

Introduction and Overview BSM

Werner Porod

Universität Würzburg

Before Higgs discovery

- something must be at the TeV scale as electroweak symmetry is broken

After Higgs discovery and several searches for new physics

- data described very well by SM
- no hint where at which scale to expect new physics

Before Higgs discovery

- something must be at the TeV scale as electroweak symmetry is broken

After Higgs discovery and several searches for new physics

- data described very well by SM
- no hint where at which scale to expect new physics

EFT vs full models		
full models	vs	EFT
<p><i>“maximum optimism”</i></p> <ul style="list-style-type: none">SM extension addresses a problem deemed “serious”<ul style="list-style-type: none">tuningdark matterbaryogenesis...“solution” induces correlation in data, e.g. “light top partners”		<p><i>“maximum scepticism”</i></p> <ul style="list-style-type: none">deform SM in most general yet consistent way<ul style="list-style-type: none">only SM fields are lightSM symmetriesmaybe extra simplifications to make things manageable“any” correlation in data that can be induced from non-resolved scales in principle

Ch. Englert
summary of last week

Most of the data can be explained (extremely well) by the SM, but

- Flavour
 - hierarchy of fermion masses, in particular ν
 - mixing pattern: small mixing for q versus large mixing for ν

- Anomalies

- $(g - 2)_\mu$
- b-physics:

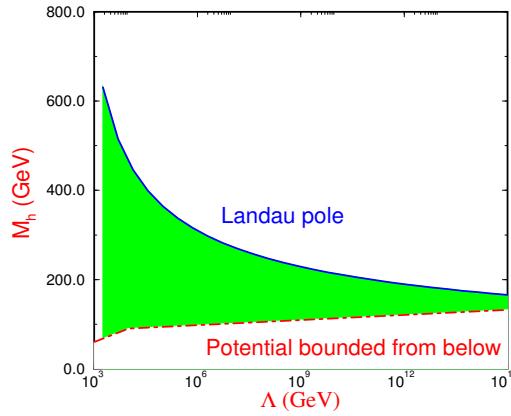
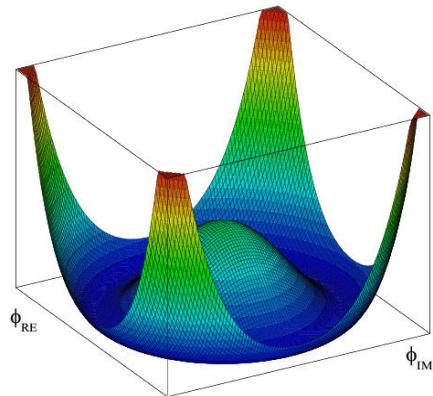
$$R_{D^{(*)}} = \frac{\Gamma(\bar{B} \rightarrow D^{(*)}\tau\bar{\nu})}{\Gamma(\bar{B} \rightarrow D^{(*)}l\bar{\nu})} \quad (l = e, \mu) \quad , \quad R_{K^{(*)}} = \frac{\Gamma(\bar{B} \rightarrow \bar{K}^{(*)}\mu^+\mu^-)}{\Gamma(\bar{B} \rightarrow \bar{K}^{(*)}e^+e^-)}$$

- ...

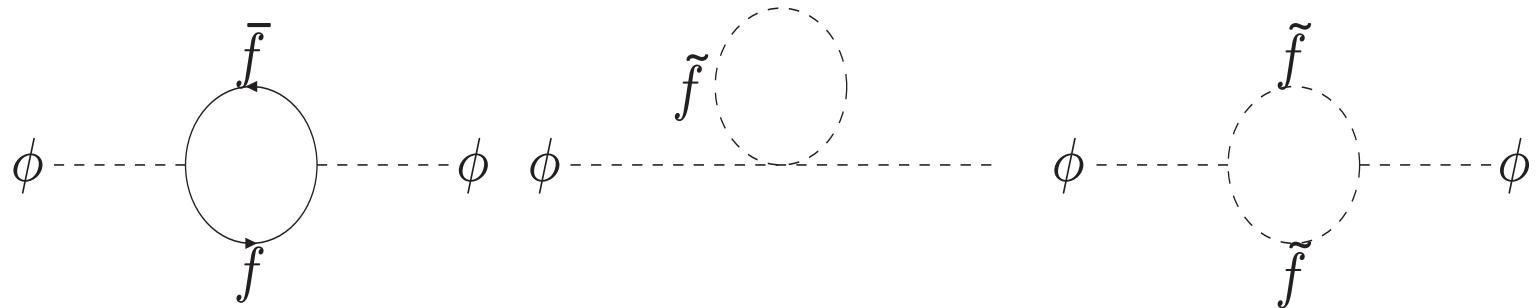
- Cosmology

- Dark matter
- Baryon asymmetry of the Universe

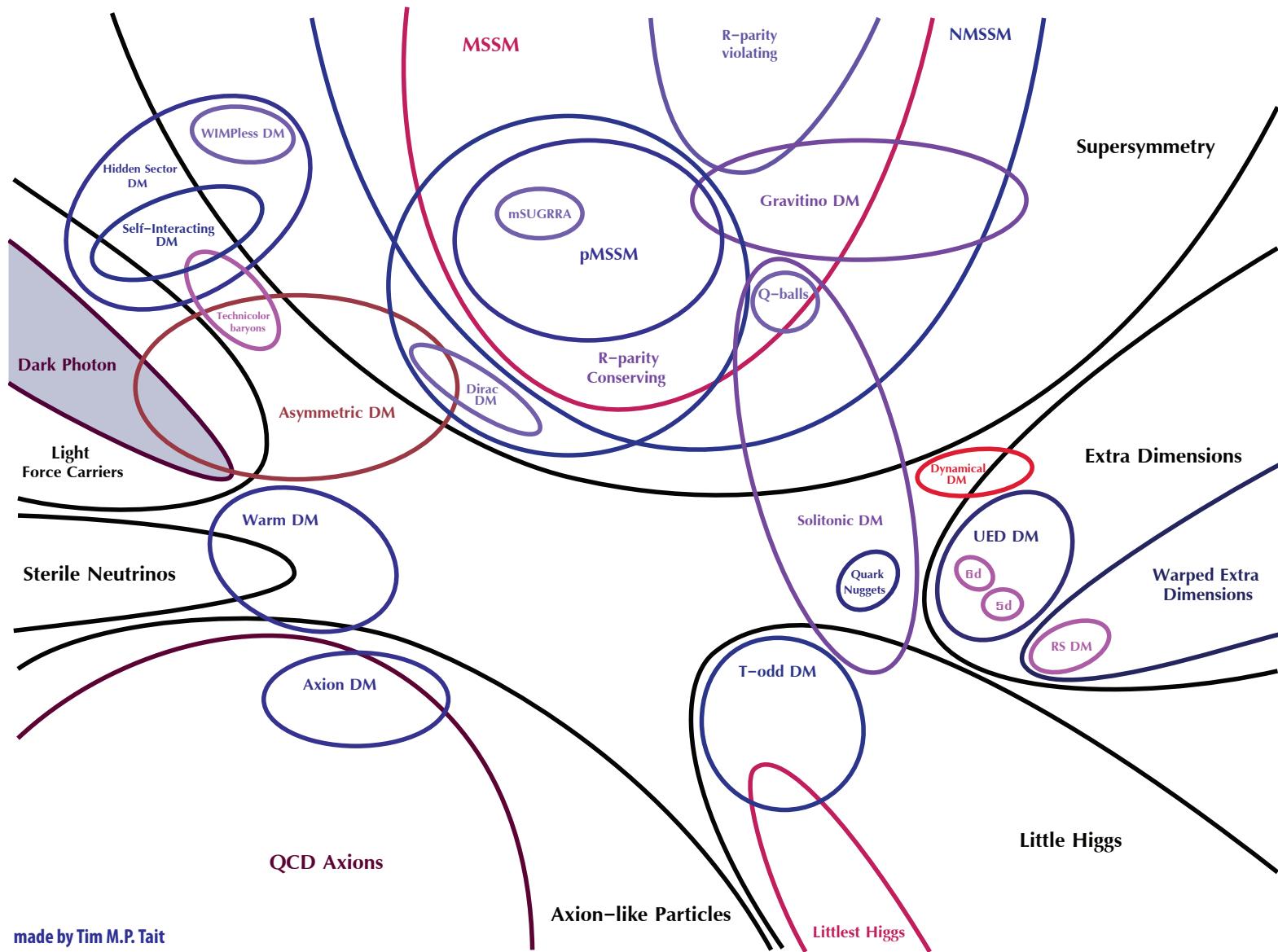
- SM & $m_h = 125.1 \text{ GeV}$: potentially meta-stable (G. Degrassi *et al.*, arXiv:1205.6497)



- "Why does electroweak symmetry break?" or "Why is $\mu^2 < 0$ in the SM?"
- Hierarchy problem (?)

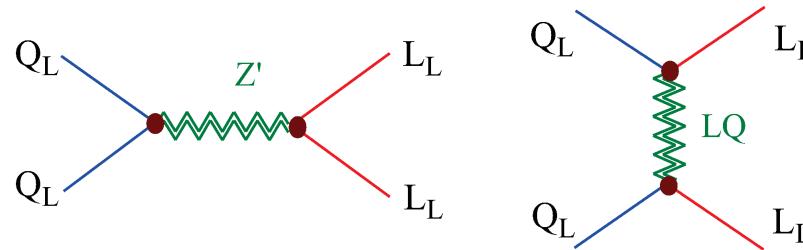


$\delta m_h^2 \propto \Lambda^2$: Sensitivity to highest mass scale of unknown physics



made by Tim M.P. Tait

J.L. Feng et al., arXiv:1401.6085



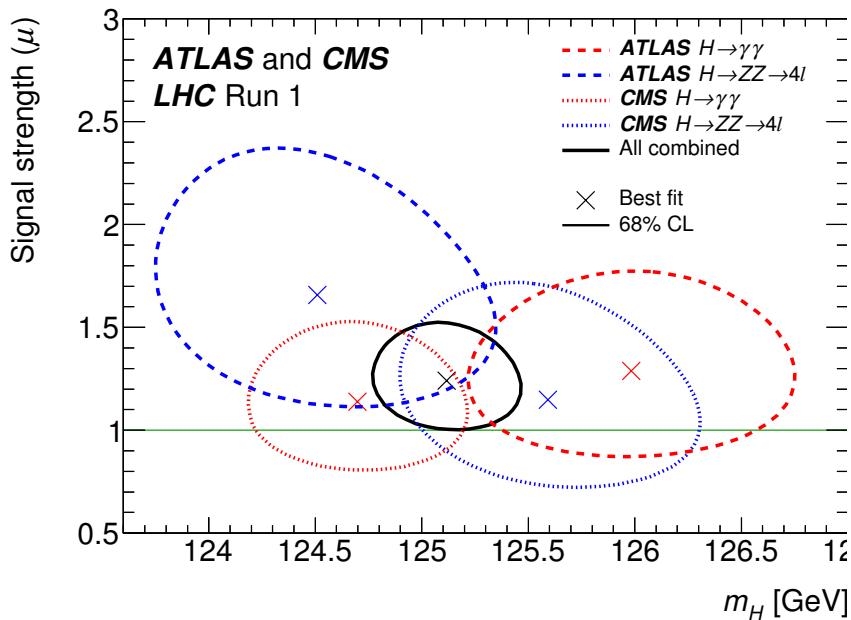
Roads to UV completions

Non-perturbative TeV-scale dynamics
(non-renormalizable models)

