nu} - 2\Lambda g_{\mu\nu}), \quad R_{\mu\nu} = 0, \quad T_{\mu\nu} = \frac{1}{2\kappa^2} \Lambda g_{\mu\nu}$$

- *RR sector, gauge fixed on Spin(D, D)  $\times$  Spin(k):*

$$R_{\mu\nu} = -\frac{1}{2\kappa^2} \Lambda g_{\mu\nu}, \quad R_{\tilde{\mu}\tilde{\nu}} = -\frac{1}{2\kappa^2} \Lambda g_{\tilde{\mu}\tilde{\nu}}$$

$$R_{\mu\tilde{\nu}} = -\frac{1}{2\kappa^2} \Lambda g_{\mu\tilde{\nu}}, \quad R_{\tilde{\mu}\nu} = -\frac{1}{2\kappa^2} \Lambda g_{\tilde{\mu}\nu}, \quad T_{\mu\nu} = 0,$$

where  $\mathcal{X} = \mathcal{I}^{n \times n} \oplus \mathcal{I}^{k \times k} \oplus \mathcal{I}^{(D-n-k) \times (D-n-k)}$  for the RR case and  $\mathcal{I}^{D \times D}$  for “II-extended” connections, ( $\mathcal{D}_A \mathcal{X}^{AB} = 0$  and  $\mathcal{F} = C^2 C^2 C^2$  is its charge conjugate [7]).

- *Spinor field:*  $\mathcal{L}_M = \partial_\mu \psi \partial^\mu \psi + m \psi \psi$ ,  $K_{\mu\nu} = \frac{1}{2} (\partial_\mu \psi \partial_\nu \psi - \partial_\nu \psi \partial_\mu \psi)$ ,  $T_{\mu\nu} = 0$ .

- *Green-Schwarz superstring (to one-fermion):*

$$e^{-2\alpha} \mathcal{L}_{\text{matter}} = \frac{1}{2} \int d^D x \sqrt{-g} \left[ -\frac{1}{2} \partial_\mu \psi \partial^\mu \psi + \mathcal{L}_M - e^{\alpha} \mathcal{X}^{AB} (\partial_\mu \mathcal{X}_{AB} - \partial_{\tilde{\mu}} \mathcal{X}_{AB}) \right] \partial^\mu \psi \partial_\mu \psi,$$

$$K_{\mu\nu} = \frac{1}{2} \int d^D x \sqrt{-g} \left[ \partial_\mu \psi \partial_\nu \psi + \partial_\nu \psi \partial_\mu \psi - \partial_\mu \psi \partial_\nu \psi - \partial_\nu \psi \partial_\mu \psi \right] \partial^\mu \psi \partial_\mu \psi,$$

where  $\mathcal{X}^{AB} = \delta^{\mu\nu} \delta^{\tilde{\mu}\tilde{\nu}} + \theta^{\mu\tilde{\nu}} \delta^{\tilde{\mu}\hat{\alpha}} + \theta^{\tilde{\mu}\hat{\alpha}} \delta^{\mu\alpha} - \mathcal{A}^{\mu\tilde{\nu}} - \mathcal{A}^{\tilde{\mu}\hat{\alpha}}$  (doubled yet gauged [9]).

## Gravitational effect

The regular optical solution to the D = 4 Einstein Double Field Equations shows that String Gravity modifies GR (Schwarzschild geometry), in particular at “short” dimensionless scale,  $r/R$ , at distance normalized by mass length. Spacetime constant. This might show new light upon the dark matter-energy problem, as they arise essentially from “thin-dimension” observations. Furthermore, it would be interesting to view the IR field and DFT dilaton as “dark gravitons” which propagate from the generic location of point particles, which should be defined in string frame [9].

Electric	Point	Hydrogen	Helium	Earth	Solar System	Milky Way	Universe
(0, 0, 0)	(0, 0, 0)	(1, 1, 1)	(1, 1, 1)	(1, 1, 1)	(1, 1, 1)	(1, 1, 1)	(1, 1, 1)


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# Gravitational effect

- The regular spherical solution to the  $D = 4$  Einstein Double Field Equations shows that Stringy Gravity modifies GR (Schwarzschild geometry), in particular at “short” dimensionless scales,  $R/MG$ , *i.e.* distance normalized by mass times Newton constant.

This might shed new light upon the dark matter/energy problems, as they arise essentially from “short distance” observations:

	Electron ( $R \simeq 0$ )	Proton	Hydrogen Atom	Billiard Ball	Earth	Solar System ( $1\text{AU}/M_{\odot}G$ )	Milky Way (visible)	Galaxy Cluster	Universe ( $M \propto R^3$ )
$R/(MG)$	$0^+$	$7.1 \times 10^{38}$	$2.0 \times 10^{43}$	$2.4 \times 10^{26}$	$1.4 \times 10^9$	$1.0 \times 10^8$	$1.5 \times 10^6$	$\sim 10^5$	$0^+$

- Furthermore, it would be intriguing to view the  $B$ -field and DFT dilaton  $d$  as ‘dark gravitons’, since they decouple from the geodesic motion of point particles, which should be defined in string frame.