- Scalar LQ as PNG: Gripaios, '10; Gripaios, Nardecchia, Renner, '14
- Vector LQ (or W', Z') as technifermion resonances: Barbieri et al. '15; Buttazzo et al. '16; Barbieri et al. '17
- W', Z' as Kaluza-Klein excitations (e.g. from warped extra dim.): Megias, Quiros, Salas '17; Megias, Panico, Pujolas, Quiros '17

Perturbative TeV-scale dynamics
(renormalizable models)

- Renormalizable models with scalar mediators (LQ, but also RPV-SUSY): Hiller, Schmaltz, '14; Becirevic et al. '16; Fajfer et al. '15-'17; Dorsner et al. '17; Crivellin, Müller, Ota '17; Altmannshofer, Dev, Soni, '17
- Gauge models: Cline, Camalich '17; Calibbi, Crivellin, Li, '17; Assad, Fornal, Grinstein, '17; Di Luzio, Greljo, Nardecchia, '17

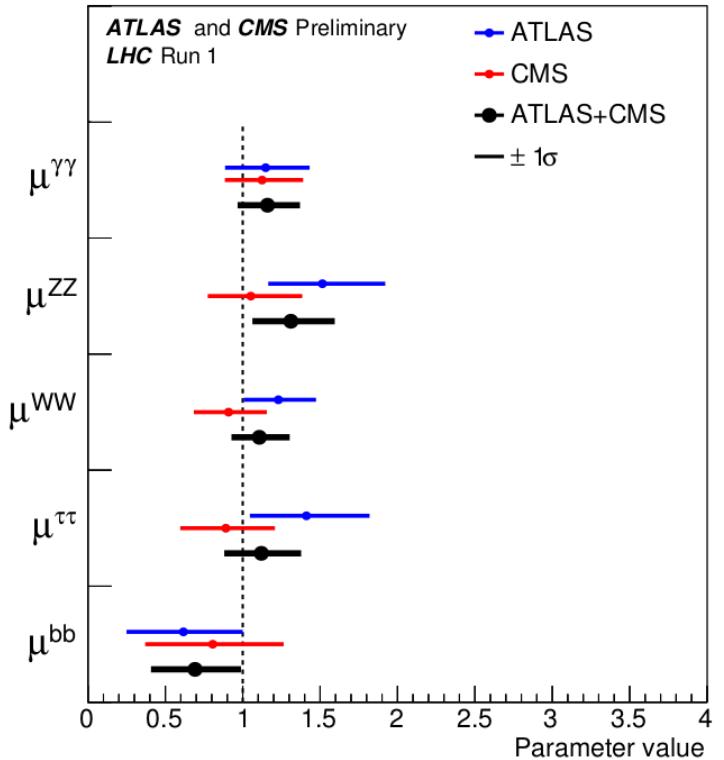


$m_H = 125.09 \pm 0.21 \text{ (stat)} \pm 0.11 \text{ (sys)} \text{ GeV}$

run 1, PRL **114** (2015) 191803

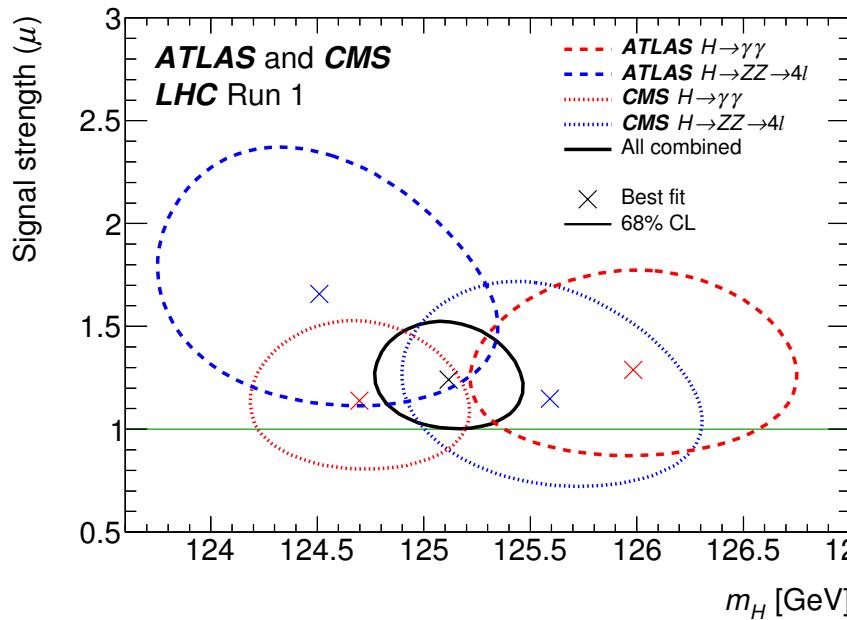
$m_H = 125.25 \pm 0.20 \text{ (stat)} \pm 0.08 \text{ (sys)} \text{ GeV}$

run 2, talk by R. Nicolaïdou @ Moriond'17, QCD



ATLAS-CONF-2015-044

CMS-PAS-HIG-15-002



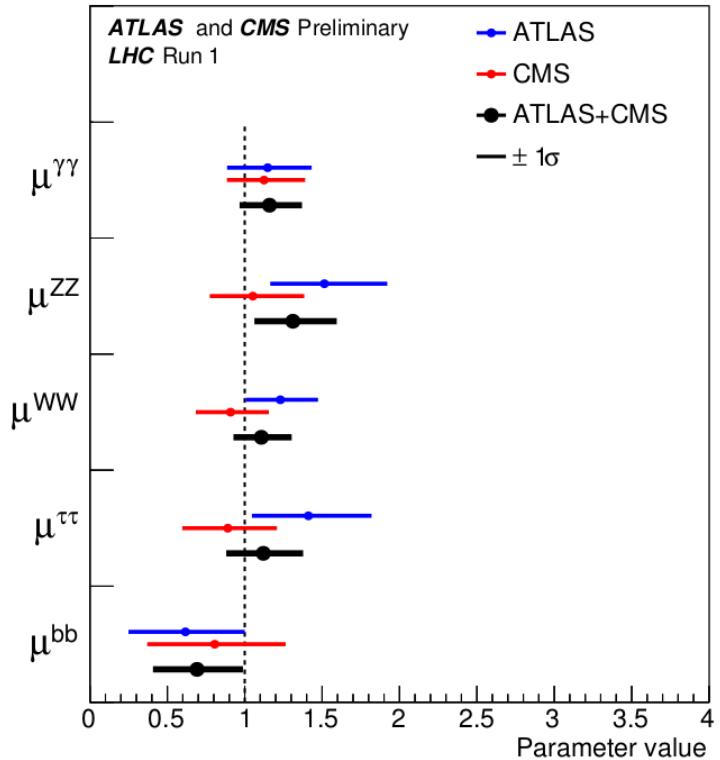
$m_H = 125.09 \pm 0.21 \text{ (stat)} \pm 0.11 \text{ (sys)} \text{ GeV}$

run 1, PRL 114 (2015) 191803

$m_H = 125.25 \pm 0.20 \text{ (stat)} \pm 0.08 \text{ (sys)} \text{ GeV}$

run 2, talk by R. Nicolaïdou @ Moriond'17, QCD

$(125 \text{ GeV})^2 \simeq m_Z^2 + (86 \text{ GeV})^2 \Rightarrow \text{large corrections within MSSM}$



ATLAS-CONF-2015-044

CMS-PAS-HIG-15-002

- after EWSB:

neutral CP-even: h, H

neutral CP-odd: A

charged: H^+, H^-

- Higgs masses:

at tree level

$$m_A, \tan \beta = v_u/v_d$$

$$m_h \leq m_Z$$

Ellis et al; Okada et al; Haber,Hempfling;
Hoang et al; Carena et al; Heinemeyer et al;
Zhang et al; Brignole et al; Harlander et al;
Kant,Harlander,Mihaila,Steinhauser;...

at higher order:

$$m_h^2 \simeq m_Z^2 \cos^2(2\beta) + \frac{3m_t^4}{4\pi^2 v^2} \left[\ln \left(\frac{M_S^2}{m_t^2} \right) + \frac{X_t^2}{M_S^2} \left(1 - \frac{X_t^2}{12M_S^2} \right) \right]$$

$$M_S^2 = m_{\tilde{t}_1} m_{\tilde{t}_2}, \quad X_t = A_t - \mu \cot \beta$$

$$m_H, m_A, m_{H^+} : O(v) \dots O(TeV)$$

$$m_{H^+}^2 = m_A^2 + m_W^2$$

$$v^2 = v_u^2 + v_d^2 = 4m_W^2/g^2$$

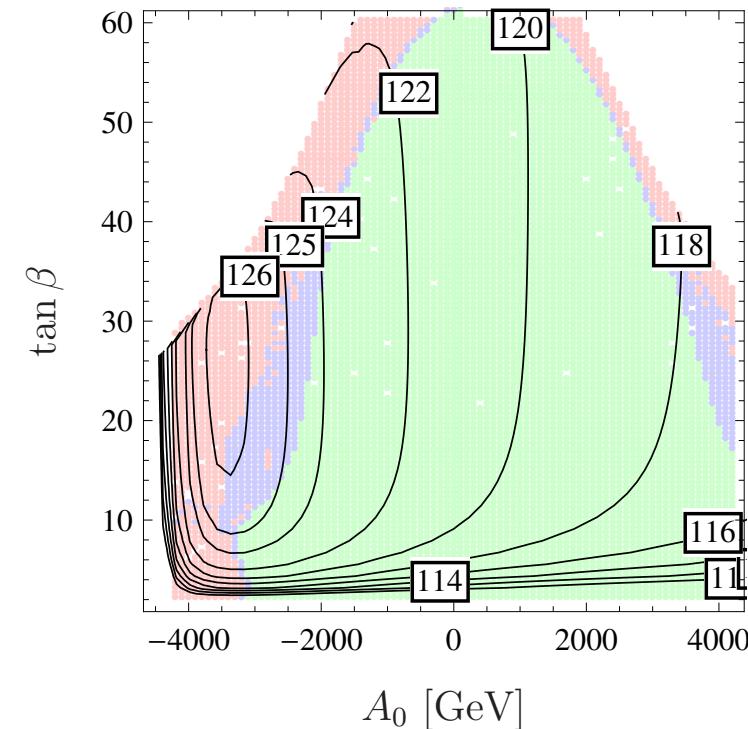
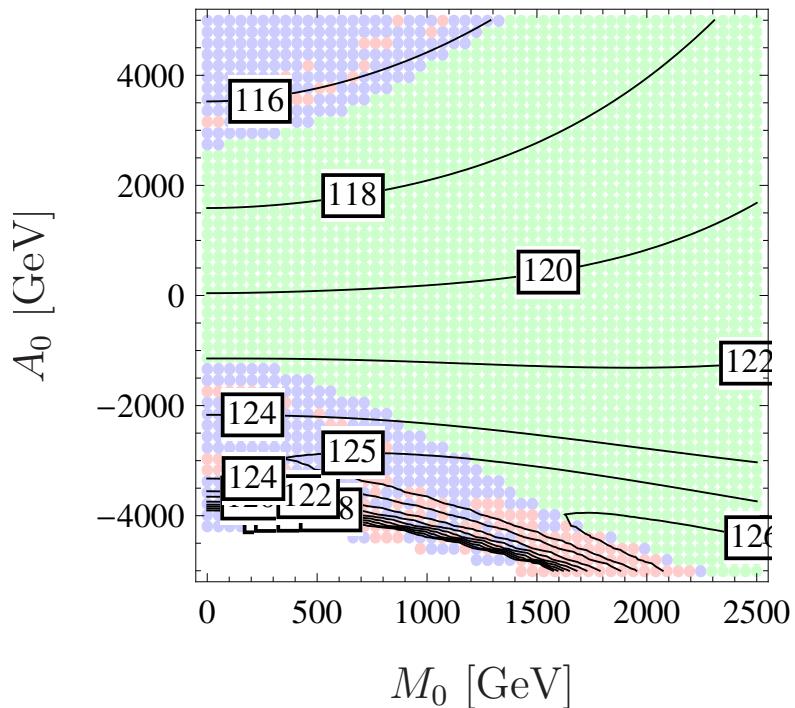
decoupling limit: $m_A \gg v, \tan \beta \gg 1$;

current precision including leading 3-loop effects: $O(\text{GeV})$

$m_h = 125.2 \text{ GeV}$ \Rightarrow large loop contributions
 \Rightarrow heavy stops and/or large left-right mixing for stops

- GMSB: $m_{\tilde{t}_1} \gtrsim 6 \text{ TeV}$,
 M. A. Ajaib, I. Gogoladze, F. Nasir, Q. Shafi, arXiv:1204.2856
 more complicated models based on P. Meade, N. Seiberg and D. Shih,
 arXiv:0801.3278 \Rightarrow allow additional terms
 e.g. S. Knappen, D. Redigolo, arXiv:1606.07501 $m_{\tilde{t}_1} \simeq m_{\tilde{b}_1} \gtrsim 1 \text{ TeV}$ if $M_{\text{mess}} \gtrsim 10^{15} \text{ GeV}$
- CMSSM, NUHM models: $|A_0| \simeq 2m_0$,
 H. Baer, V. Barger and A. Mustafayev, arXiv:1112.3017; M. Kadastik *et al.*,
 arXiv:1112.3647; O. Buchmueller *et al.*, arXiv:1112.3564; J. Cao, Z. Heng, D. Li,
 J. M. Yang, arXiv:1112.4391; L. Aparicio, D. G. Cerdeno, L. E. Ibanez,
 arXiv:1202.0822; J. Ellis, K. A. Olive, arXiv:1202.3262; ...
CMSSM fit to data P. Bechtle et al., arXiv:1508.05951: best fit point with
 $m_{\tilde{g}}, m_{\tilde{q}} \gtrsim 2 \text{ TeV}, m_{\tilde{l}_R} \simeq 600 \text{ GeV}, m_{\tilde{\chi}_1^0} \simeq 450 \text{ GeV}$
- general high scale models: $A_0 \simeq -(1 - 3) \max(M_{1/2}, m_{Q_3}, m_{U_3}) @ M_{\text{GUT}}$
 among other cases, details in F. Brümmer, S. Kraml and S. Kulkarni, arXiv:1204.5977

- SUSY models contain many scalars \Rightarrow complicated potential
- usually some parameters (μ, B) are chosen to obtain correct EWSB
- does not exclude the existence of other minima breaking charge and/or color!



$$M_{1/2} = 1 \text{ TeV}, \tan \beta = 10, \mu > 0$$

$$M_{1/2} = M_0 = 1 \text{ TeV}$$

J.E. Camargo-Molina, B. O'Leary, W.P., F. Staub, arXiv:1309.7212

$$\mathcal{L}_{MSSM} = \mu \tilde{H}_u \tilde{H}_d + \text{h.c.} + (m_{H_u}^2 + |\mu|^2) |H_u|^2 + (m_{H_d}^2 + |\mu|^2) |H_d|^2 + \dots$$

basic idea: contributions to m_Z should not contain large cancellations among each other

$$\frac{m_Z^2}{2} = \frac{m_{H_d}^2 + \Sigma_d - (m_{H_u}^2 + \Sigma_u) \tan^2 \beta}{\tan^2 \beta - 1} - |\mu|^2 \simeq -(m_{H_u}^2 + \Sigma_u) - |\mu|^2$$

'standard fine-tuning measure'[†]

$$\Delta_{FT} = \max[c_i] , \quad c_i = \left| \frac{\partial \ln m_Z^2}{\partial \ln p_i} \right| = \left| \frac{p_i}{m_Z^2} \frac{\partial m_Z^2}{\partial p_i} \right|$$

requiring at most a tuning at the per-cent level one finds

$$|\mu|^2 \simeq m_Z^2 \simeq |m_{H_u}^2| , \quad m_{\tilde{t}_1} \lesssim 1 \text{ TeV} , \quad m_{\tilde{g}} \lesssim 1\text{-}2 \text{ TeV}$$

[†] Ellis, Enqvist, Nanopoulos, Zwirner 1986; Barbieri, Guidice 1988

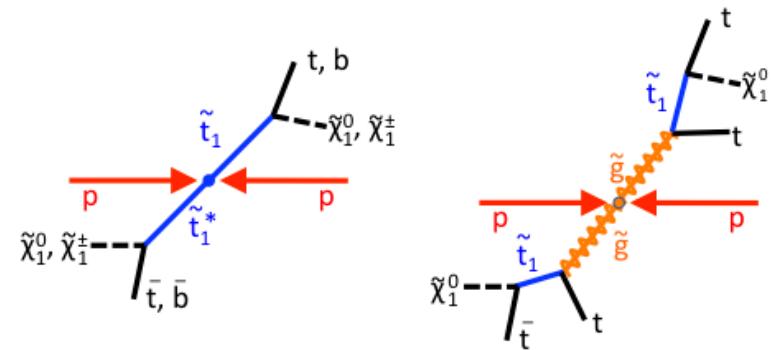
several studies, see e.g. S. Sekmen et al., arXiv:1109.5119; A. Arbey, M. Battaglia, A. Djouadi and F. Mahmoudi, arXiv:1211.4004; M. Cahill-Rowley, J. Hewett, A. Ismail and T. Rizzo, arXiv:1308.0297

- generic signatures are well known: multi-lepton, multi-jets + missing E_T
- sub-class of general MSSM: ‘natural SUSY’
see e.g. M. Papucci, J. T. Ruderman and A. Weiler, arXiv:1110.6926;
H. Baer, V. Barger, P. Huang, A. Mustafayev, X. Tata, arXiv:1207.3343
keep only SUSY particles light needed for ‘natural Higgs’:

$$\tilde{t}_1, \tilde{b}_1, \tilde{g}, \tilde{\chi}_1^0, \tilde{\chi}_1^+ \simeq \tilde{h}_{1,2}^0, \tilde{h}^+ \simeq \tilde{h}^+$$

$$\Rightarrow 100 \text{ MeV} \lesssim m_{\tilde{\chi}_1^+} - m_{\tilde{\chi}_1^0} \simeq m_{\tilde{\chi}_2^0} - m_{\tilde{\chi}_1^0} \lesssim 5 - 10 \text{ GeV}$$

$$\begin{aligned}\tilde{g} &\rightarrow \tilde{t}_1 t, \tilde{b}_1 b \\ \tilde{t}_1 &\rightarrow t \tilde{\chi}_{1,2}^0, b \tilde{\chi}_1^+, W^+ \tilde{b}_1 \\ \tilde{b}_1 &\rightarrow b \tilde{\chi}_{1,2}^0, t \tilde{\chi}_1^-, W^- \tilde{t}_1\end{aligned}$$

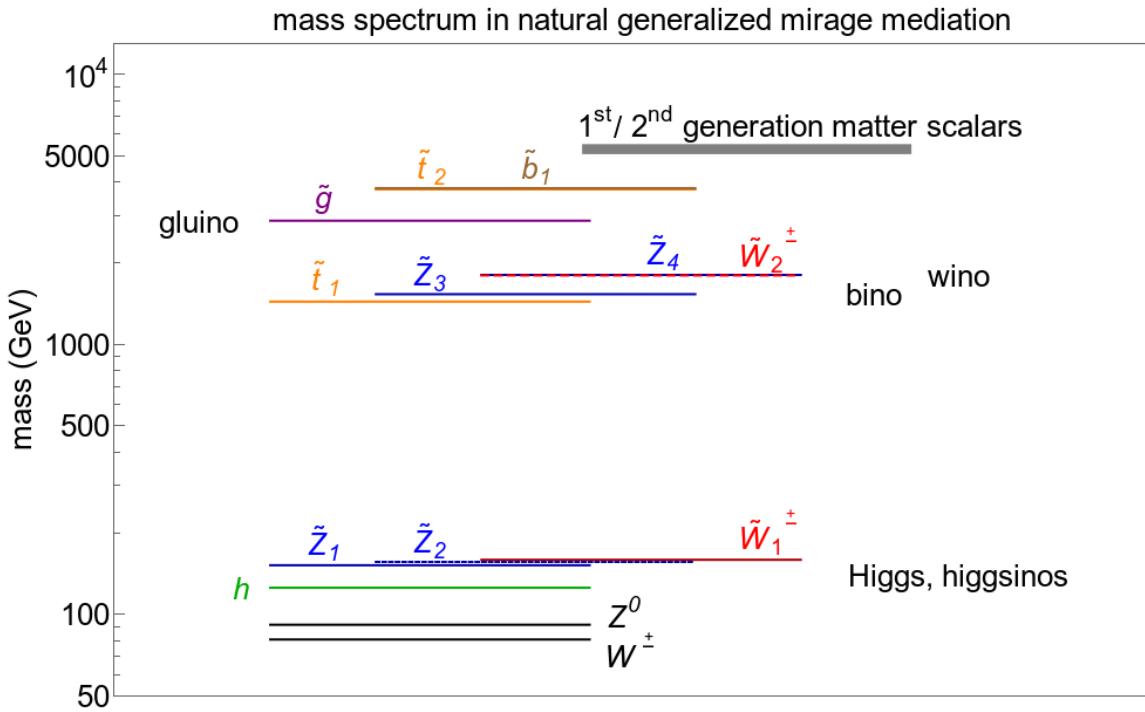


BRs depend on the nature of \tilde{t}_1 and \tilde{b}_1

Higgsino mass: $\mu + \mu'$ with soft SUSY breaking parameter: $\mathcal{L} = \mu' \tilde{H}_u \tilde{H}_d$

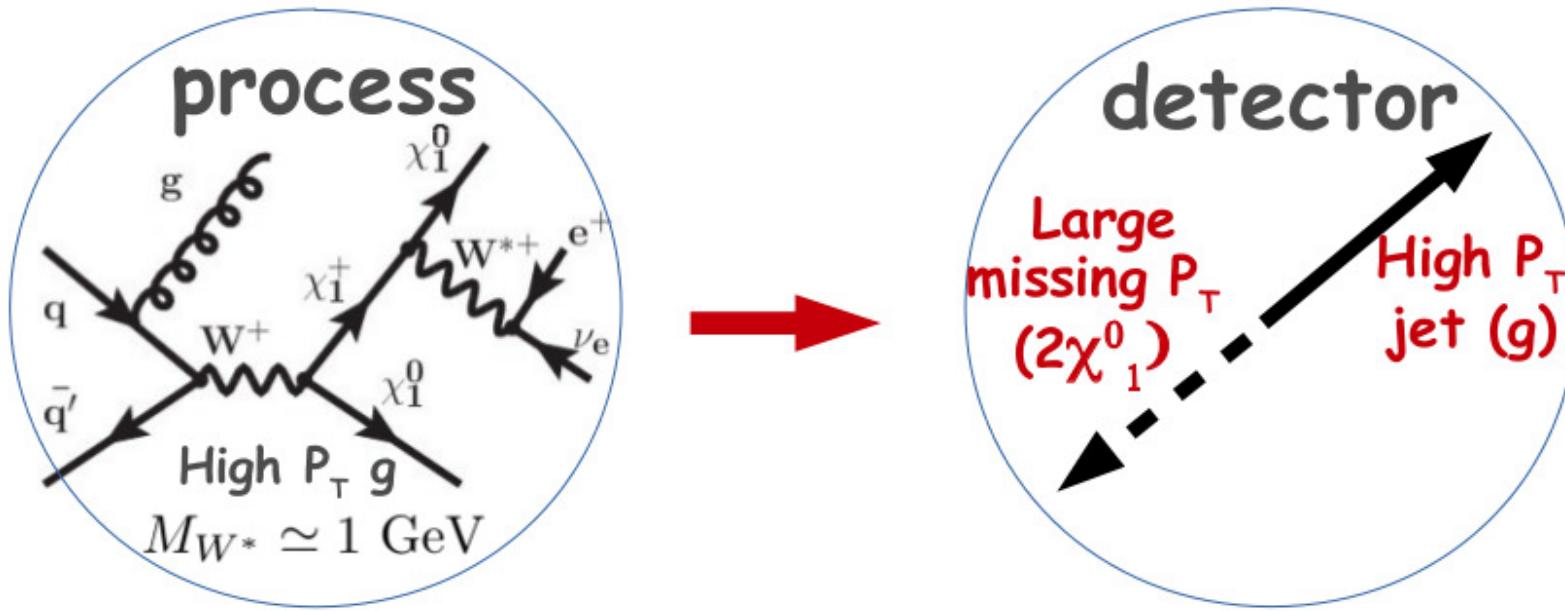
Different sources for soft SUSY breaking: moduli & AMSB

main consequence: gaugino masses unify at a (vastly) different scale than gauge couplings



H. Baer, V. Barger, H. Serce and X. Tata, arXiv:1610.06205

Most challenging case: only higgsinos accessible but nothing else
and ΔM too small for any leptonic signature

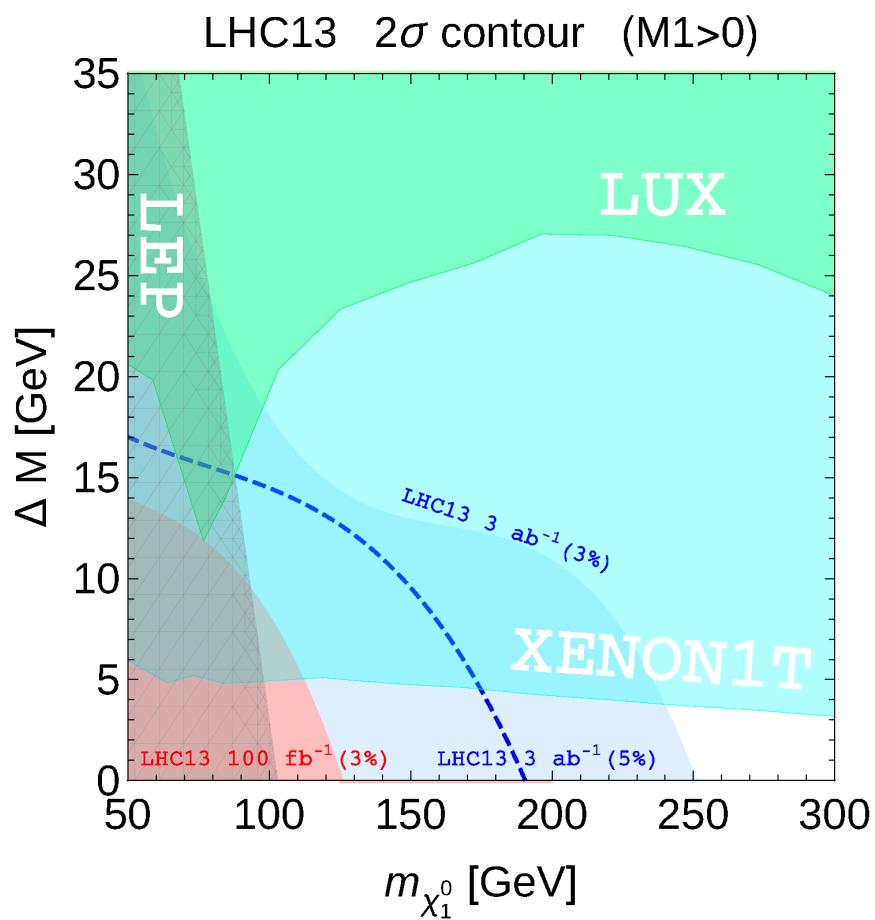


The only way to probe compressed higgsinos is a mono-jet signature:
'Where the Sidewalk Ends? ...' Alves, Izaguirre, Wacker 2011

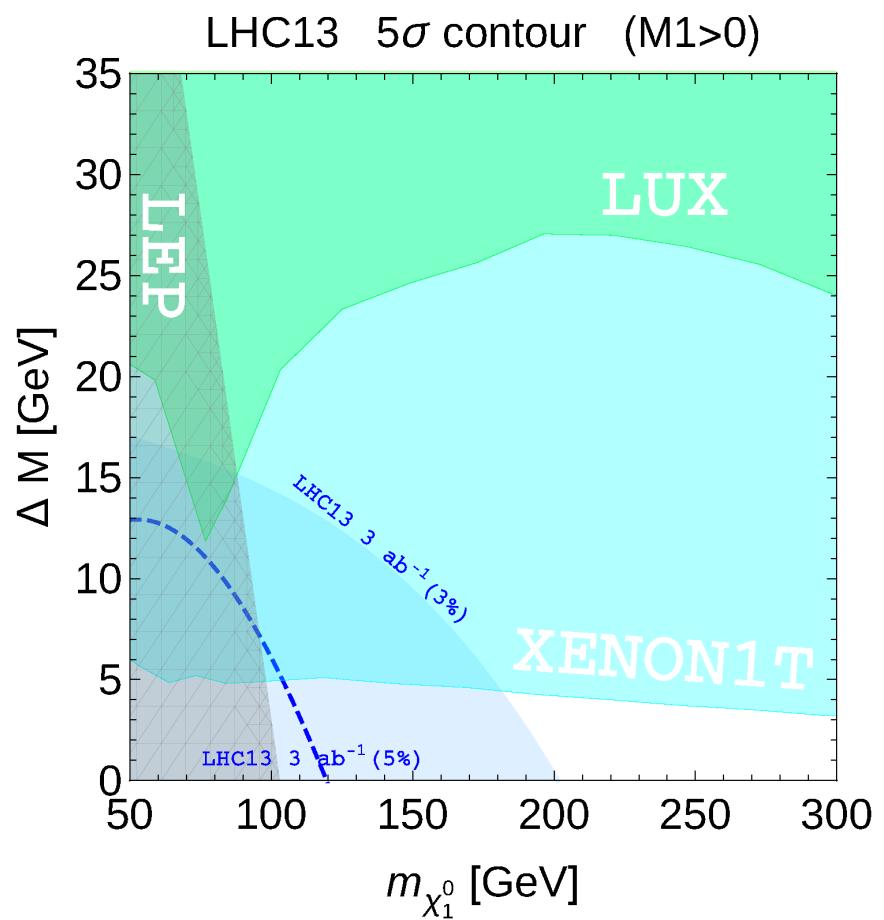
related work C. Han et al., arXiv:1310.4274; P. Schwaller, J. Zurita, arXiv:1312.7350;
Z. Han et al, arXiv:1401.1235; H. Baer et al., arXiv:1401.1162, ...

	$Z(\nu\bar{\nu})j$	$W(l\nu)j$	$\mu = 100 \text{ GeV}$	$\mu = 200 \text{ GeV}$
Initial # of events	$3.15 \cdot 10^6$	$1.25 \cdot 10^7$	$3.63 \cdot 10^5$	$6.45 \cdot 10^3$
$p_T^j > 200 \text{ GeV} \mid \eta^j \mid < 2.4$	$1.05 \cdot 10^6$	$4.11 \cdot 10^6$	$1.73 \cdot 10^5$	3528
Jet veto ($n \geq 3$)	$8.7 \cdot 10^5$	$3.13 \cdot 10^6$	$1.33 \cdot 10^5$	2691
$\Delta\phi(j_1, j_2) < 2.5$	$7.2 \cdot 10^5$	$2.3 \cdot 10^6$	$1.10 \cdot 10^5$	2320
Veto e^\pm, μ^\pm, τ^\pm	$7.2 \cdot 10^5$	$6.8 \cdot 10^5$	$1.08 \cdot 10^5$	2301
$E_T^{\text{miss}} > 200 \text{ GeV}$	$6.4 \cdot 10^5$	$4.3 \cdot 10^5$	9846	2188
$E_T^{\text{miss}} > 600 \text{ GeV}$	4353	1002	171	93
$E_T^{\text{miss}} > 800 \text{ GeV}$	694	0	37	22

exclusion reach



discovery reach



D. Barducci, A. Belyaev, A. Bharucha, WP, V. Sanz, arXiv:1504.02472

- additional D-term contributions to m_h at tree-level

extra $U(1)_\chi$: $m_{h,tree}^2 \leq m_Z^2 + \frac{1}{4}g_\chi^2 v^2$

- Origin of R -parity $R_P = (-1)^{2s+3(B-L)}$

$$\Rightarrow SO(10) \rightarrow SU(3)_C \times SU(2)_L \times SU(2)_R \times U(1)_{B-L}$$

$$\rightarrow SU(3)_C \times SU(2)_L \times U(1)_R \times U(1)_{B-L}$$

$$\cong SU(3)_C \times SU(2)_L \times U(1)_Y \times U(1)_\chi$$

or $E(8) \times E(8) \rightarrow SU(3)_C \times SU(2)_L \times U(1)_Y \times U(1)_{B-L}$

- Neutrino masses

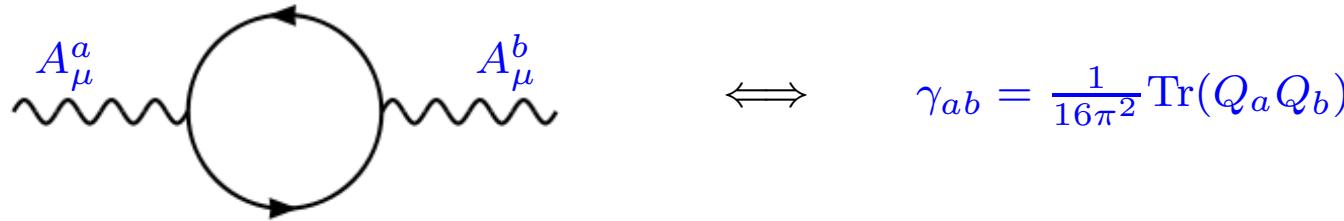
$$B - L \text{ anomaly free} \Rightarrow \nu_R$$

usual seesaw, inverse seesaw

$U(1)_a \times U(1)_b$ models allow for

(B. Holdom, PLB 166 (1986) 196)

$$\mathcal{L} \supset -\chi_{ab} \hat{F}^{a,\mu\nu} \hat{F}_{\mu\nu}^b$$



equivalent

$$D_\mu = \partial_\mu - i \begin{pmatrix} Q_a \\ Q_b \end{pmatrix} \begin{pmatrix} g_{aa} & g_{ab} \\ g_{ba} & g_{bb} \end{pmatrix} \begin{pmatrix} A_\mu^a \\ A_\mu^b \end{pmatrix}$$

both $U(1)$ unbroken \Rightarrow chose basis with e.g. $g_{ba} = 0$

affects also RGE running of soft SUSY parameters:

R. Fonseca, M. Malinsky, W.P., F. Staub, arXiv:1107.2670

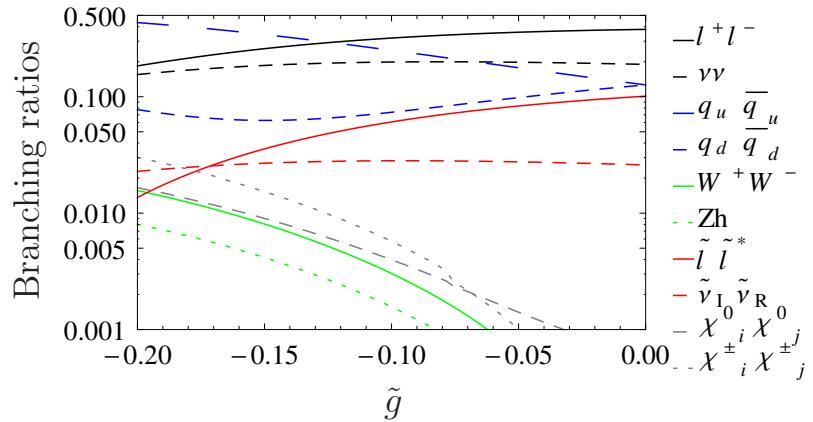
Superfield	Generations	$U(1)_Y \times SU(2)_L \times SU(3)_C \times U(1)_{B-L}$
\hat{Q}	3	$(\frac{1}{6}, \mathbf{2}, \mathbf{3}, \frac{1}{6})$
\hat{D}	3	$(\frac{1}{3}, \mathbf{1}, \overline{\mathbf{3}}, -\frac{1}{6})$
\hat{U}	3	$(-\frac{2}{3}, \mathbf{1}, \overline{\mathbf{3}}, -\frac{1}{6})$
\hat{L}	3	$(-\frac{1}{2}, \mathbf{2}, \mathbf{1}, -\frac{1}{2})$
\hat{E}	3	$(1, \mathbf{1}, \mathbf{1}, \frac{1}{2})$
$\hat{\nu}$	3	$(0, \mathbf{1}, \mathbf{1}, \frac{1}{2})$
\hat{H}_d	1	$(-\frac{1}{2}, \mathbf{2}, \mathbf{1}, 0)$
\hat{H}_u	1	$(\frac{1}{2}, \mathbf{2}, \mathbf{1}, 0)$
$\hat{\eta}$	1	$(0, \mathbf{1}, \mathbf{1}, -1)$
$\hat{\bar{\eta}}$	1	$(0, \mathbf{1}, \mathbf{1}, 1)$

$$\begin{aligned}
 W = & Y_u^{ij} \hat{U}_i \hat{Q}_j \hat{H}_u - Y_d^{ij} \hat{D}_i \hat{Q}_j \hat{H}_d - Y_e^{ij} \hat{E}_i \hat{L}_j \hat{H}_d + \mu \hat{H}_u \hat{H}_d + Y_\nu^{ij} \hat{L}_i \hat{H}_u \hat{\nu}_j \\
 & - \mu' \hat{\eta} \hat{\bar{\eta}} + Y_x^{ij} \hat{\nu}_i \hat{\eta} \hat{\nu}_j
 \end{aligned}$$

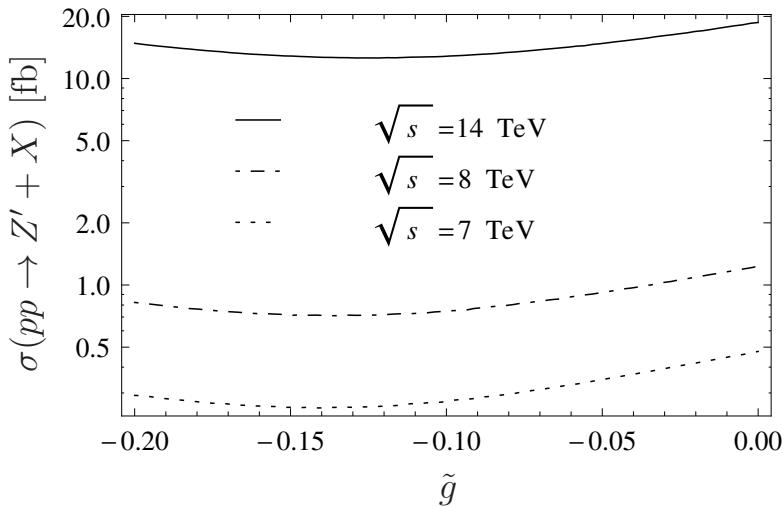
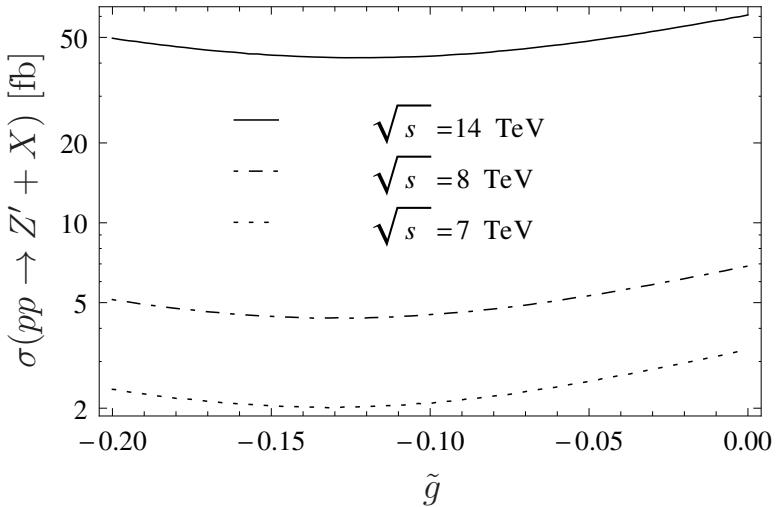
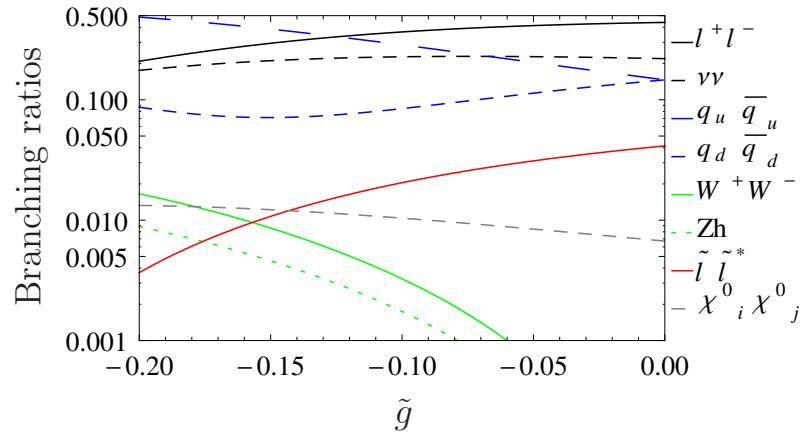
based on B. O'Leary, W.P., F. Staub, arXiv:1112.4600

K -factor for leptonic channels: ~ 1.2

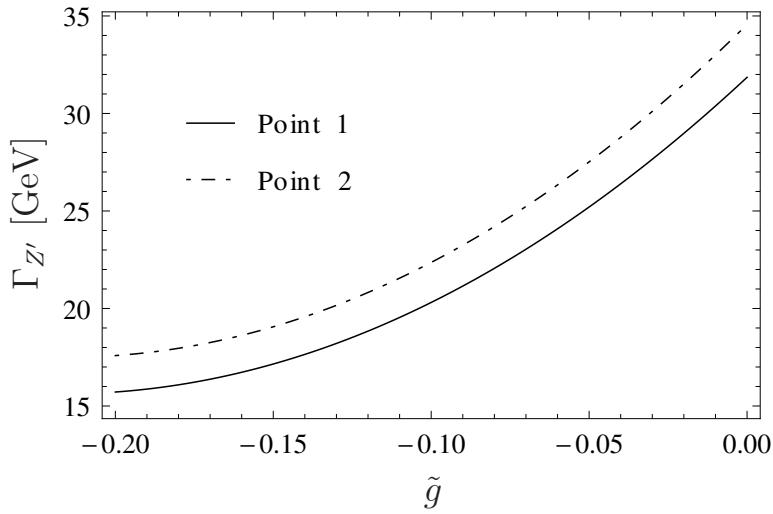
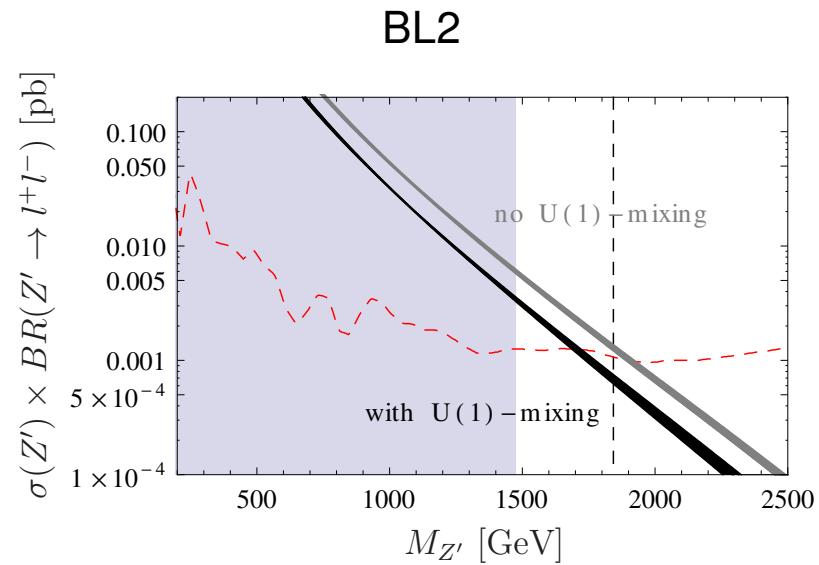
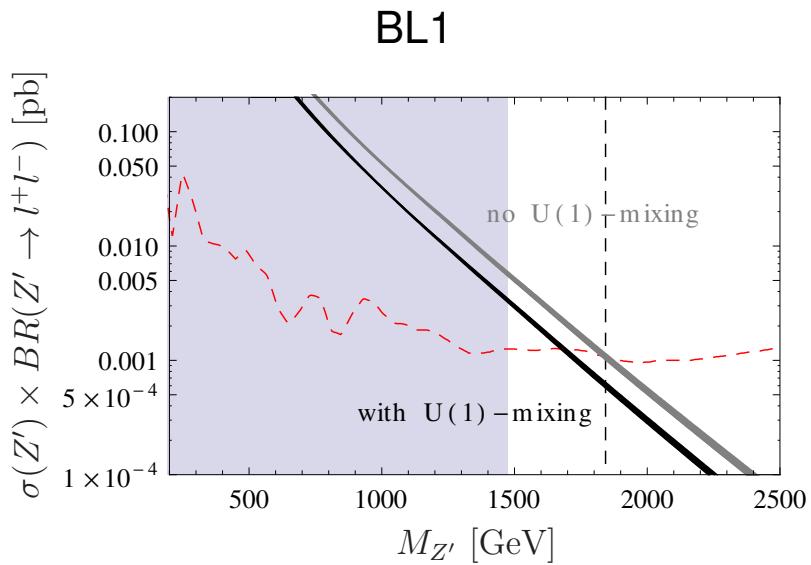
BL1



BL2



M. Krauss, B. O'Leary, W.P., F. Staub, arXiv:1206.3513



No.	$\tilde{g} = -0.11$	$\tilde{g} = 0$
BL1	1680 GeV	1840 GeV
BL2	1700 GeV	1910 GeV

Several programs with partially different philosophies

- CheckMATE, M. Drees et al., arXiv:1312.2591
- Fastlim, M. Papucci et al., arXiv:1402.0492
- MadAnalysis 5, E. Conte et al., arXiv:1206.1599
- Rivet, A. Buckley et al., arXiv:1003.0694
- SModelS, S. Kraml et al., arXiv:1312.4175

$$\mathcal{W}_{eff} = \mathcal{W}_{MSSM} + \frac{1}{2}(M_R)_{ij} \hat{\nu}_{R,i} \hat{\nu}_{R,j}$$

$$+ (Y_\nu)_{ij} \hat{L}_i \cdot \hat{H}_u \hat{\nu}_{R,j}$$

$$(Y_\nu)_{\ell 5} = \pm (Z_\ell^{NH})^* \sqrt{\frac{2m_3 M_5}{v_u}} \cosh \gamma_{56} e^{\mp i\theta_{56}}$$

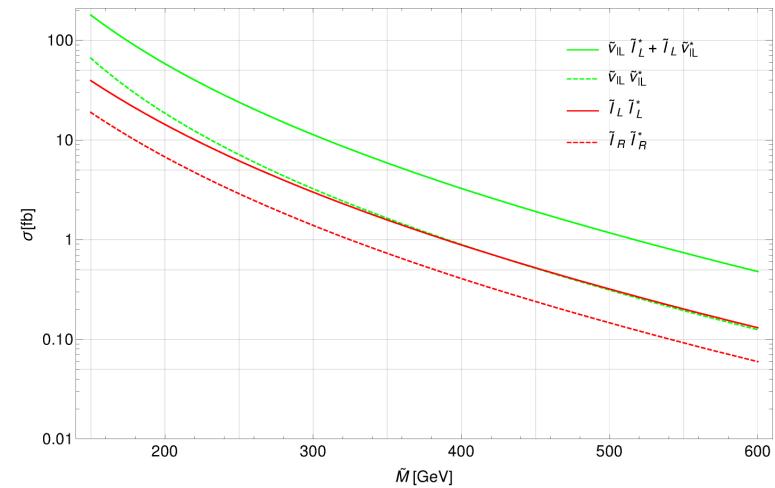
$$(Y_\nu)_{\ell 6} = -i(Z_\ell^{NH})^* \sqrt{\frac{2m_3 M_6}{v_u}} \cosh \gamma_{56} e^{\mp i\theta_{56}}$$

$$R = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \phi_{56} & \sin \phi_{56} \\ 0 & -\sin \phi_{56} & \cos \phi_{56} \end{pmatrix}$$

$$\phi_{56} \in \mathbb{C}$$

$$m_{\nu_h, i} \simeq M_{i-3}, M_4 = O(\text{keV}), \\ M_5 \simeq M_6 = O(\text{few - 100 GeV})$$

search for sleptons

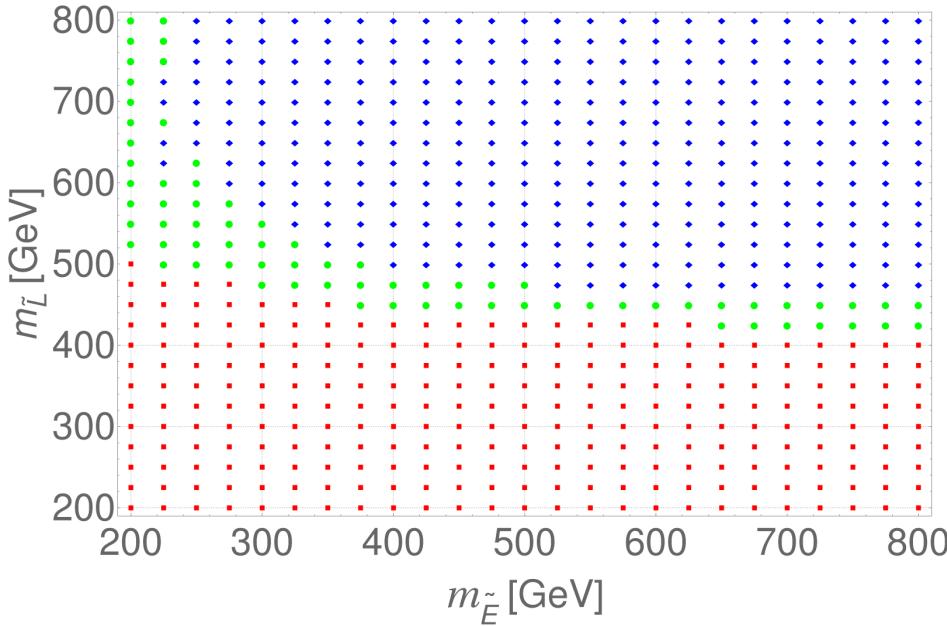


LHC, 13 TeV, tree-level
for searches: \times K-factor 1.17
(B. Fuks et al., arXiv:1304.0790)

dominant decays:

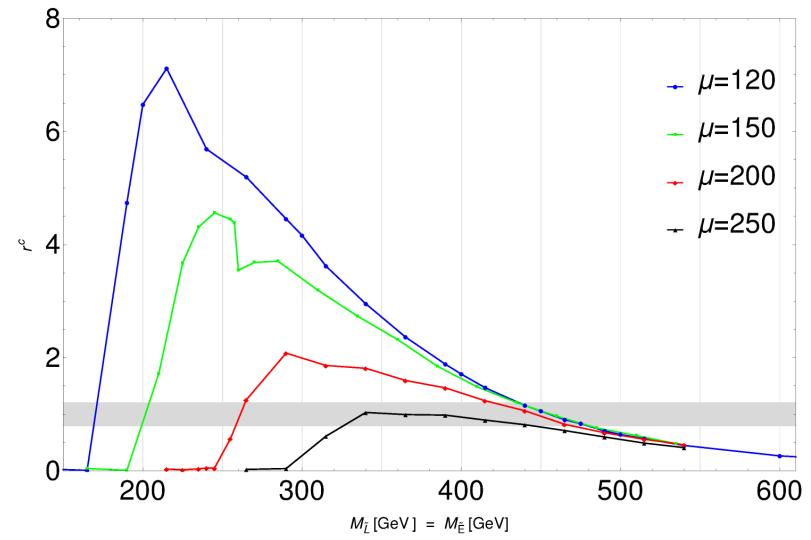
$$\tilde{l}_L \rightarrow l \tilde{\chi}_1^0, \nu \tilde{\chi}_1^-$$

$$\tilde{\nu}_L \rightarrow l^- \tilde{\chi}_1^+, \nu \tilde{\chi}_1^0$$



$\mu = 120 \text{ GeV}$, $\tan \beta = 10$

■ excluded, ● ambiguous, ◆ allowed , via



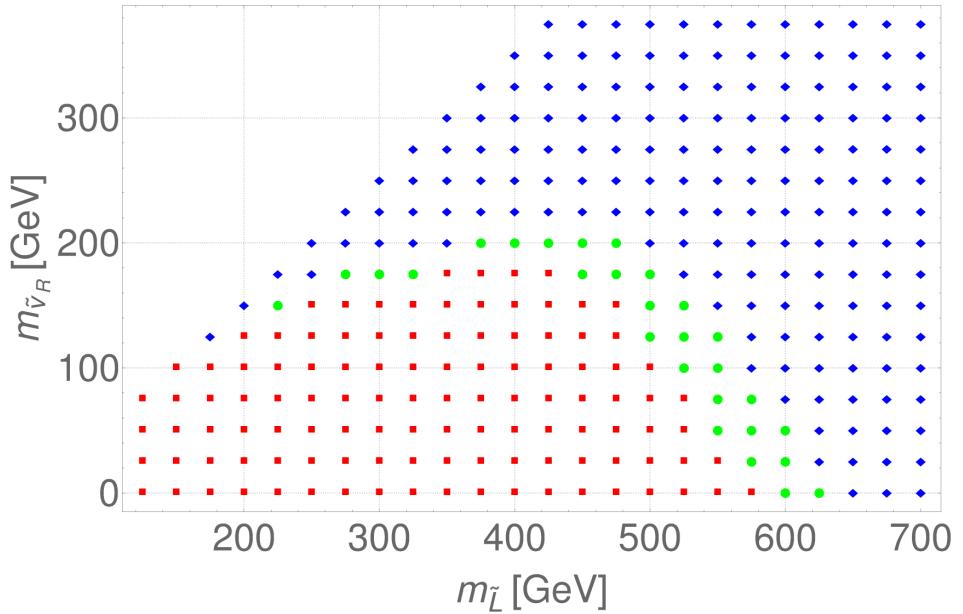
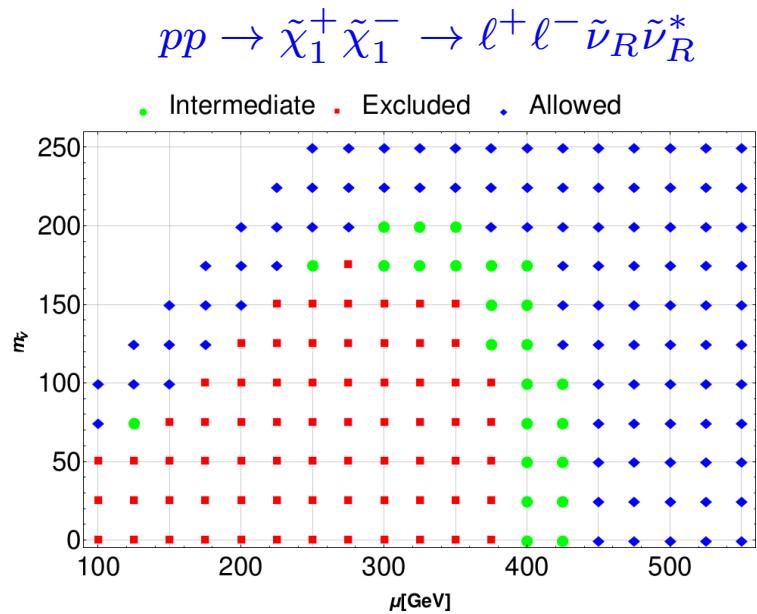
$m_{\tilde{L}} = m_{\tilde{E}}$, $\tan \beta = 10$

8+13 TeV data (13.9 fb^{-1})

using CheckMATE 2.0

Nh. Cerna-Velazco, Th. Faber, J. Jones, WP arXiv:1705.06583

additional constraint

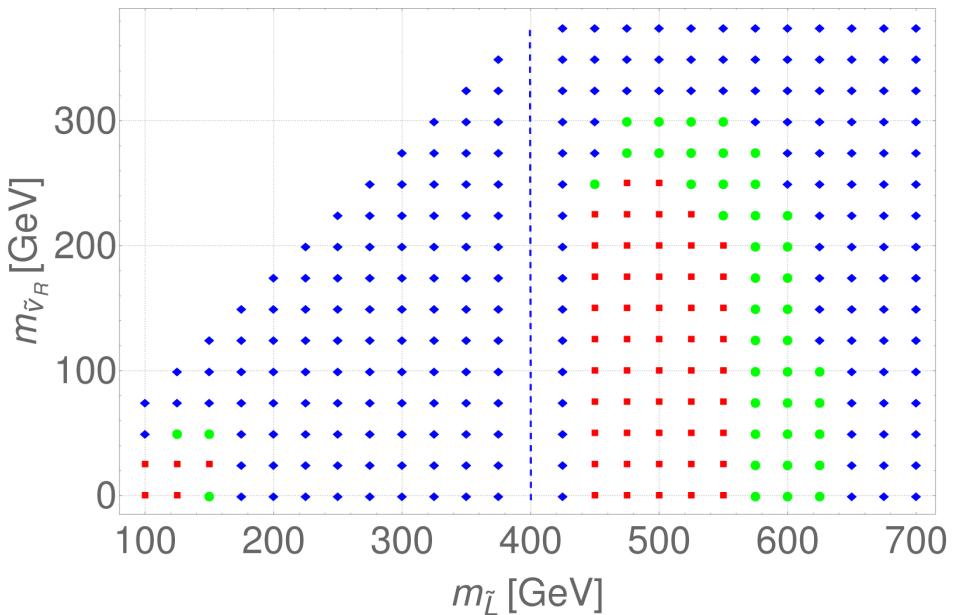
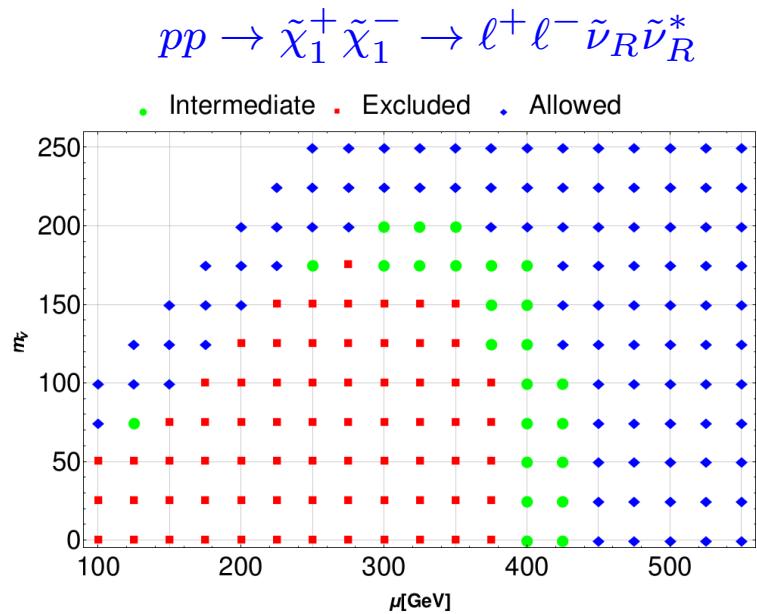


8+13 TeV data (13.9 fb^{-1})
using CheckMATE 2.0
 $m_{\tilde{\nu}_R} = 20 \text{ GeV}$

$\mu = 25 + m_{\tilde{\nu}} < m_{\tilde{l}} \simeq m_{\tilde{L}} = m_{\tilde{E}}$
 $M_1 = M_2 = 2 \text{ TeV}, \tan \beta = 6$

Nh. Cerna-Velazco, Th. Faber, J. Jones, WP arXiv:1705.06583

additional constraint



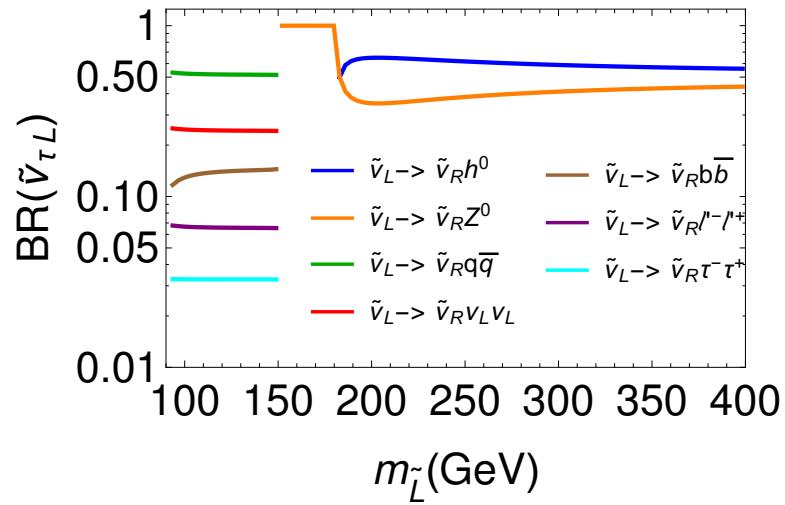
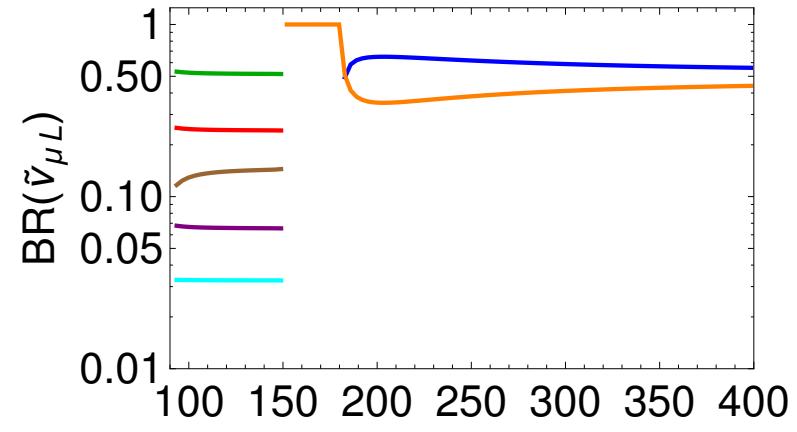
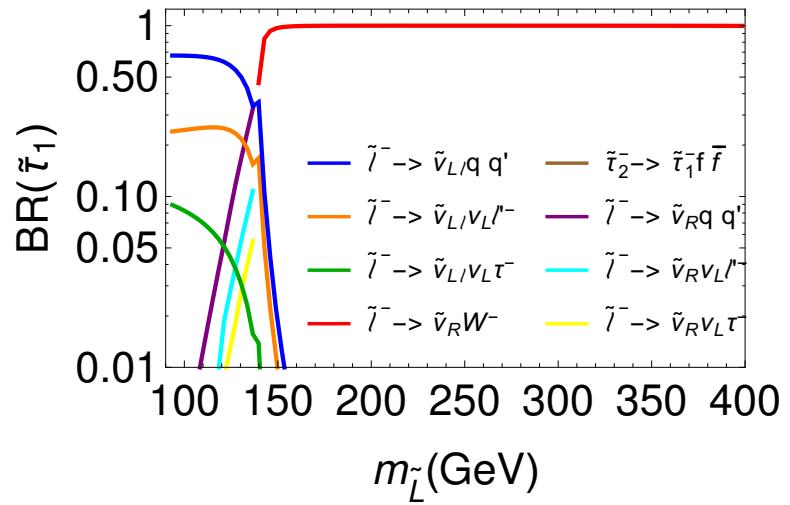
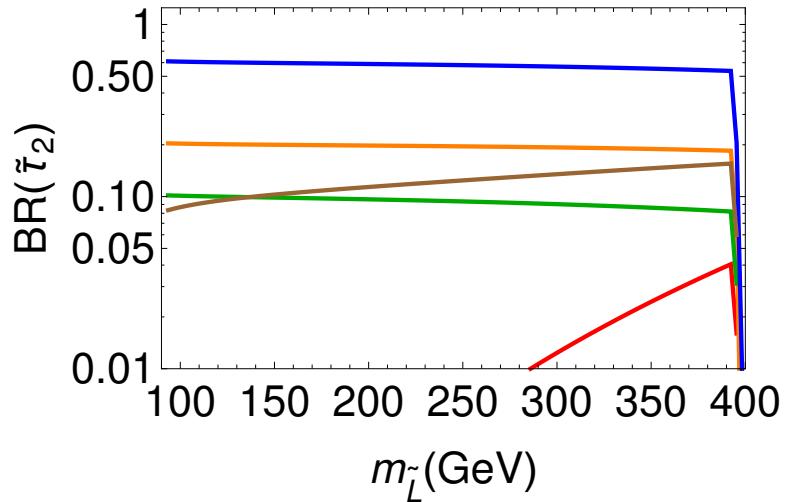
8+13 TeV data (13.9 fb^{-1})
using CheckMATE 2.0
 $m_{\tilde{\nu}_R} = 20 \text{ GeV}$

$\mu = 400 \text{ GeV}$, $m_{\tilde{l}} \simeq m_{\tilde{L}} = m_{\tilde{E}}$
 $M_1 = M_2 = 2 \text{ TeV}$, $\tan \beta = 6$

Nh. Cerna-Velazco, Th. Faber, J. Jones, WP arXiv:1705.06583

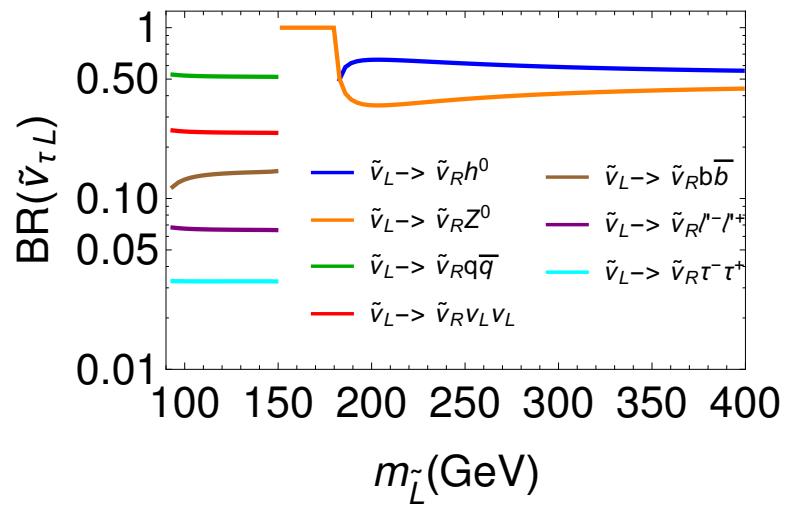
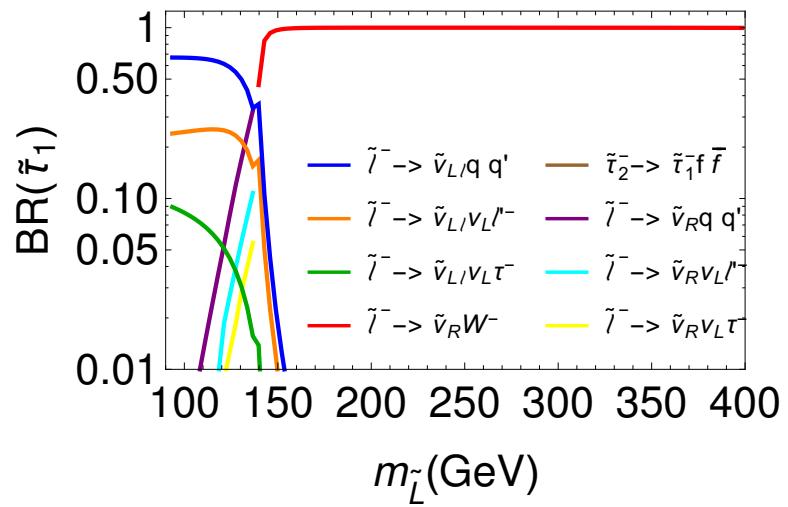
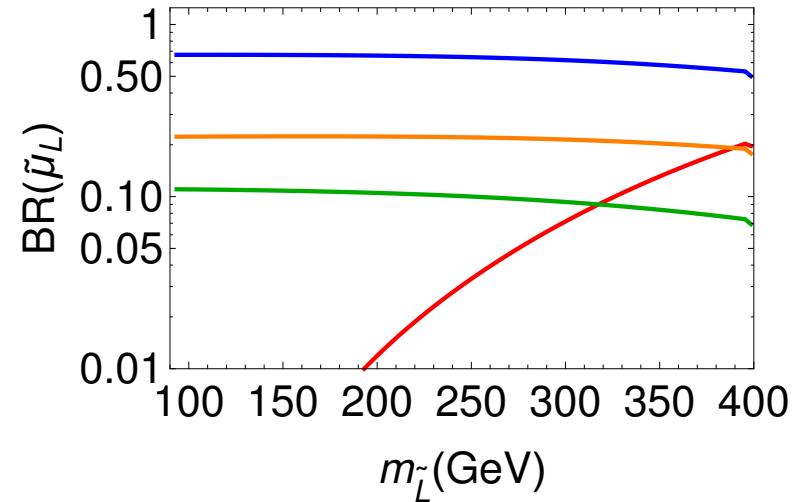
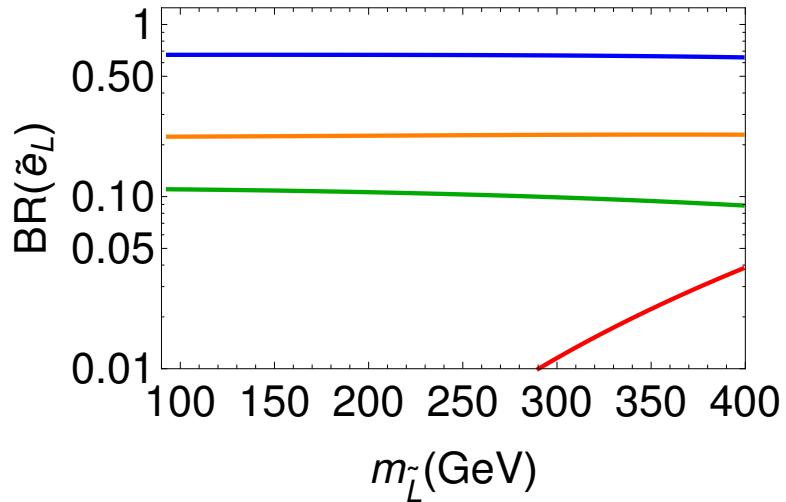
- no conclusive BSM signal so far: need good knowledge of tails of distributions
- QCD corrections (and partially EW corrections) are known for production and decays of several BSM particles
 - strongly interacting SUSY particles
 - electroweak SUSY particles
 - leptoquarks
 - $pp \rightarrow Z' \rightarrow l^+l^-$, $pp \rightarrow W' \rightarrow l\nu$
 - extended Higgs sectors
 - ...

for $\mu = 400 \text{ GeV} > m_{\tilde{L}} = m_{\tilde{E}}$, $\tan \beta = 6$, $M_1, M_2 \geq 500 \text{ GeV}$



Nh. Cerna-Velazco, Th. Faber, J. Jones, WP arXiv:1705.06583

for $\mu = 400 \text{ GeV} > m_{\tilde{L}} = m_{\tilde{E}}$, $\tan \beta = 6$, $M_1, M_2 \geq 500 \text{ GeV}$



Nh. Cerna-Velazco, Th. Faber, J. Jones, WP arXiv:1705.06583